

Ultracold dipolar bosonic molecules

Workshop on Quantum Simulations with Ultracold Atoms

ICTP Trieste

Tetsu Takekoshi, 20.07.12







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Nägerl (PI)

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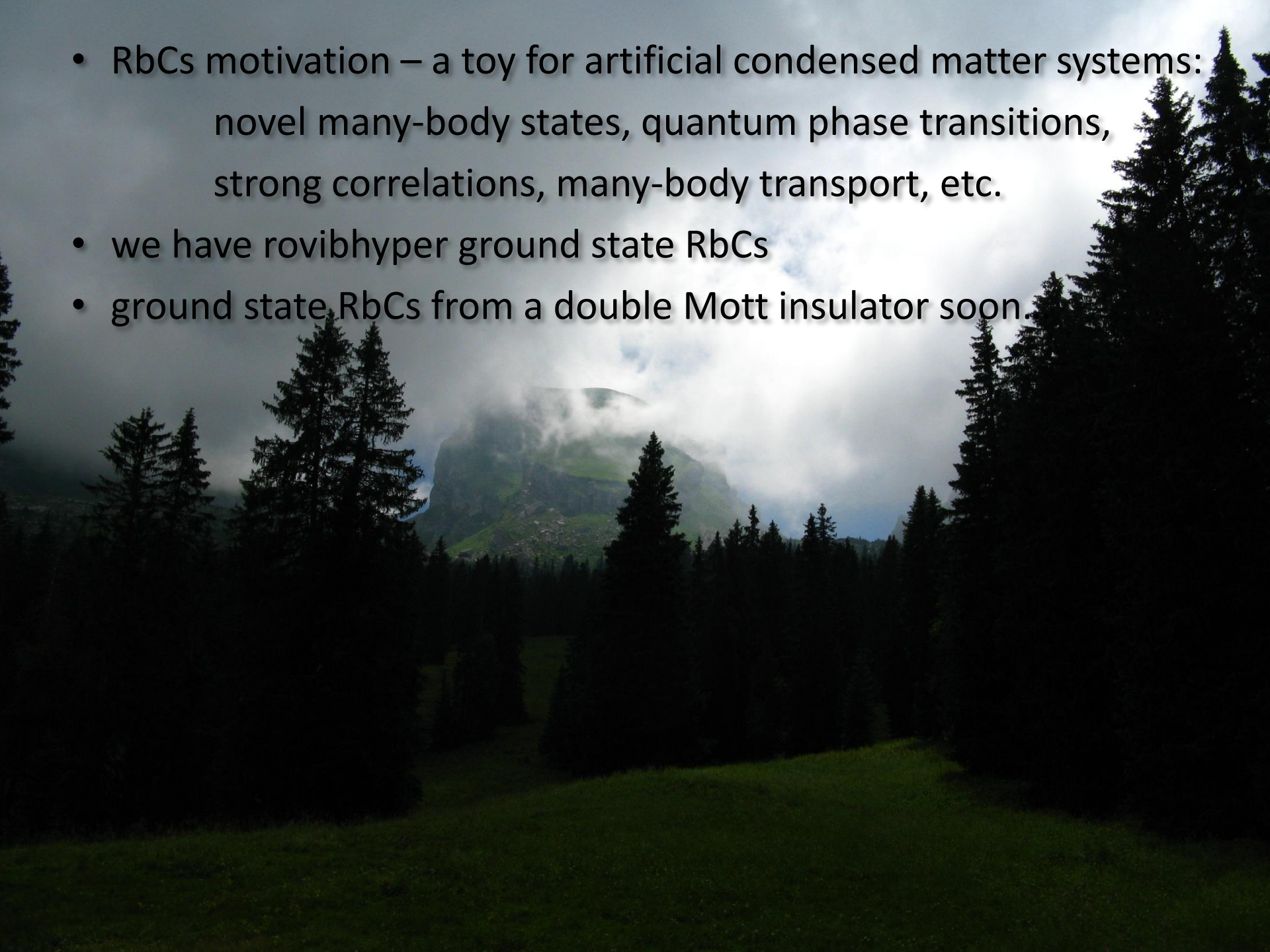
Michael
Kugler
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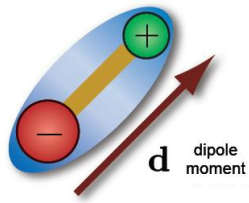
Verena
Pramhaas
(Masters)

Carl Hippler
(Masters)

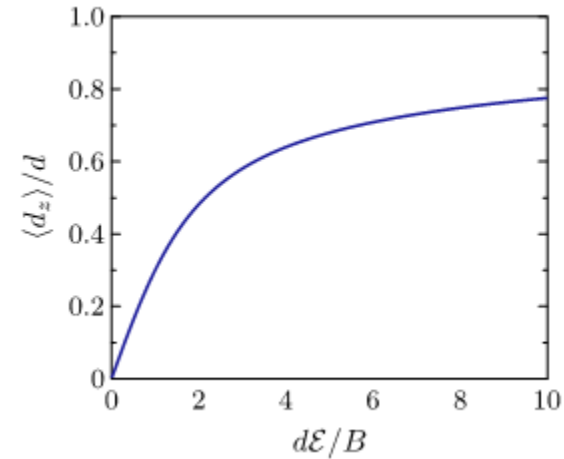
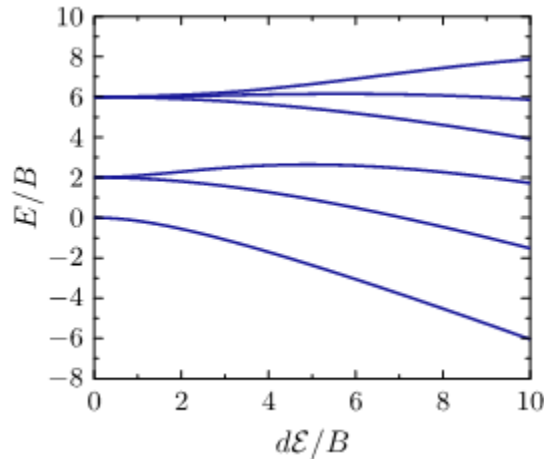
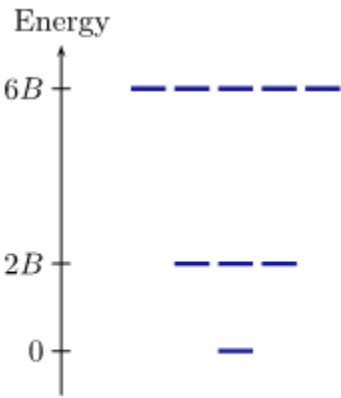
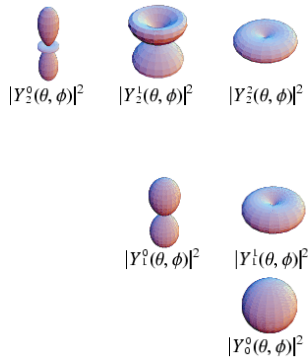
- RbCs motivation – a toy for artificial condensed matter systems:
novel many-body states, quantum phase transitions,
strong correlations, many-body transport, etc.
- we have rovibhyper ground state RbCs
- ground state RbCs from a double Mott insulator soon.



Why degenerate dipolar gases?



1Σ rigid rotor



$$\hat{H}_{\text{rot}} = B\hat{J}^2$$

$$\hat{H} = \hat{H}_{\text{rot}} - \hat{d} \cdot \mathbf{E} = \hat{H}_{\text{rot}} - d\mathcal{E} \cos\theta$$

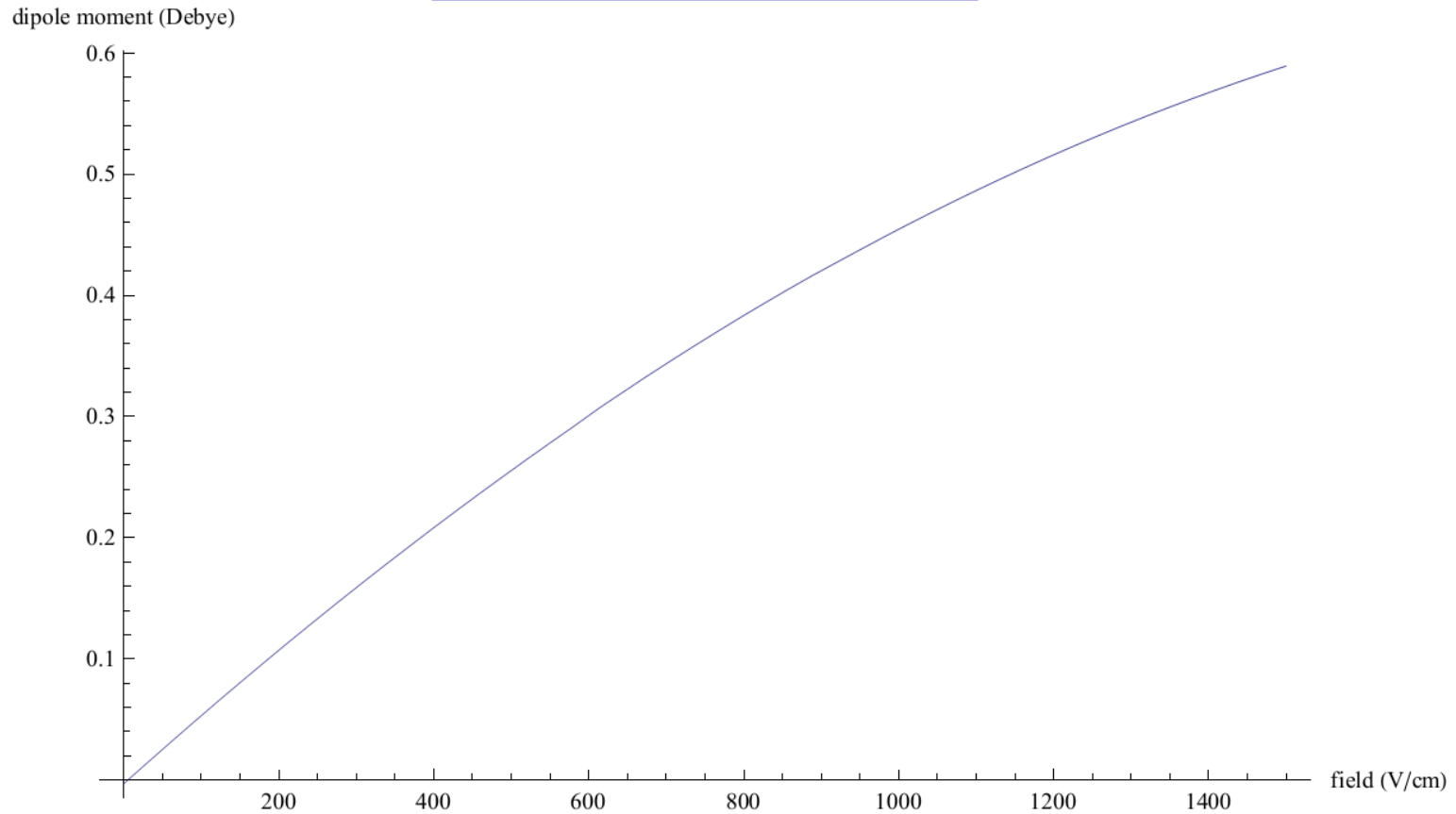
$$\mathbf{E} = \mathcal{E} \mathbf{e}_z$$

Review articles:

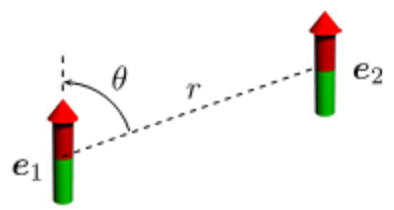
T. Lahaye, C. Menotti, L. Santos, M. Lewenstein, and T. Pfau, Rep. Prog. Phys. **72**, 126401 (2009).

M.A. Baranov, Phys. Rep. **464**, 71 (2008).

87RbCs lowest rotational level



Why degenerate dipolar gases?



$$U_{dd}(\mathbf{r}) = \frac{C_{dd}}{4\pi} \frac{1 - 3 \cos^2 \theta}{r^3}$$



$$C_{dd} = \frac{d^2}{\epsilon_0} \quad \text{electric dipoles}$$

$$C_{dd} = \mu_0 \mu^2 \quad \text{magnetic dipoles}$$

$$\frac{\mu_0 \mu^2}{d^2 / \epsilon_0} \sim \alpha^2 \sim 10^{-4}$$

$$a_{dd} \equiv \frac{C_{dd} m}{12\pi \hbar^2} \quad \epsilon_{dd} \equiv \frac{a_{dd}}{a} = \frac{C_{dd}}{3g}$$

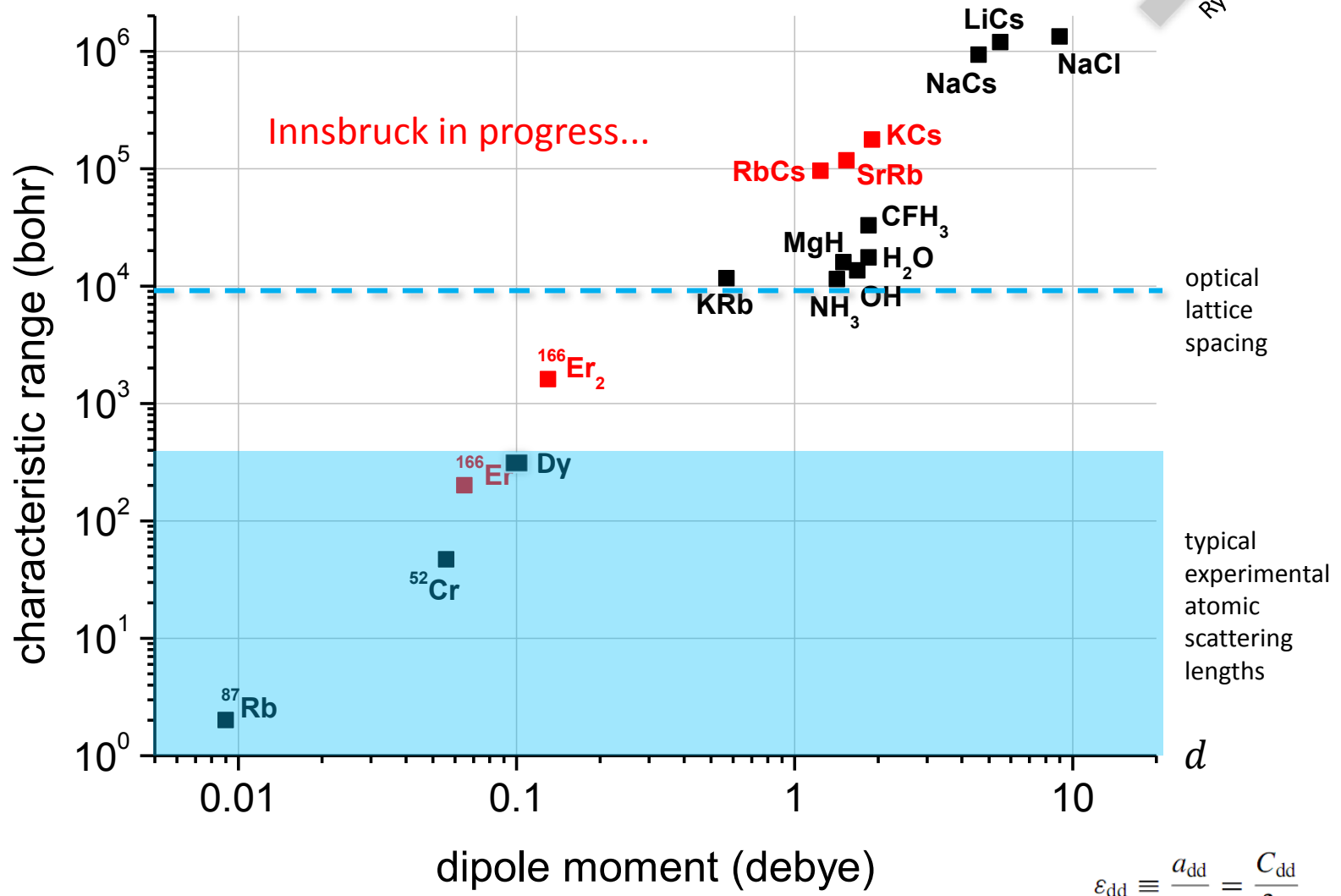
Review articles:
 T. Lahaye *et al.*, Rep. Prog. Phys. **72**, 126401 (2009).
 M.A. Baranov, Phys. Rep. **464**, 71 (2008).

Why degenerate dipolar gases?

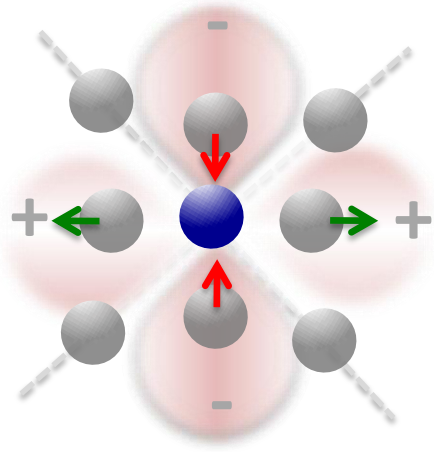


$$a_{dd} = \frac{md^2}{4\pi\epsilon_0\hbar^2}$$

dipolar gases

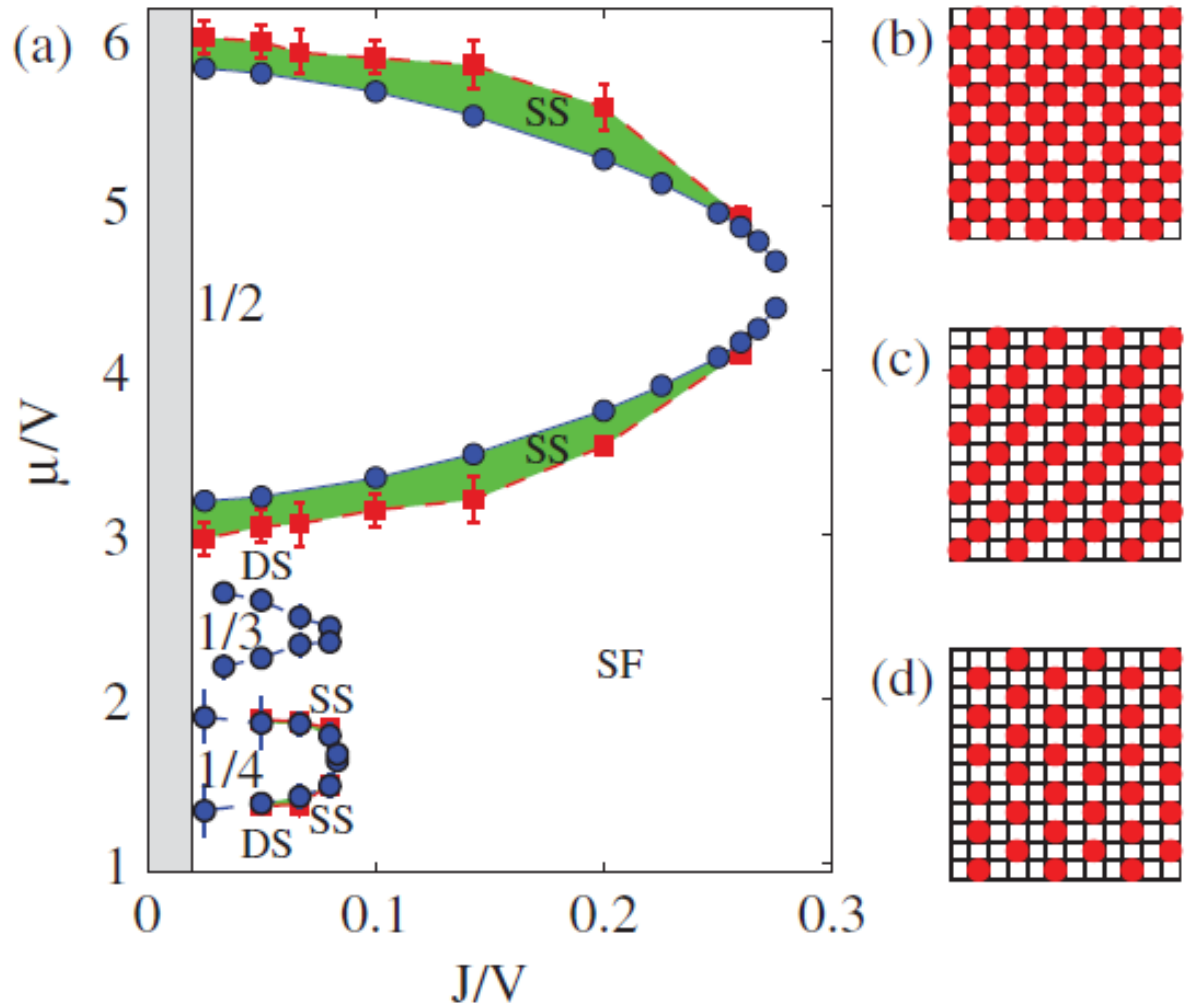


$$\epsilon_{dd} \equiv \frac{a_{dd}}{a} = \frac{C_{dd}}{3g}$$

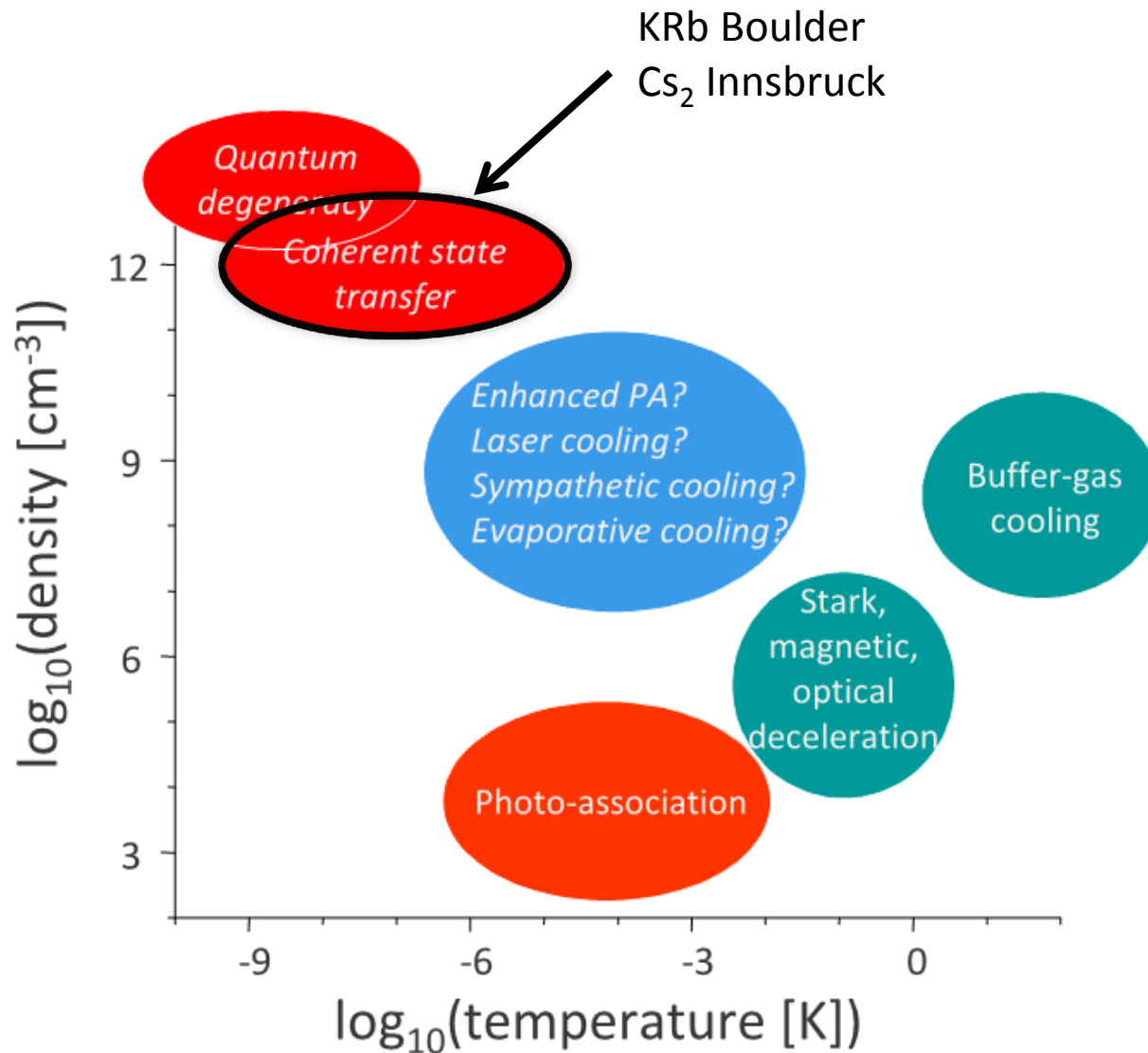


Novel effects appear – NOT contact interaction!

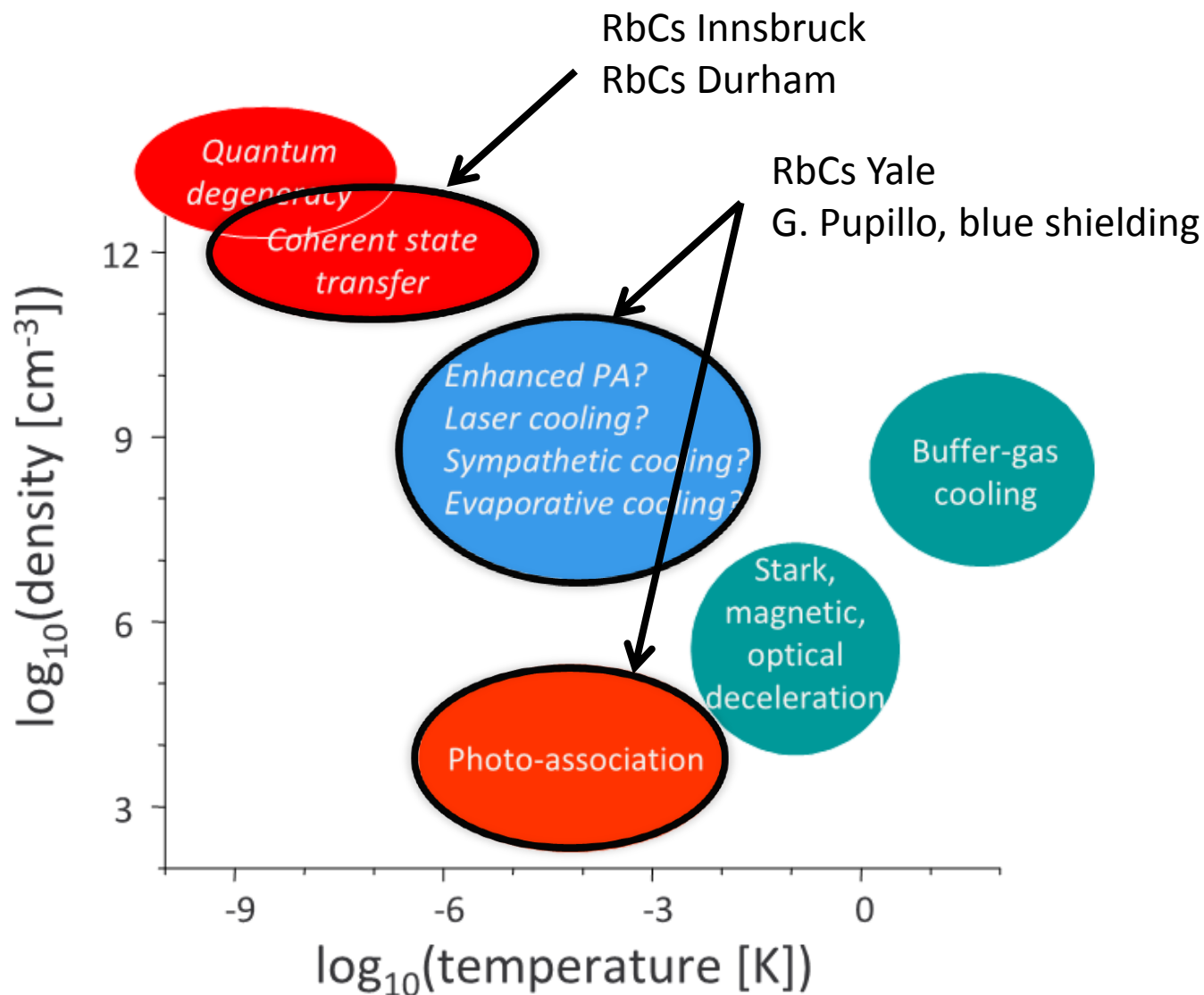
- Identical fermions interact • exotic quantum phases •
- novel spectrum of excitations • geometry-dependent interaction • ...



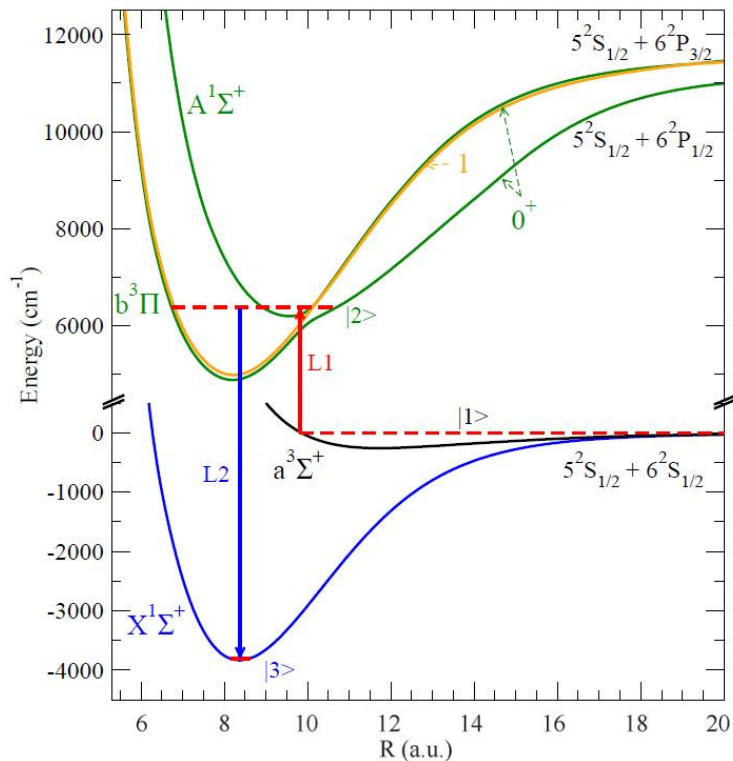
Why degenerate dipolar gases?



Why degenerate dipolar gases?

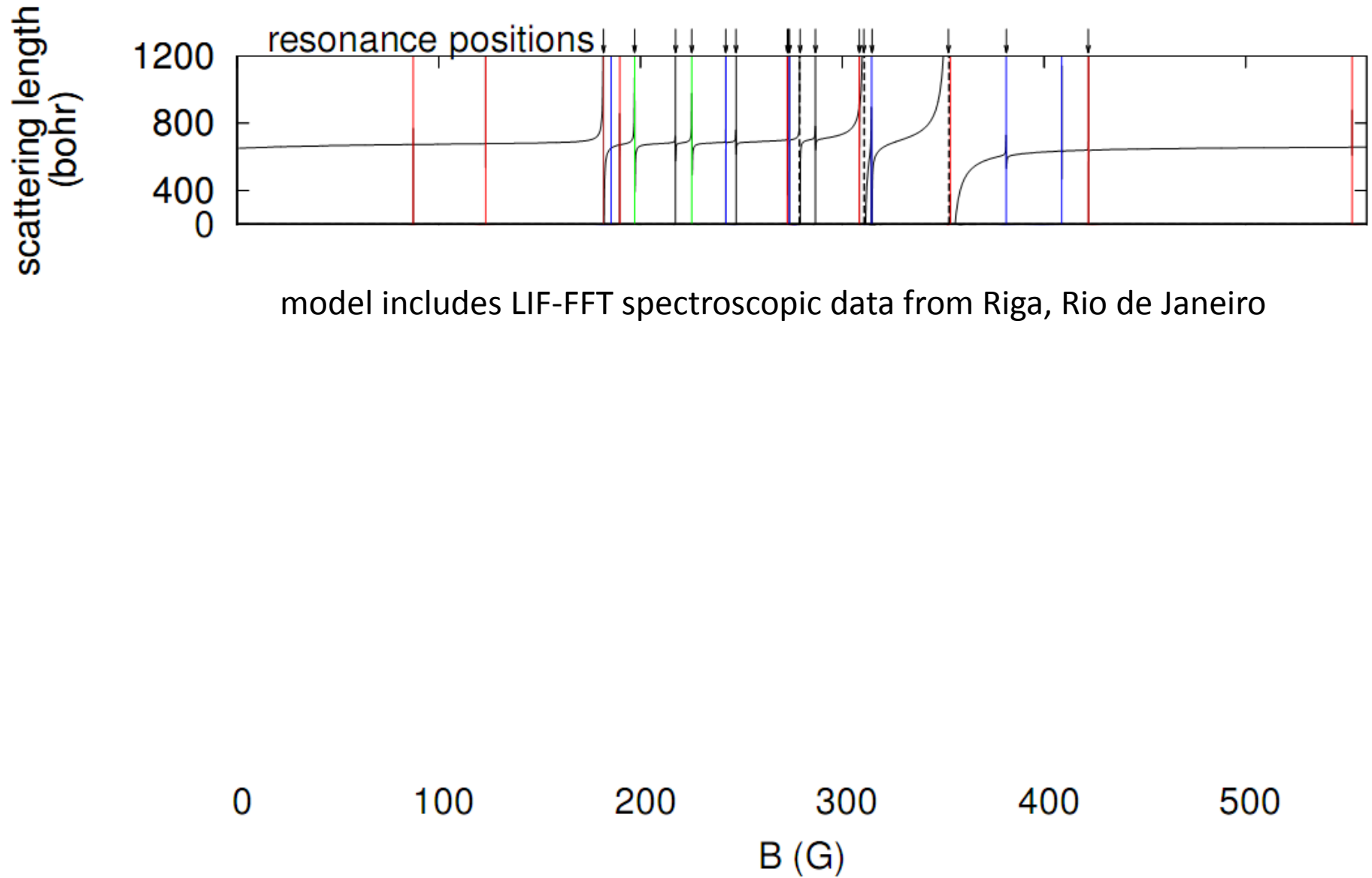


degenerate atomic gas(es) → Feshbach molecules → coherent ground state transfer

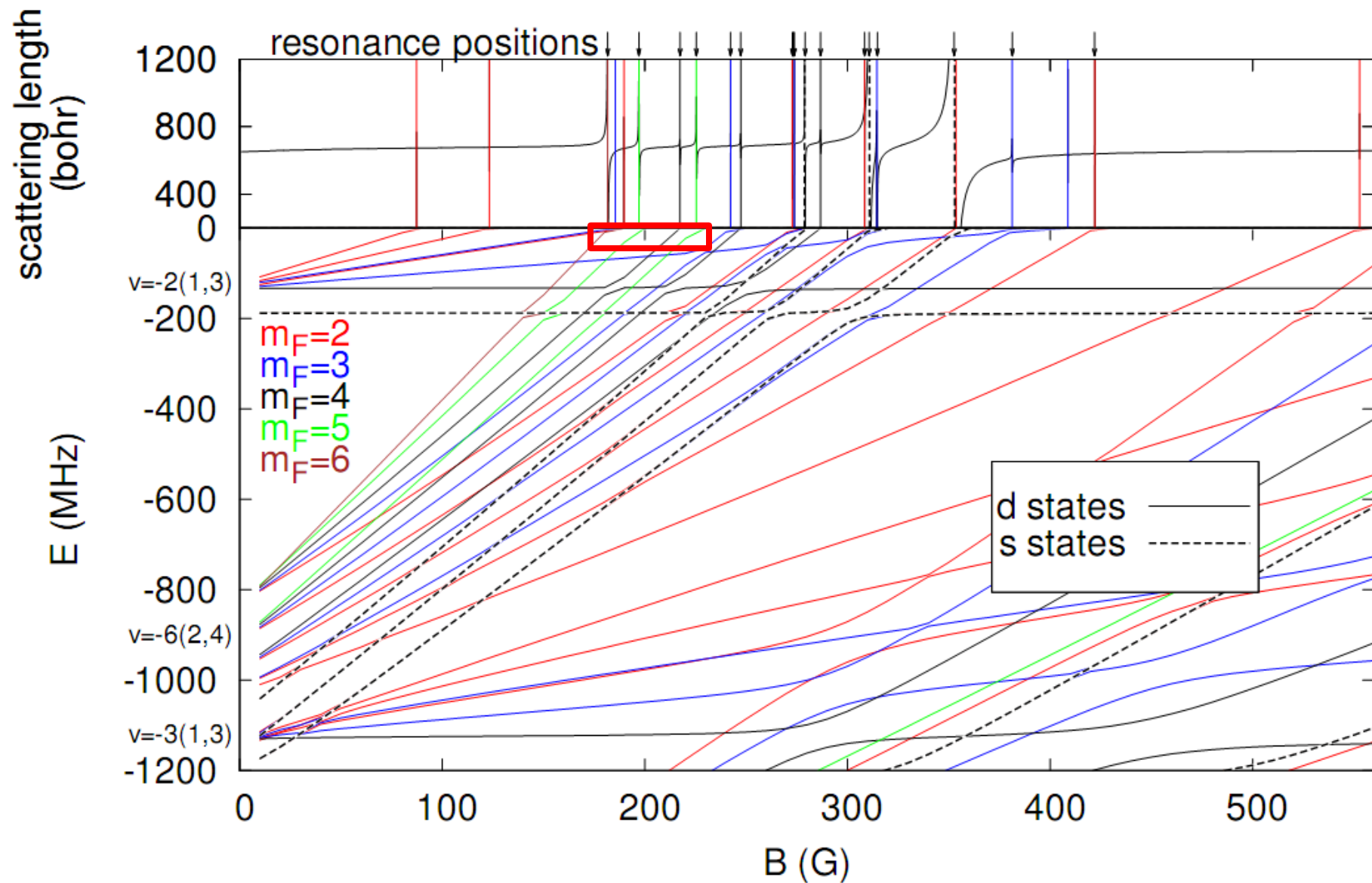


- high phase space density almost guaranteed
- high experimental complexity
- limited to "boring" molecules, i.e. dimers
- successfully applied thus far to
KRb, Cs₂, Rb₂ (³Σ), ...RbCs

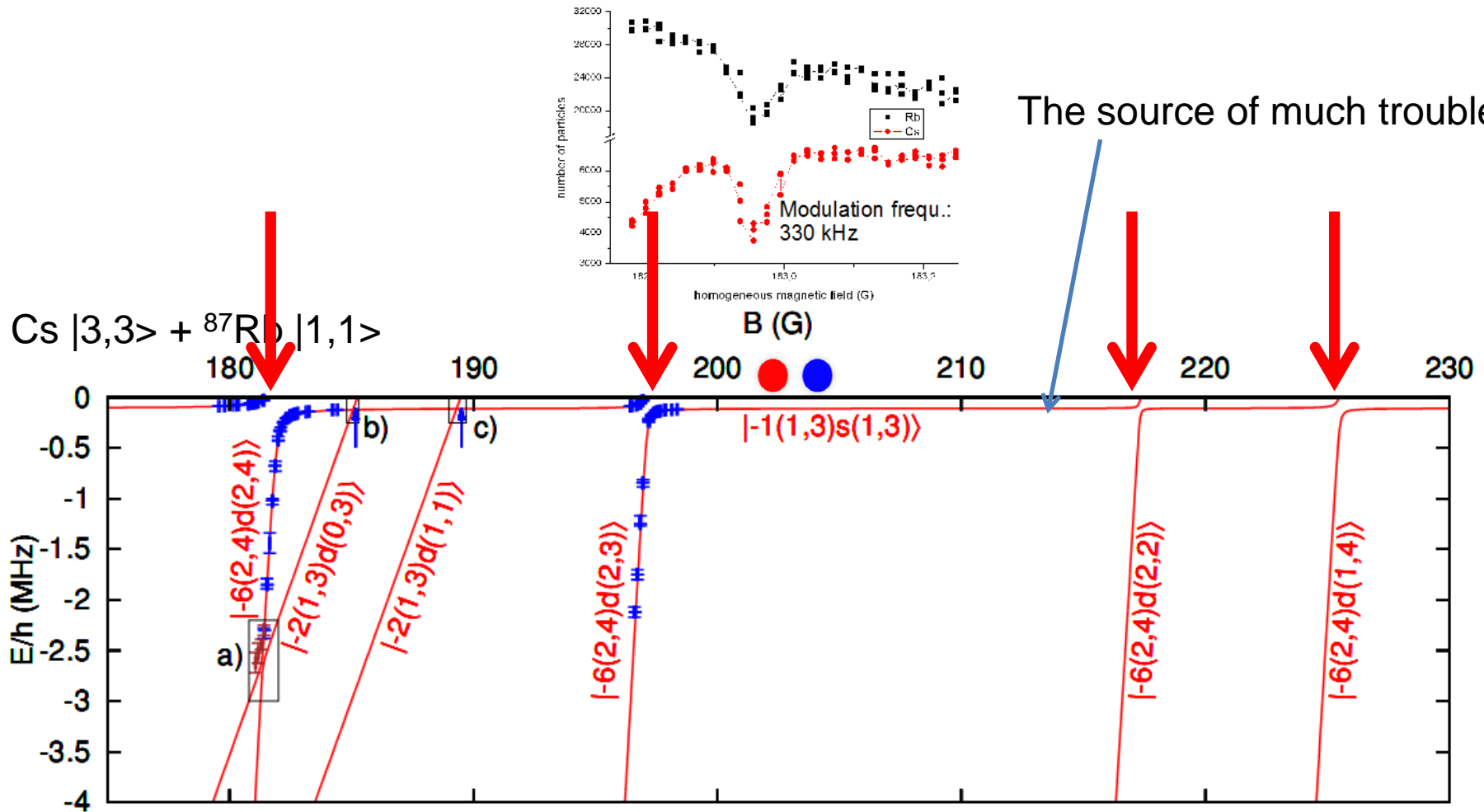


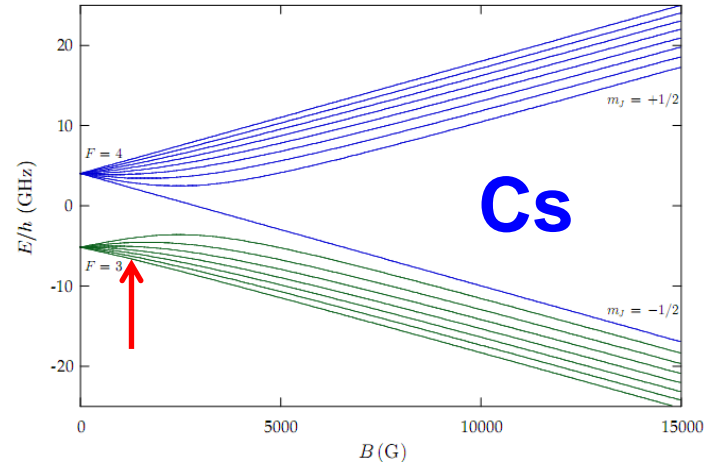
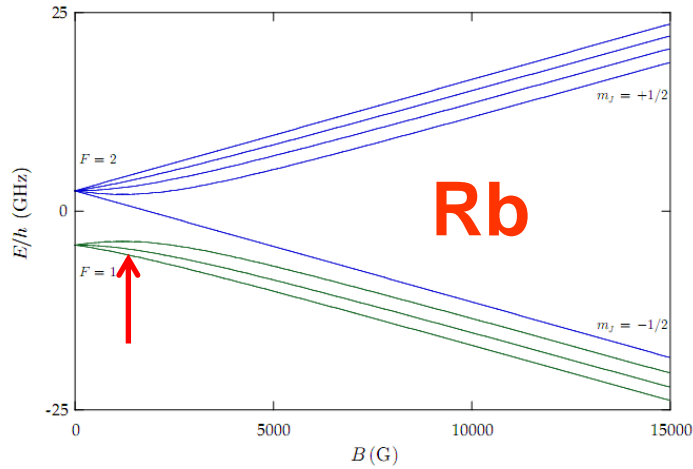


model includes LIF-FFT spectroscopic data from Riga, Rio de Janeiro

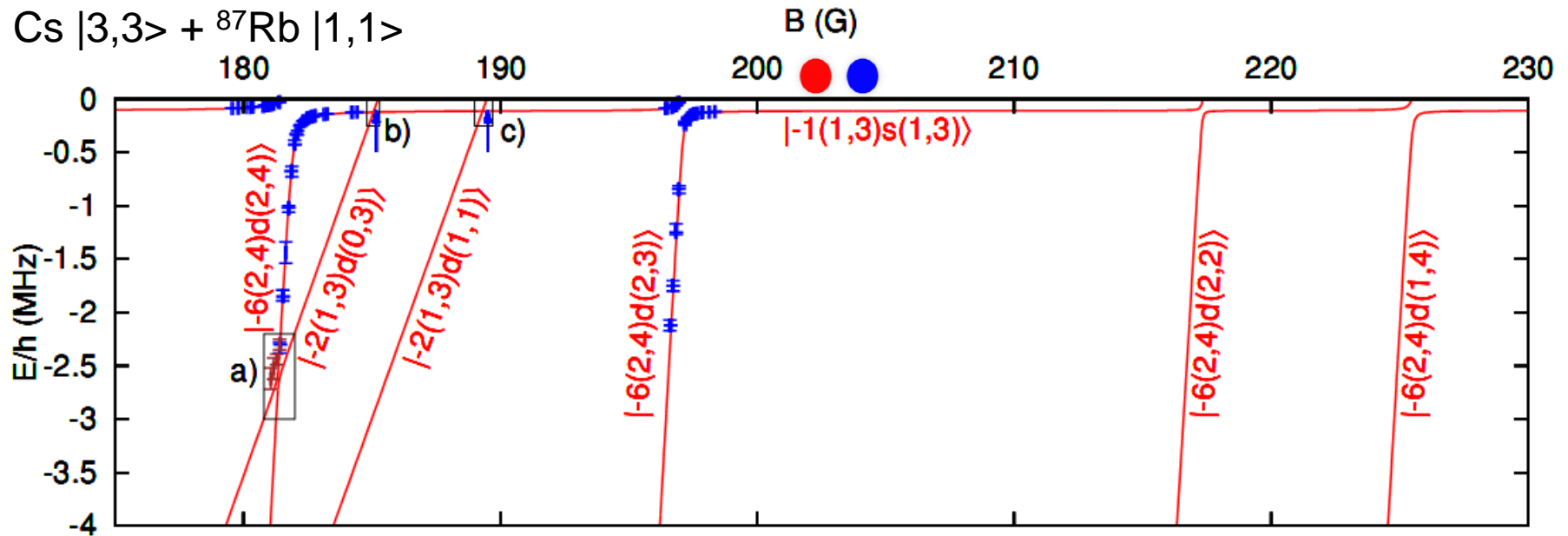


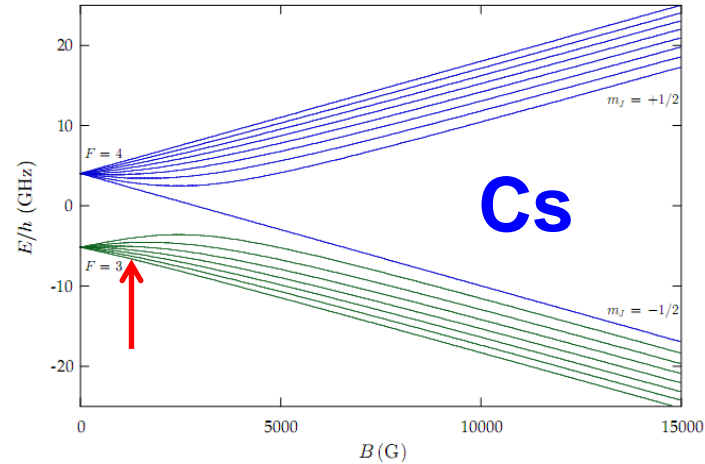
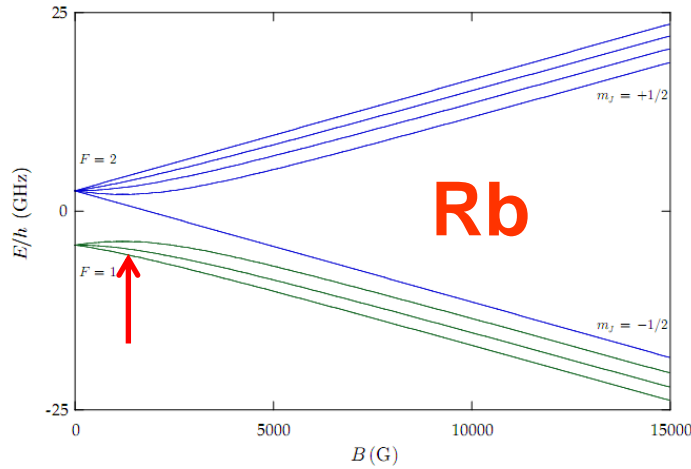
Data for model – Feshbach molecule binding energies through magnetic field modulation



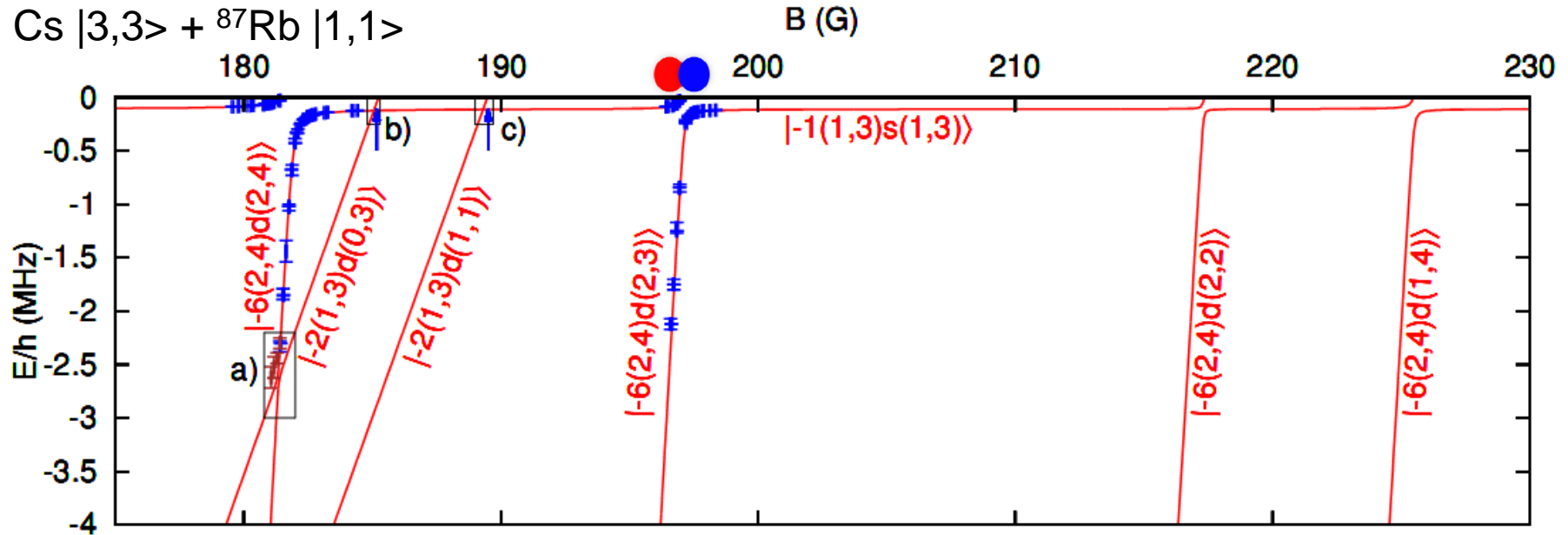


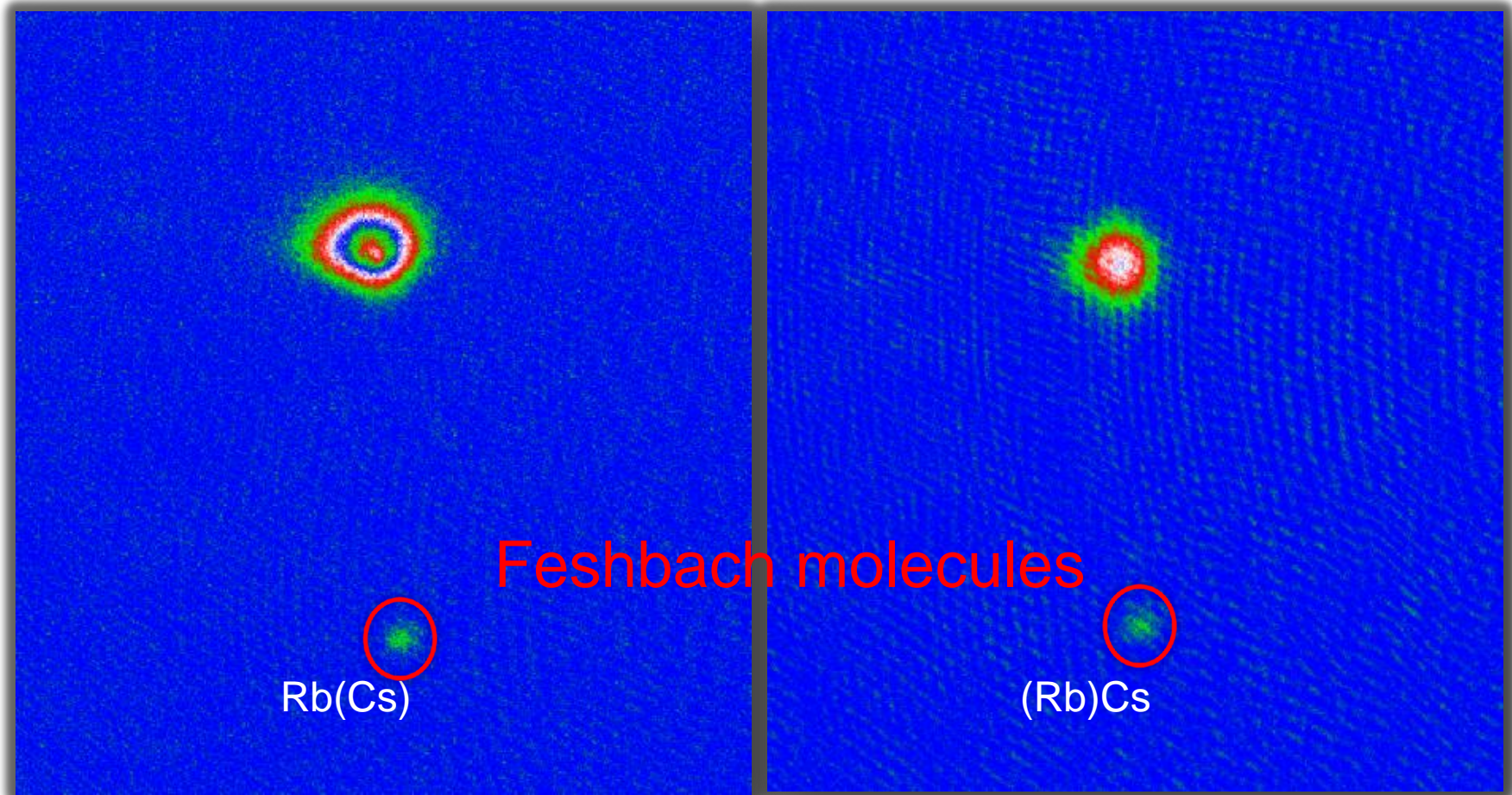
Cs $|3,3\rangle + {}^{87}\text{Rb } |1,1\rangle$





Cs $|3,3\rangle + {}^{87}\text{Rb } |1,1\rangle$





Typically 60k Cs + 150k Rb gives 4000 RbCs (we detect only atoms)

Hamiltonian:

$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 0 & \Omega_2 \\ 0 & \Omega_2 & 0 \end{bmatrix} \begin{matrix} |1\rangle \\ |2\rangle \\ |3\rangle \end{matrix}$$

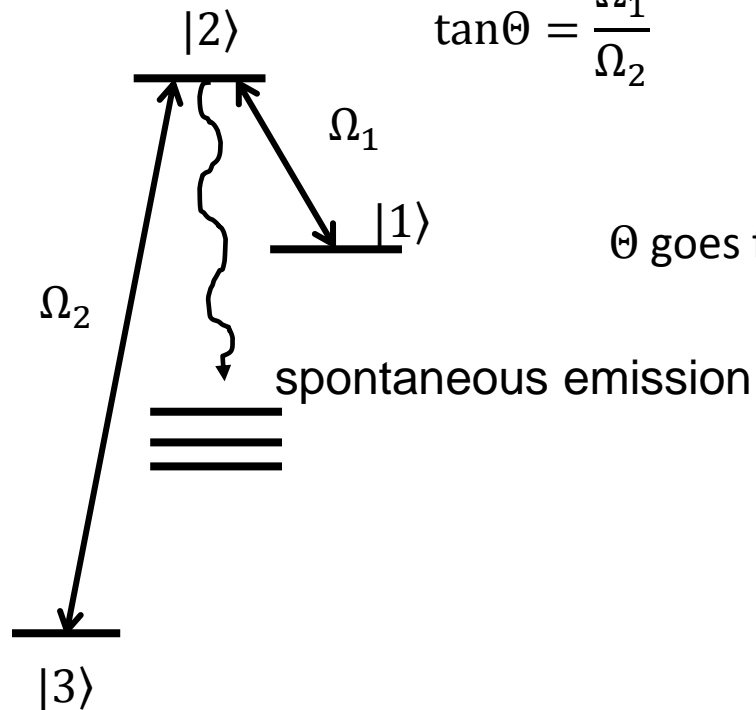
Eigenstates with light on:

$$|a^+\rangle = \sin \Theta \sin \Phi |1\rangle + \cos \Phi |2\rangle + \cos \Theta \sin \Phi |3\rangle$$

$$|a^0\rangle = \cos \Theta |1\rangle - \sin \Theta |3\rangle \text{ dark state}$$

$$|a^-\rangle = \sin \Theta \cos \Phi |1\rangle - \sin \Phi |2\rangle + \cos \Theta \cos \Phi |3\rangle$$

$$\tan \Theta = \frac{\Omega_1}{\Omega_2}$$



Θ goes from 0 to Π adiabatically

Hamiltonian:

$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 0 & \Omega_2 \\ 0 & \Omega_2 & 0 \end{bmatrix} \begin{matrix} |1\rangle \\ |2\rangle \\ |3\rangle \end{matrix}$$

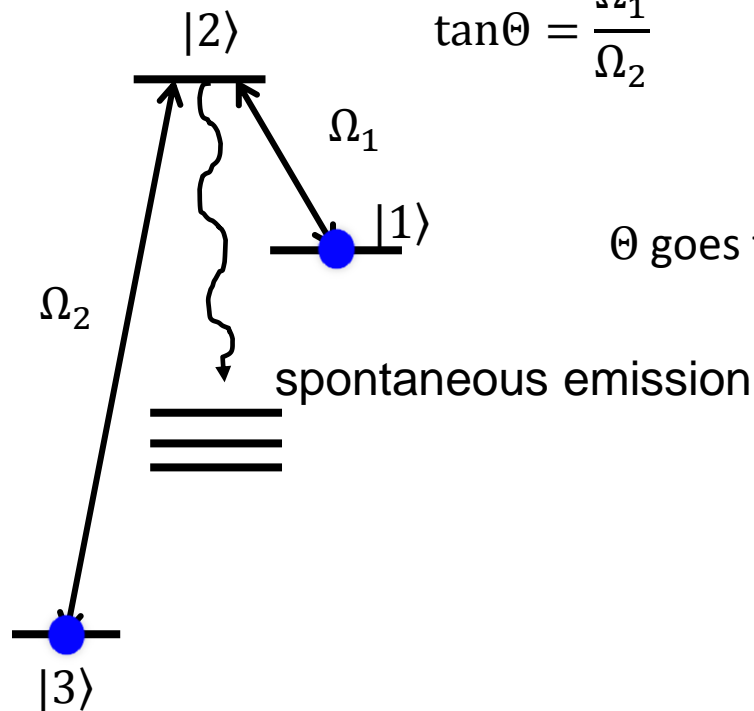
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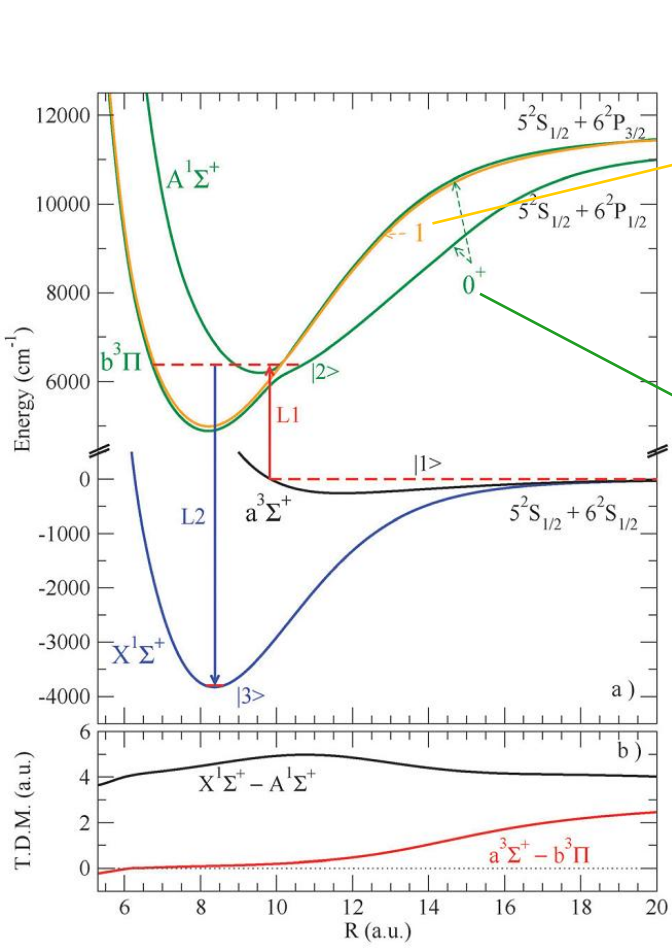
$$\tan \Theta = \frac{\Omega_1}{\Omega_2}$$



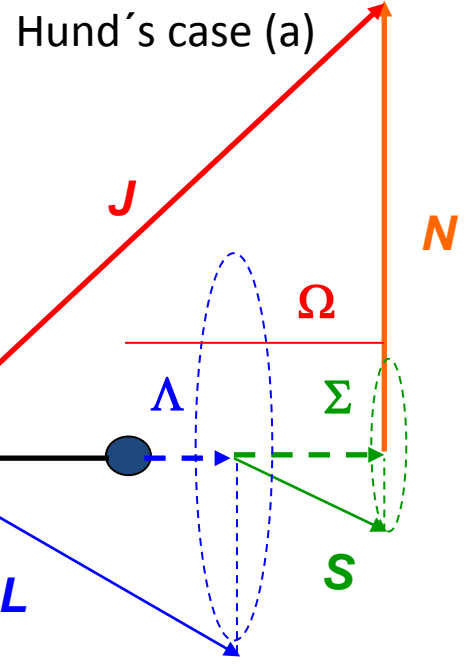
Θ goes from 0 to Π adiabatically

Ground state transfer

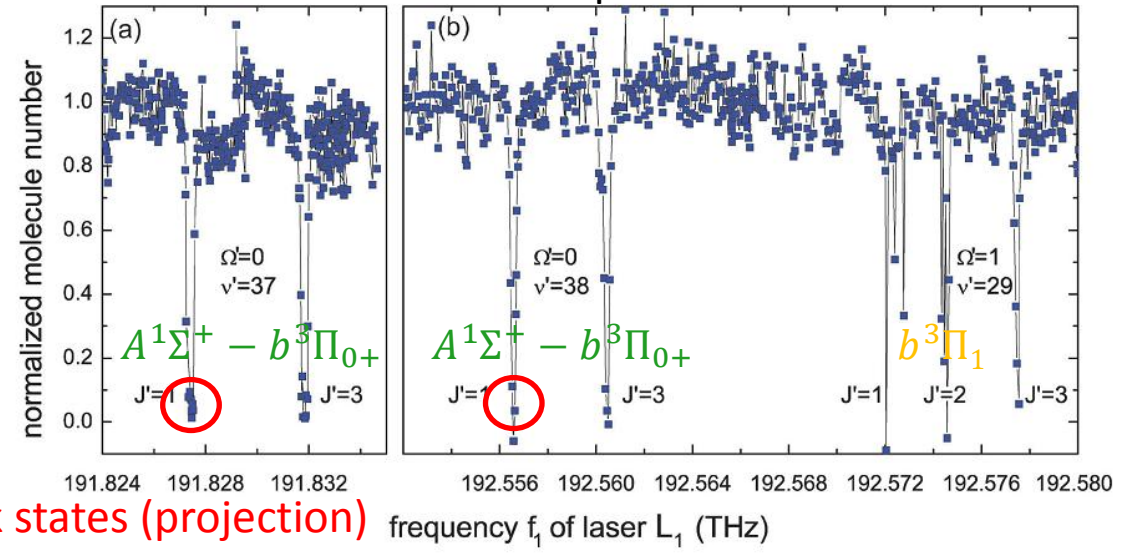
Suggested Bergeman *et al.*, PRA **67** 050501 (2003). (Riga, Rio de Janeiro)



$S = 1, \Sigma = -1, 0, 1$
 $\Lambda = 1$
 $\Omega = \Lambda + \Sigma = 0, 1, 2$
 $A^1\Sigma^+ - b^3\Pi_{0+}$
 spin-orbit mixing $\Delta\Omega = 0, \Delta J = 0$



Feshbach molecule absorption



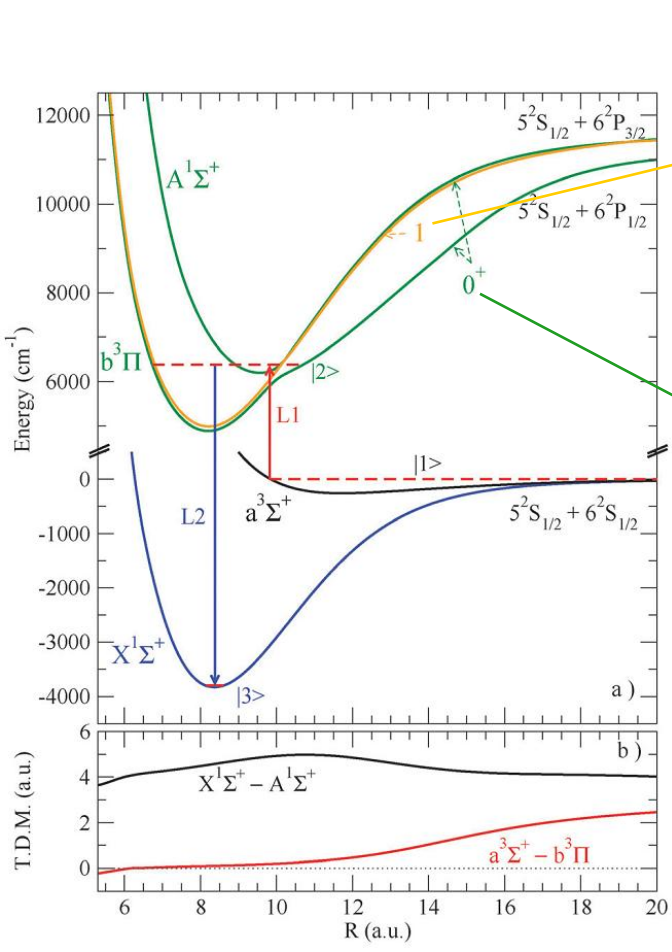
dark states (projection)

Energies and rotational constants agree well with Docenko *et al.*, PRA **81** 042511 (2010). (Riga, Rio de Janeiro)

Ground state transfer

Suggested Bergeman *et al.*, PRA **67** 050501 (2003). (Riga, Rio de Janeiro)

Hund's case (a)

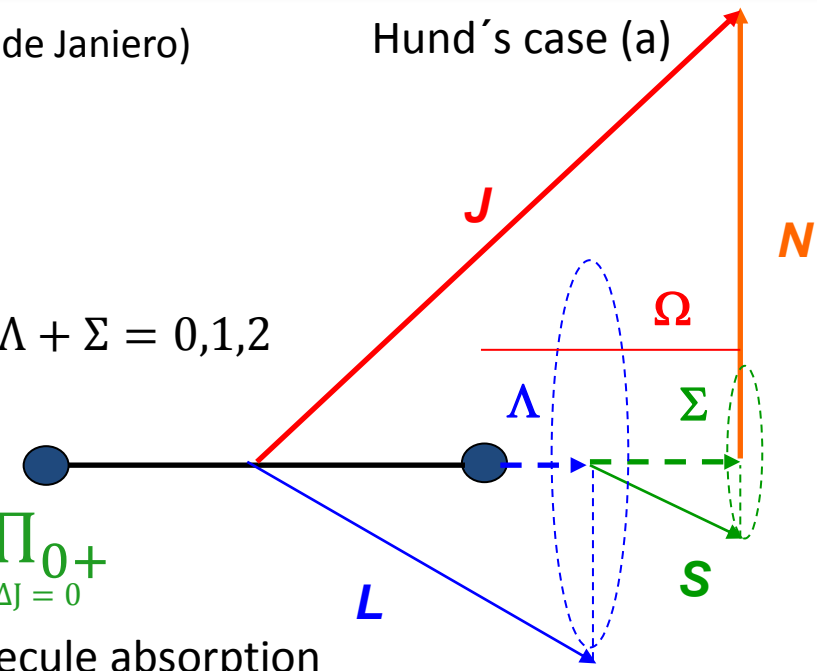
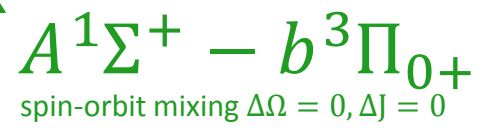


$$S = 1, \Sigma = -1, 0, 1$$

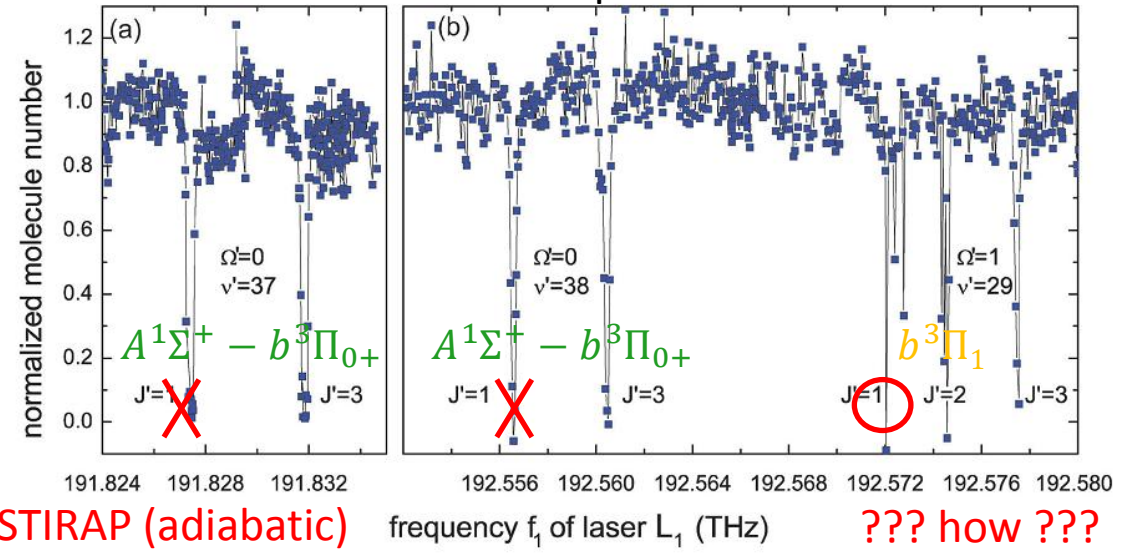


$$\Omega = \Lambda + \Sigma = 0, 1, 2$$

$$\Lambda = 1$$



Feshbach molecule absorption

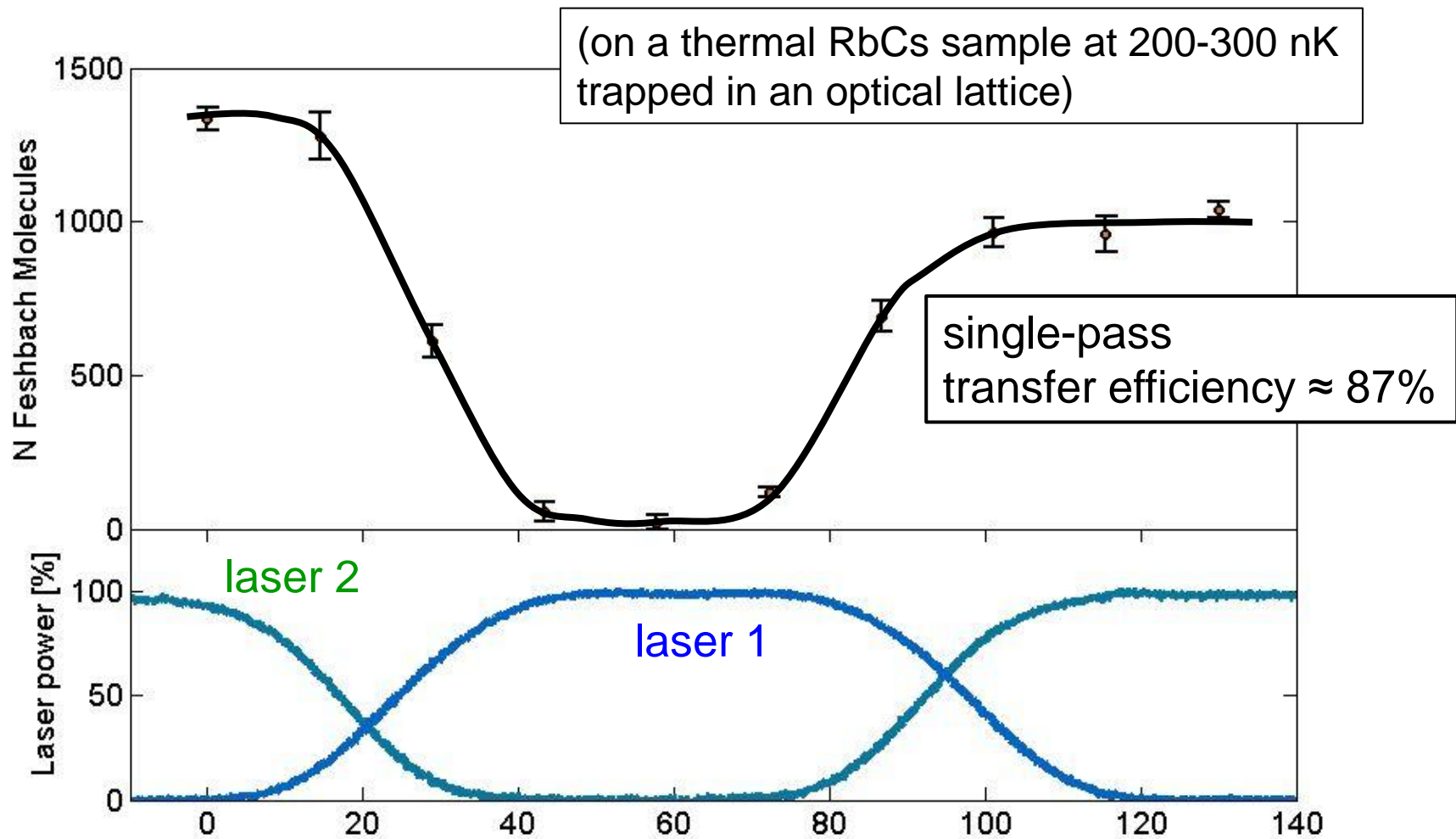


??? why ???

STIRAP (adiabatic)

??? how ???

Energies and rotational constants agree well with Docenko *et al.*, PRA **81** 042511 (2010). (Riga, Rio de Janeiro)

RbCs two-photon STIRAP to $v=0, J=0$.

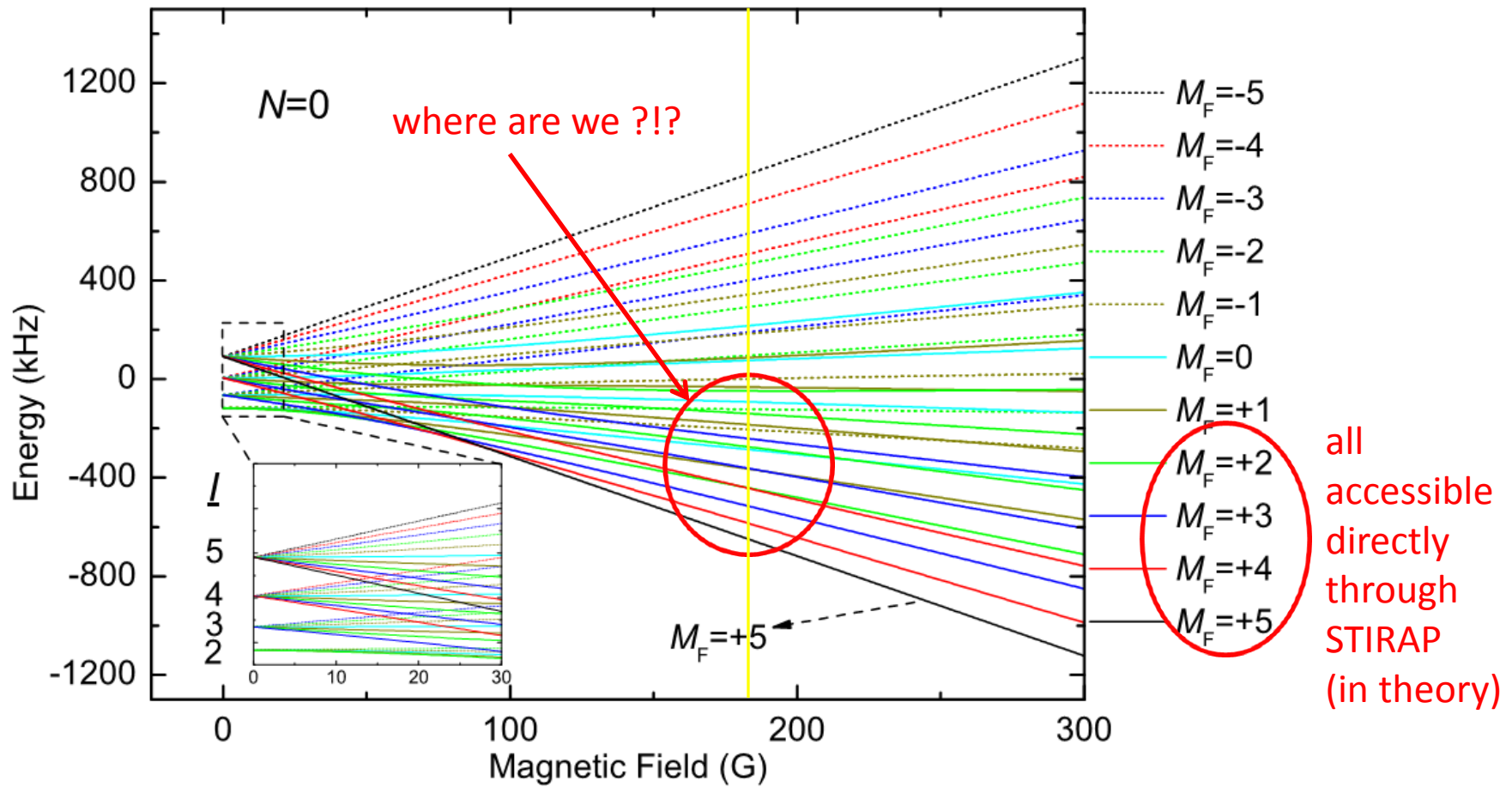
- we detect only atoms
- STIRAP references -- two optical cavities locked to Cs atomic reference laser
- estimated relative laser linewidth: 5-10kHz



Ground state transfer

Cs $|3,3\rangle + {}^{87}\text{Rb } |1,1\rangle$ in incoming s-wave collision has $M_F=4$, therefore, Feshbach molecules also have $M_F=4$

${}^{87}\text{Rb } {}^{133}\text{Cs}$

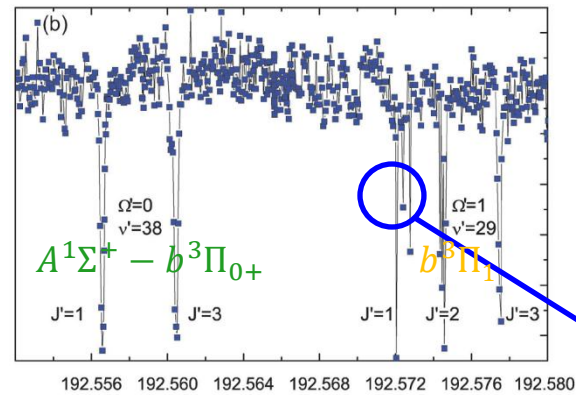


dominant terms – scalar nuclear dipole-dipole, nuclear Zeeman
 J. Aldegunde (Salamanca) and Jeremy M. Hutson (Durham)

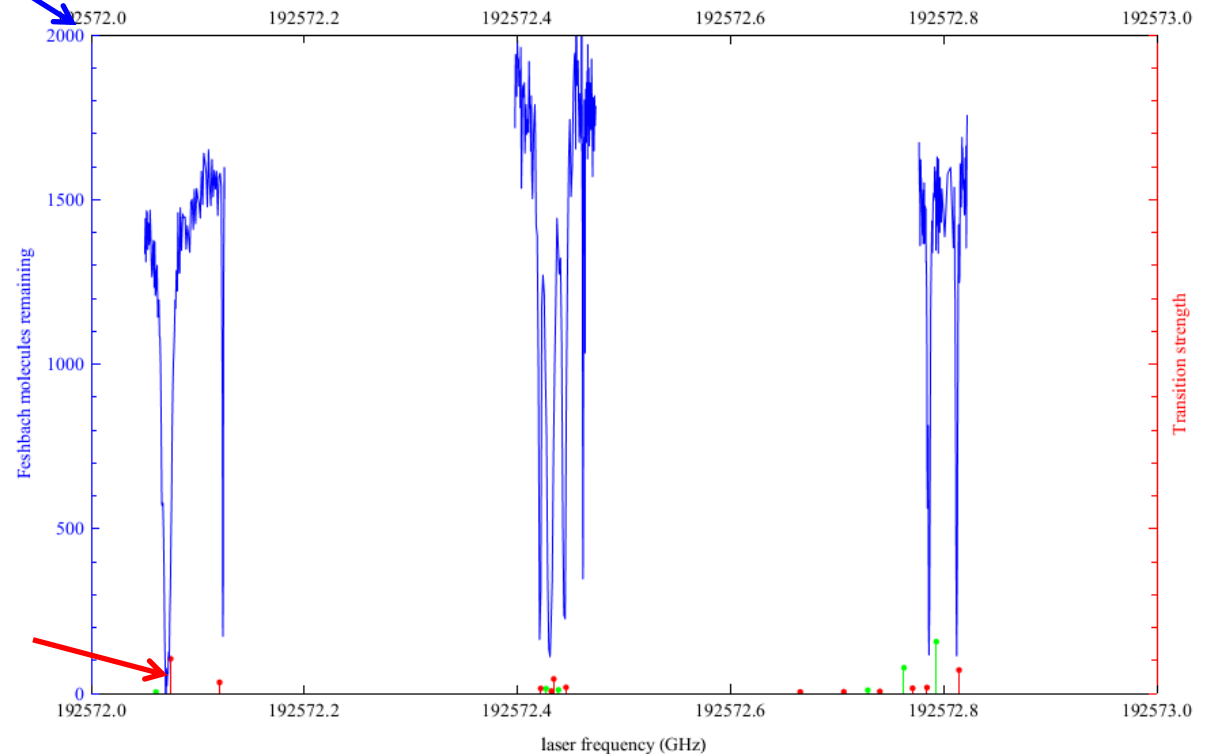
Ground state transfer

Excited state model

(help from Romain Vexieu, Anne Crubelier, Oliver Dulieu)



$^{87}\text{RbCs } b^3\Pi_1 v'=29 J=1$ 217G h-pol
 red=calculated transition strengths to M=3,5 (hpol), green=calculated transition strengths to M=4 (vpol)
 exposures 1.8V 1ms, 1.8V 2ms, 2V 1ms



$$H = H_{\text{rotation}} + H_{\text{Zeeman}} + H_{\text{hf}}$$

$$H_{\text{hf}} = a_{\text{Rb}} \mathbf{i}_{\text{Rb}} \cdot \mathbf{L} + a_{\text{Cs}} \mathbf{i}_{\text{Cs}} \cdot \mathbf{L}$$

3 parameter fit to effective Hamiltonian:

$$a_{\text{Rb}} = 127 \text{ MHz} \times h$$

$$a_{\text{Cs}} = 74 \text{ MHz} \times h$$

overall frequency shift

(1st unambiguous
 observation of orbital
 hyperfine in bialkalis)

FIG. 3: 217G *Feshbach* $\rightarrow b^3\Pi_1 v' = 29, J' = 1$ horizontal laser polarization. Red lines indicate calculated transition strengths.

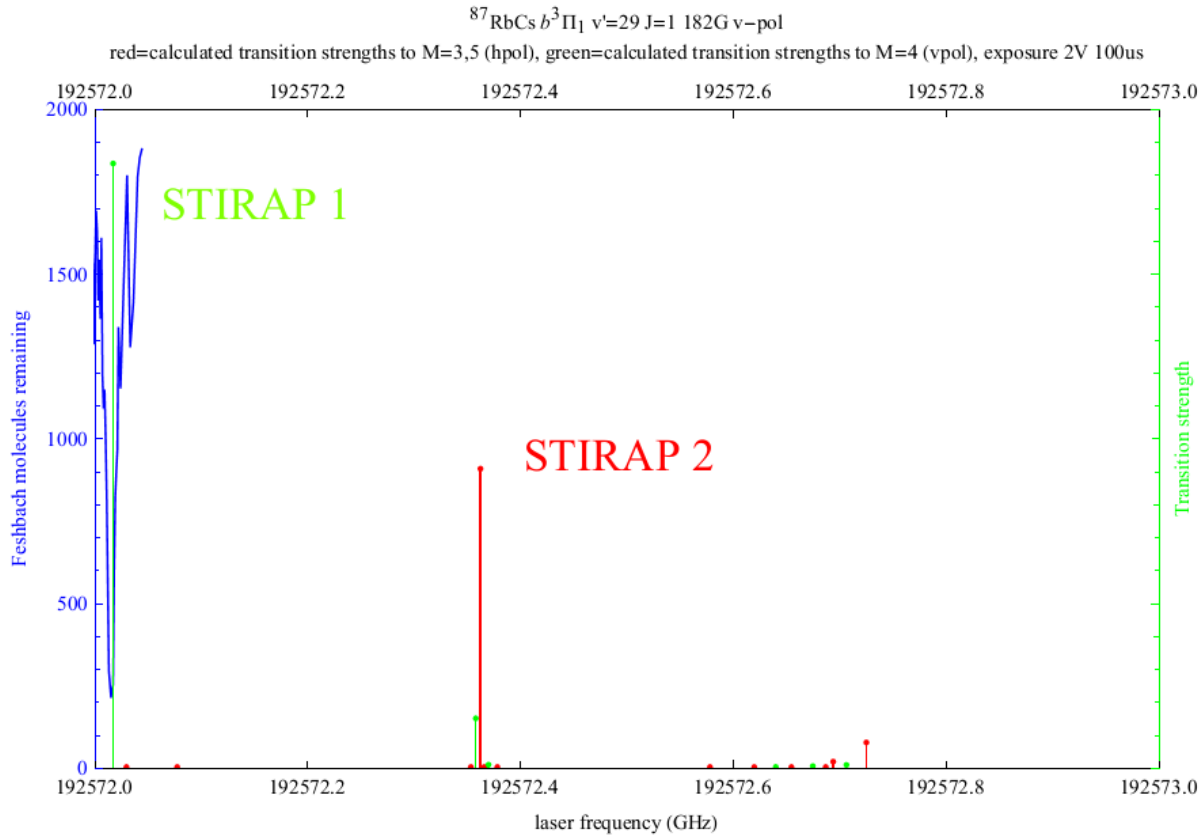


FIG. 4: 182G $Feshbach \rightarrow b^3\Pi_1 \ v' = 29, J' = 1$ vertical laser polarization. The lowest-frequency peak here (STIRAP 1) is currently used for STIRAP. Green lines indicate calculated transition strengths.

TABLE II: 182G $b^3\Pi_1 \ v' = 29, J' = 1$ expectation values. (Green lines from left to right in Fig. 4.)

$\langle \hat{n}_{i_{Rb}} \rangle$	1.45827	1.10556	0.79239	-0.21970	0.64279	1.21344
$\langle \hat{n}_{i_{Cs}} \rangle$	3.46557	2.80281	3.1212	3.25451	2.43313	1.91451
$\langle \hat{n}_J \rangle$	-0.92384	0.09163	0.08642	0.96519	0.92408	0.87205
$\langle \hat{M} \rangle$	4	4	4	4	4	4
$\langle \hat{P}(\delta) \rangle$	0.92671	0.06771	0.00303	0.00019	0.00130	0.00268
$\langle \hat{P}(\epsilon) \rangle$	0.04354	0.23208	0.57752	0.03096	0.08868	0.02533
$\langle \hat{P}(\zeta) \rangle$	0.03590	0.51521	0.31557	0.00686	0.00002	0.12472
$\langle \hat{P}(\theta) \rangle$	0.00089	0.02827	0.07236	0.70694	0.16983	0.01726
$\langle \hat{P}(\iota) \rangle$	0.00201	0.09986	0.00123	0.22289	0.40956	0.25803
$\langle \hat{P}(\kappa) \rangle$	0.00065	0.04539	0.02830	0.02986	0.32579	0.56352

(δ) denotes projection onto $m_{J'} = -1, m_{i_{Rb}} = \frac{3}{2}, m_{i_{Cs}} = \frac{7}{2}$

(ϵ) denotes projection onto $m_{J'} = 0, m_{i_{Rb}} = \frac{3}{2}, m_{i_{Cs}} = \frac{7}{2}$

(ζ) denotes projection onto $m_{J'} = 0, m_{i_{Rb}} = \frac{1}{2}, m_{i_{Cs}} = \frac{7}{2}$

(θ) denotes projection onto $m_{J'} = 1, m_{i_{Rb}} = -\frac{1}{2}, m_{i_{Cs}} = \frac{7}{2}$

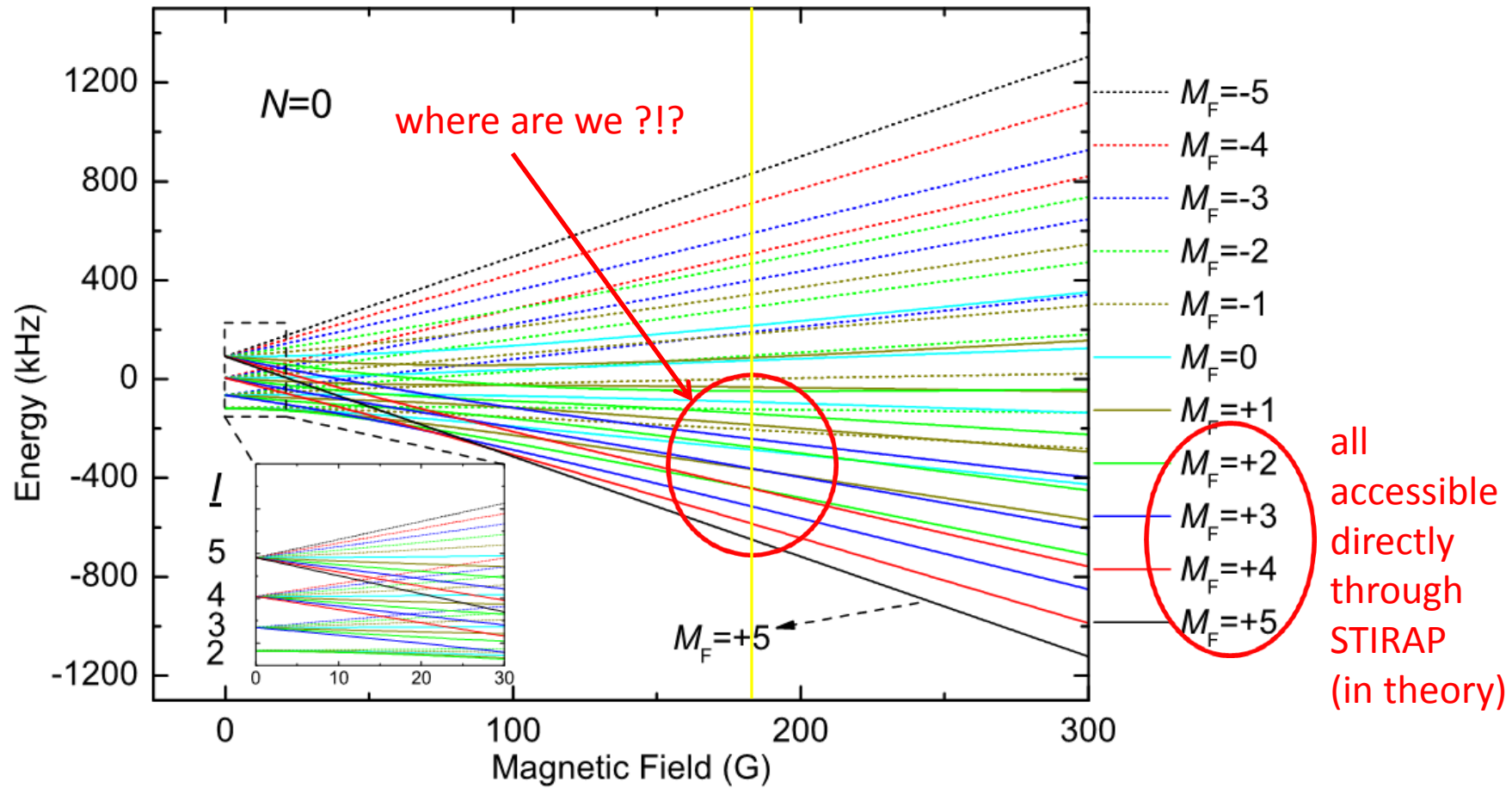
(ι) denotes projection onto $m_{J'} = 1, m_{i_{Rb}} = \frac{1}{2}, m_{i_{Cs}} = \frac{7}{2}$

(κ) denotes projection onto $m_{J'} = 1, m_{i_{Rb}} = \frac{3}{2}, m_{i_{Cs}} = \frac{5}{2}$

Ground state transfer

Cs $|3,3\rangle + {}^{87}\text{Rb } |1,1\rangle$ in incoming s-wave collision has $M_F=4$, therefore, Feshbach molecules also have $M_F=4$

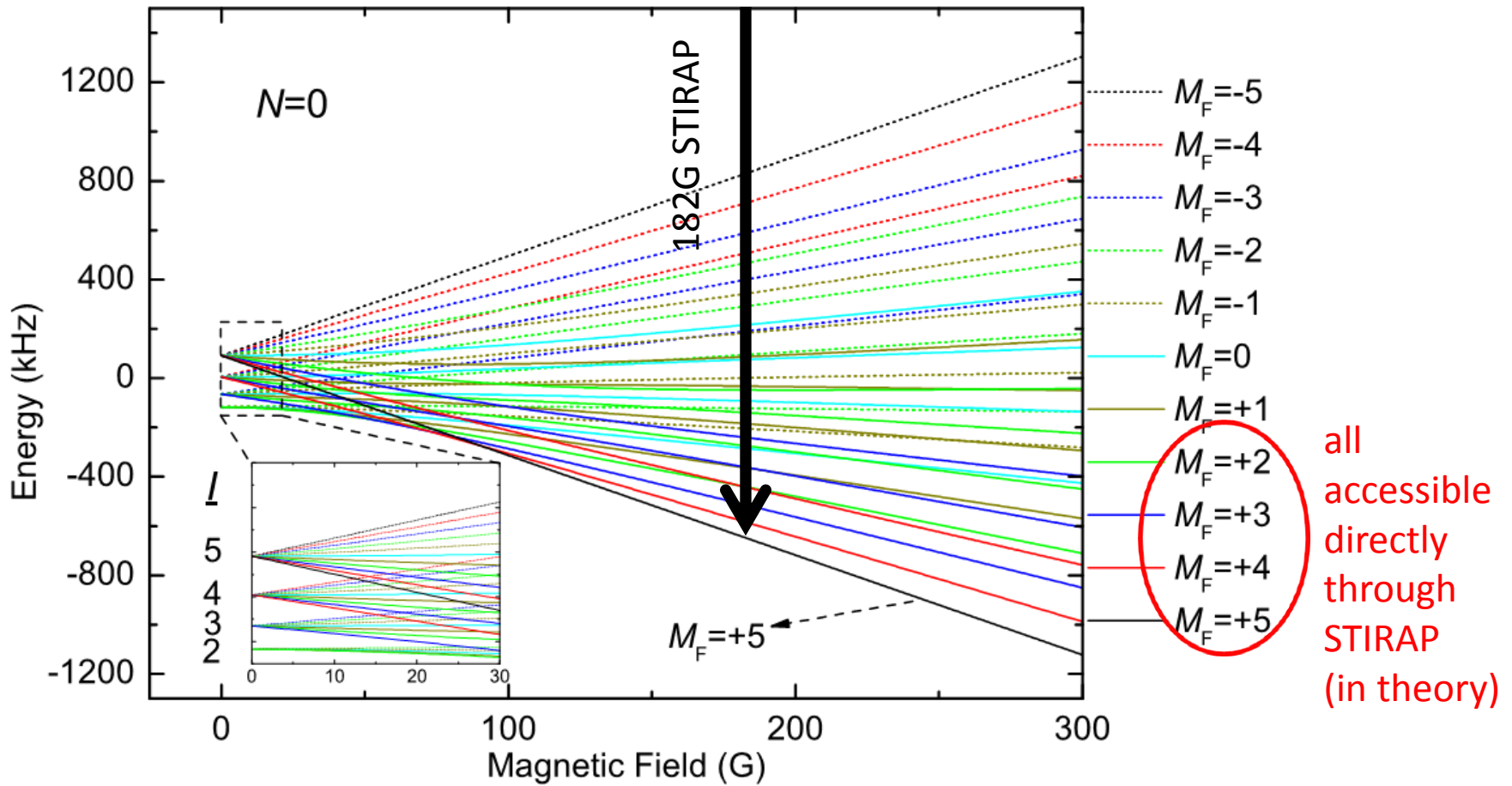
${}^{87}\text{Rb } {}^{133}\text{Cs}$

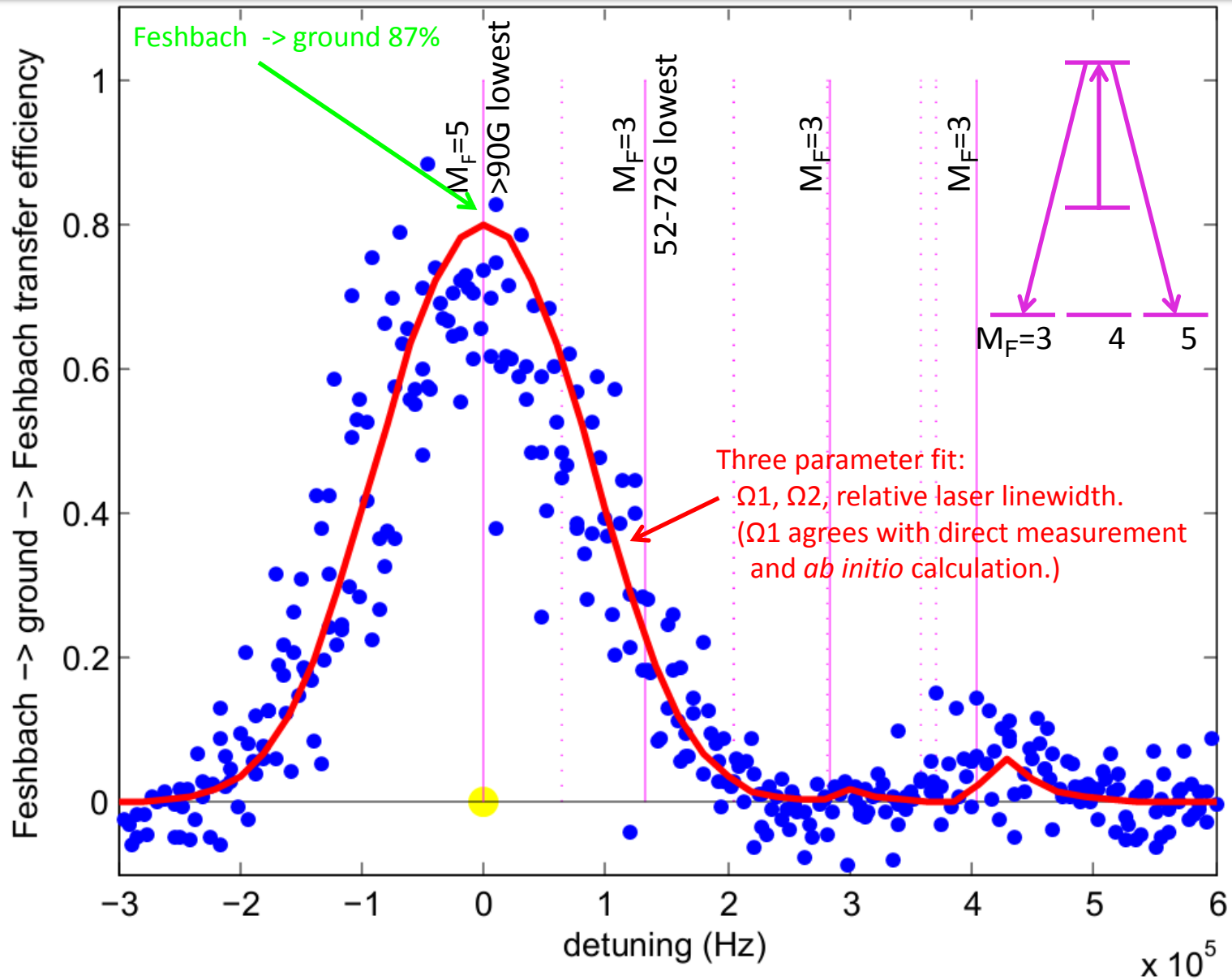


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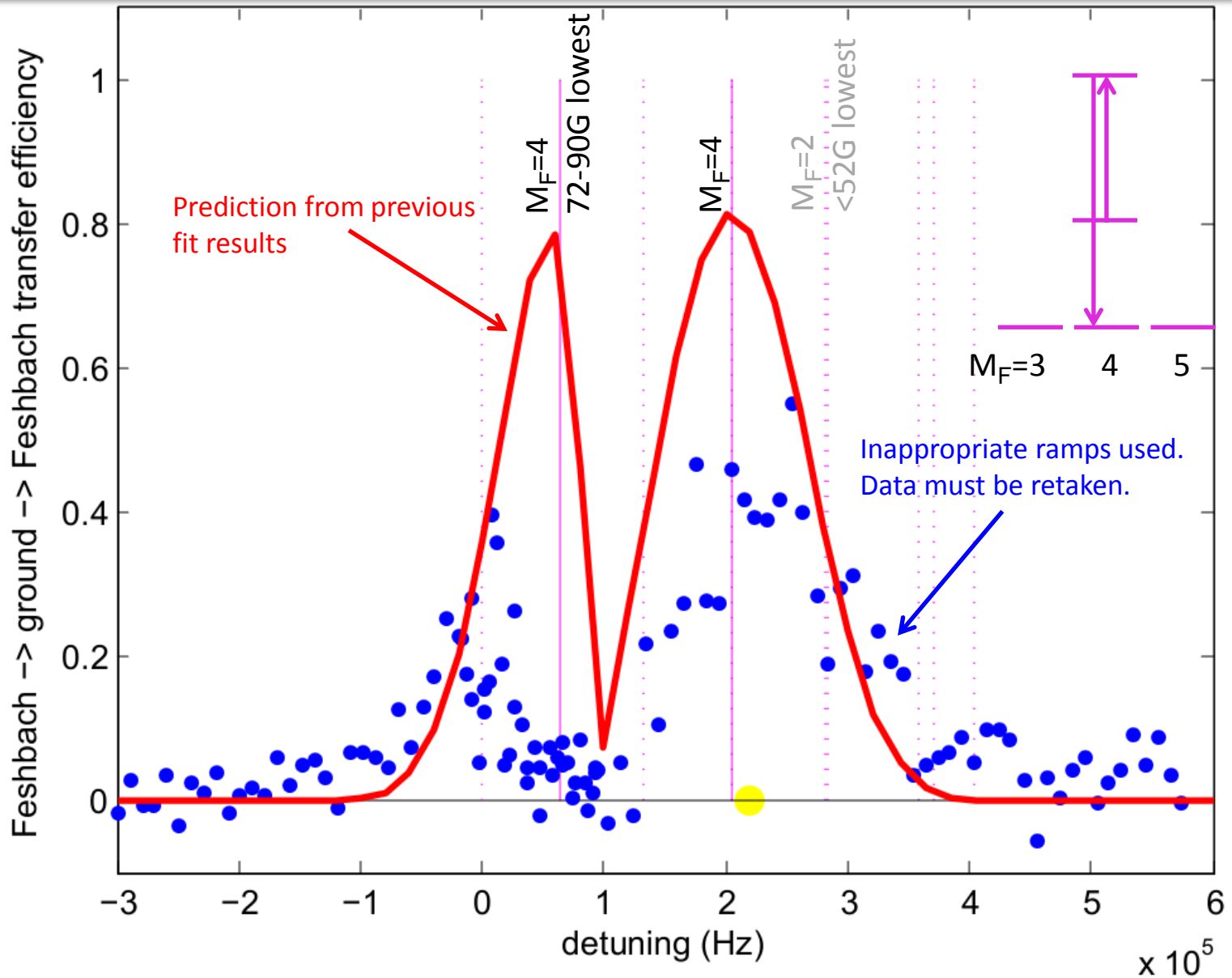
Cs $|3,3\rangle + {}^{87}\text{Rb} |1,1\rangle$ in incoming s-wave collision has $M_F=4$, therefore, Feshbach molecules also have $M_F=4$

${}^{87}\text{Rb} {}^{133}\text{Cs}$





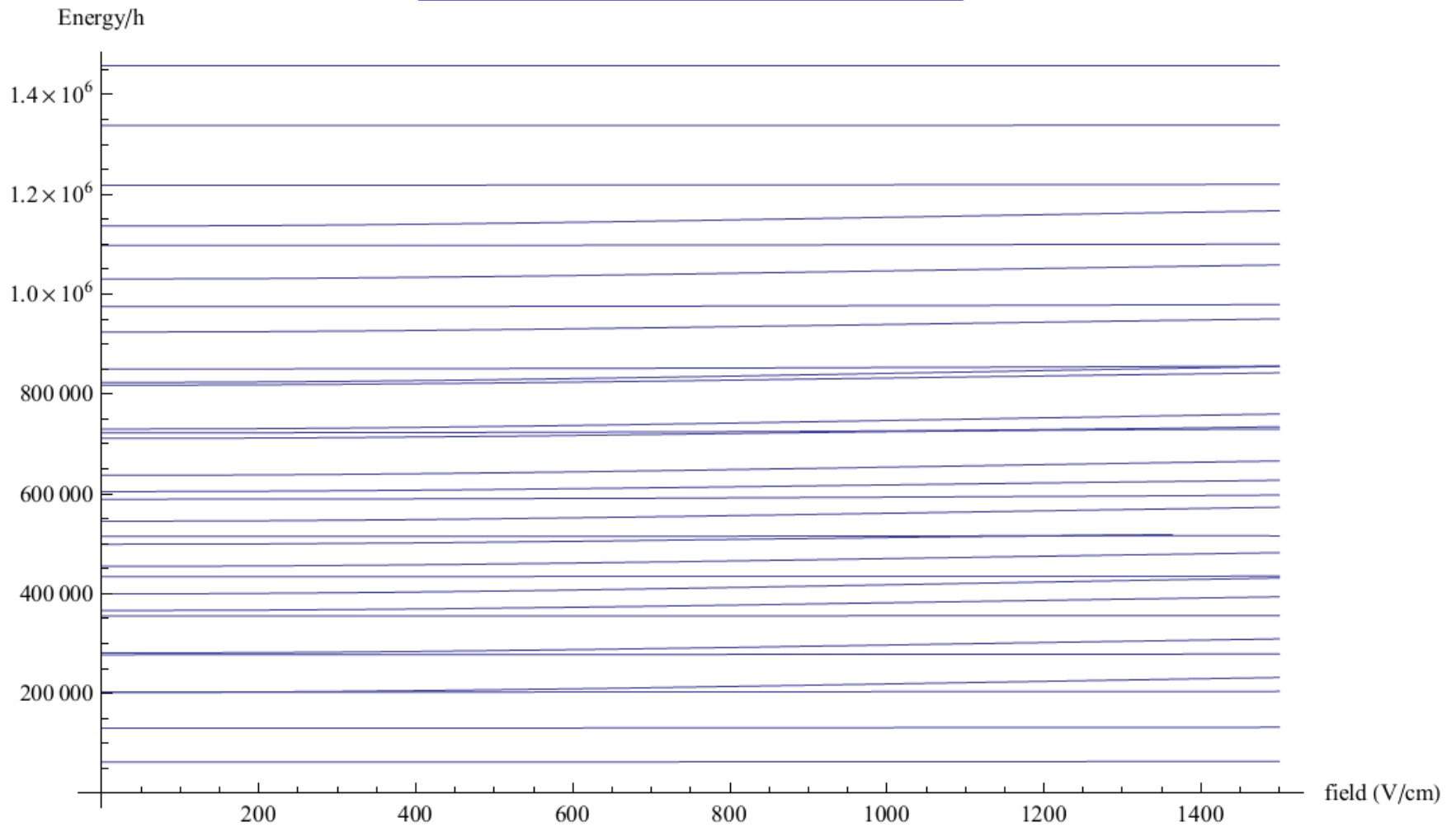
VH polarization 182G, excited state $M_F=4$.



VV polarization 182G, excited state $M=4$

87RbCs lowest rotational level

182G





Our current estimated phase-space density ~ 0.01 ?
A new method is necessary to get us to 1.

We really want something like this! (atom pairs)

VOLUME 90, NUMBER 11

PHYSICAL REVIEW LETTERS

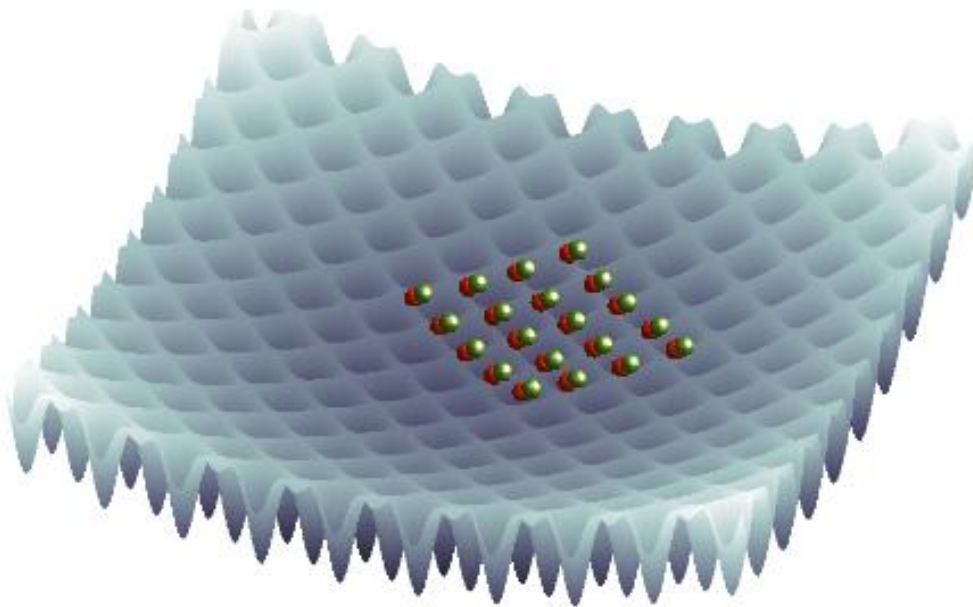
week ending
21 MARCH 2003

Creation of a Dipolar Superfluid in Optical Lattices

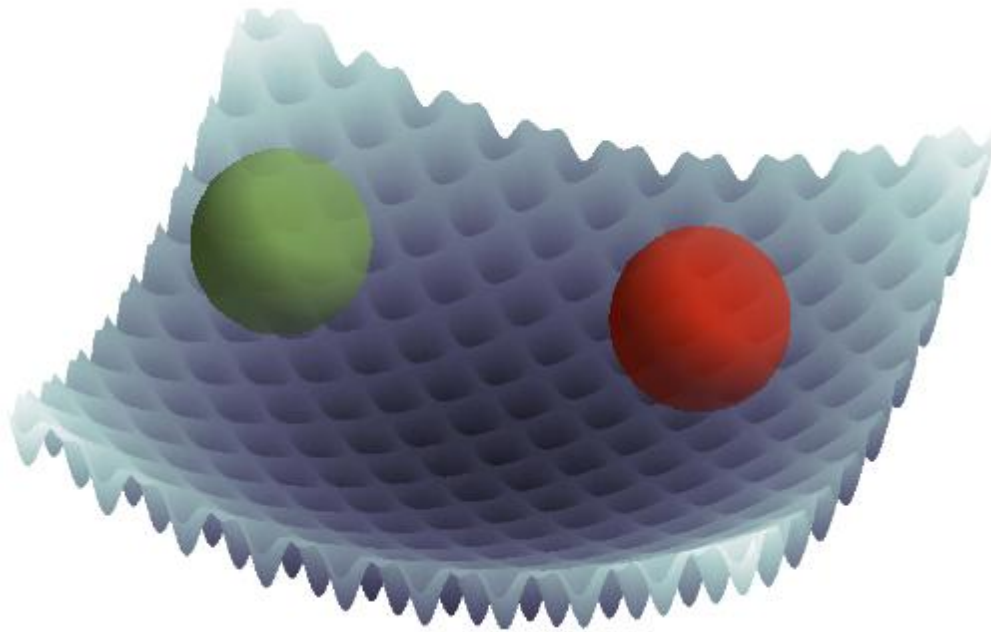
B. Damski,^{1,2} L. Santos,¹ E. Tiemann,³ M. Lewenstein,¹ S. Kotochigova,⁴ P. Julienne,⁴ and P. Zoller⁵

Noah Bray-Ali and Carl Williams
preliminary calculations

J. Freericks
numerical simulations starting



$$\hat{H} = -J \sum_{\langle ij \rangle} \hat{a}_i^\dagger \hat{a}_j - \sum_i \mu \hat{n}_i + \sum_i \frac{U}{2} \hat{n}_i (\hat{n}_i - 1)$$



$$U \propto \frac{4\pi\hbar^2 a}{m}$$

$$a_{CsCs} = 1700 a_0$$

$$a_{RbRb} = 100 a_0$$

$$a_{RbCs} = \text{tunable}$$

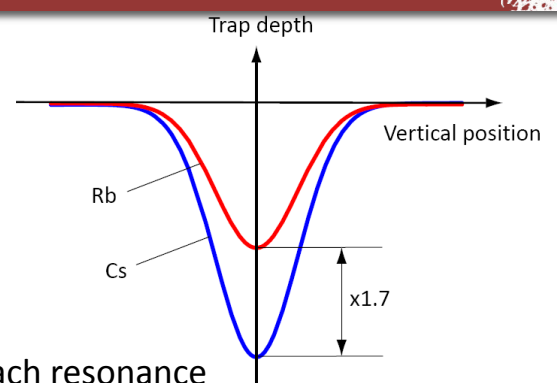
(0.2 G wide Feshbach resonance)

$$U_0^{RbRb}/J_{Rb} \ll 35 \quad U_0^{CsCs}/J_{Cs} \ll 35$$

2 superfluids

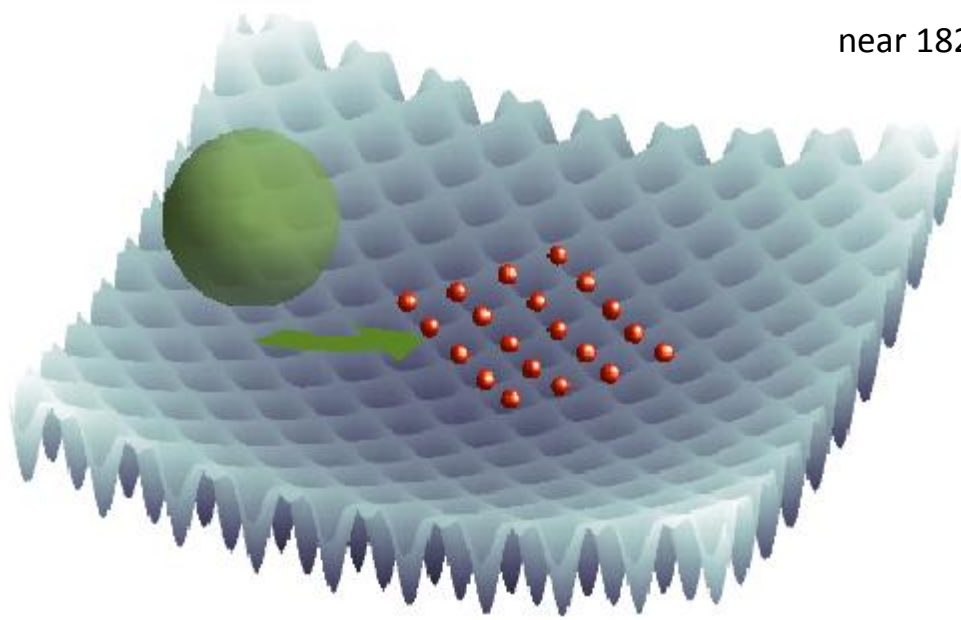
RbCs from a double Mott insulator

$$\hat{H} = -J \sum_{\langle ij \rangle} \hat{a}_i^\dagger \hat{a}_j - \sum_i \mu \hat{n}_i + \sum_i \frac{U}{2} \hat{n}_i (\hat{n}_i - 1)$$

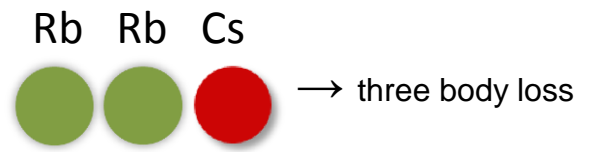


near 182G Feshbach resonance

- $U_{RbRb} \sim +k_B \times 15 \text{ nK} \quad (h \times 300 \text{ Hz})$
- $U_{CsCs} \sim +k_B \times 400 \text{ nK} \quad (h \times 8000 \text{ Hz})$
- $U_{RbCs} = \text{tunable, make negative?}$



Need to prevent site occupation by:



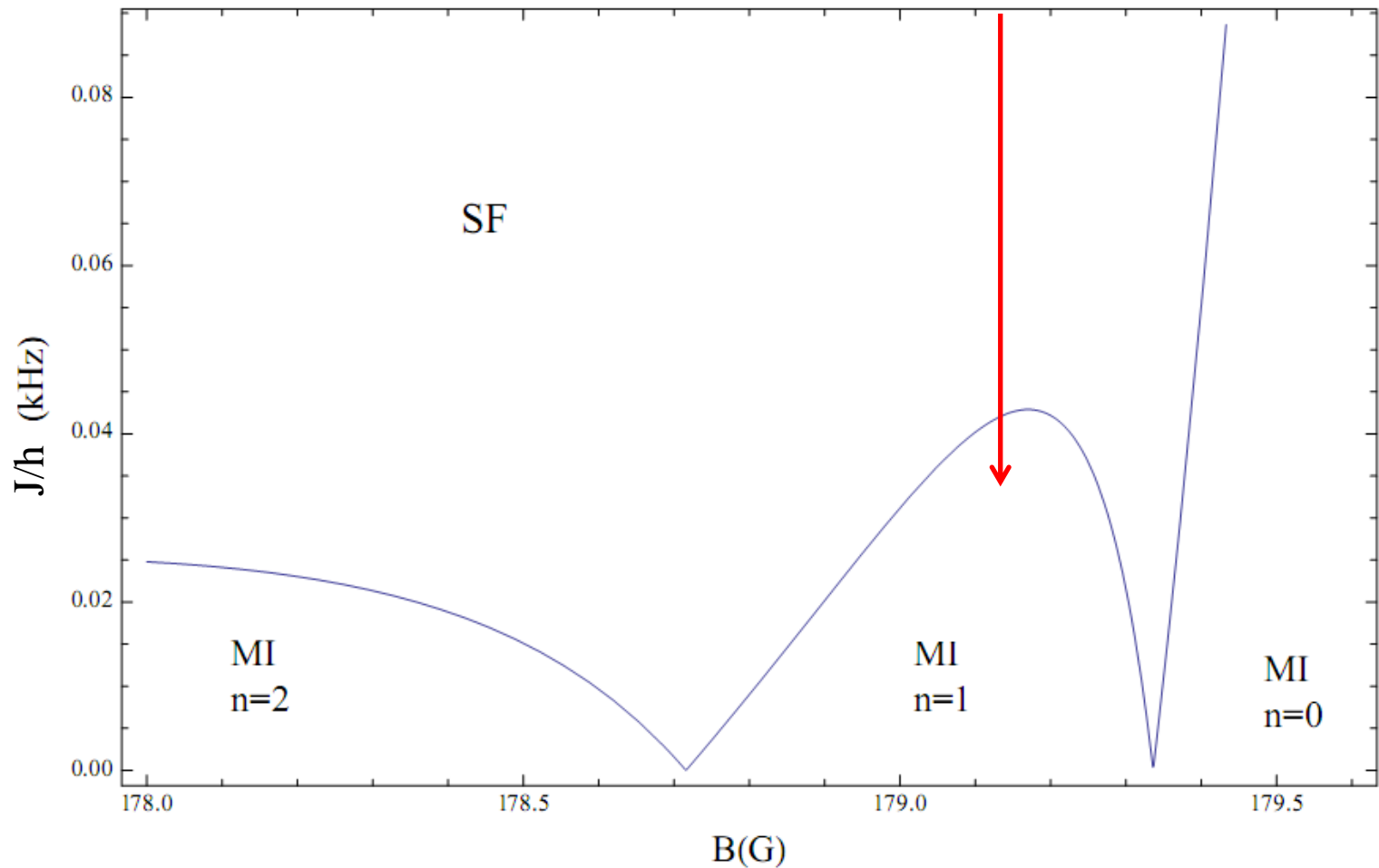
want $U_{RbRb} + U_{RbCs} > 0$

$$U_0^{RbRb} / J_{Rb} = 2500 \quad U_0^{CsCs} / J_{Cs} = 4.3$$

$$\frac{\Delta a_{RbCs}}{\Delta B} = \frac{a_{RbCs}^{background}}{\Delta} = \frac{649 a_0}{0.2 \text{ G}} = 3.3 \frac{a_0}{mG}$$

5–10 mG noise measured in lab → $\Delta a_{RbCs} = 17\text{-}33 a_0$

- Freeze out Cs
- Rb superfluid flows onto Cs
- Use onsite interactions U to prevent 3 particles per lattice site



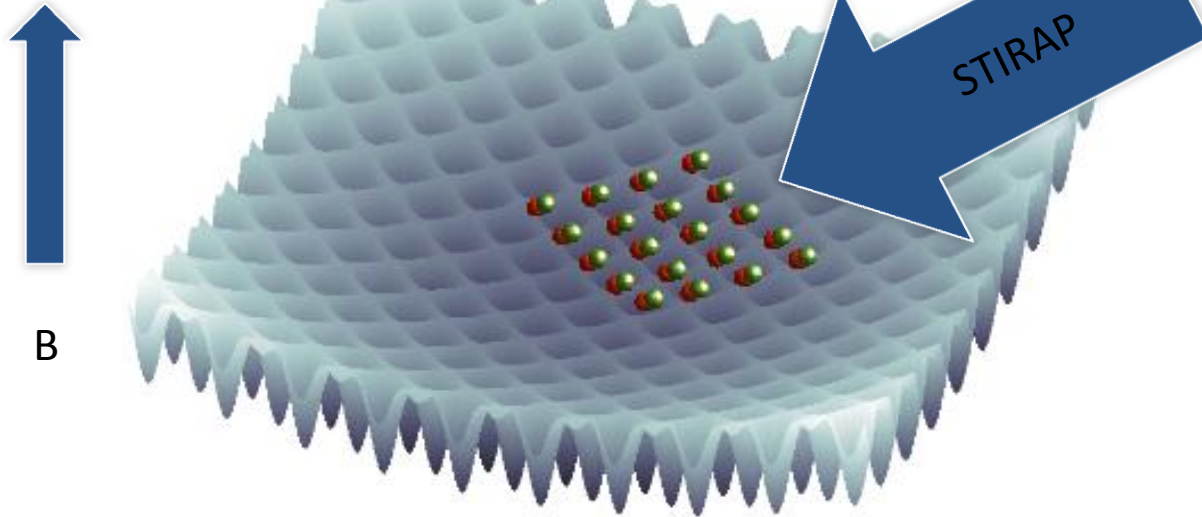
Mean-field phase diagram of Rb atoms in an optical lattice resonantly interacting with a Cs Mott insulator. J/h is the Rb tunneling rate. (Noah Bray-Ali, Carl Williams)

RbCs from a double Mott insulator

$$\hat{H} = -J \sum_{\langle ij \rangle} \hat{a}_i^\dagger \hat{a}_j - \sum_i \mu \hat{n}_i + \sum_i \frac{U}{2} \hat{n}_i (\hat{n}_i - 1) + \frac{U_{\text{NN}}}{2} \sum_{\vec{\ell}} \sum_{\langle\langle ij \rangle\rangle_{\vec{\ell}}} \frac{1}{|\vec{\ell}|^3} \hat{n}_i \hat{n}_j$$

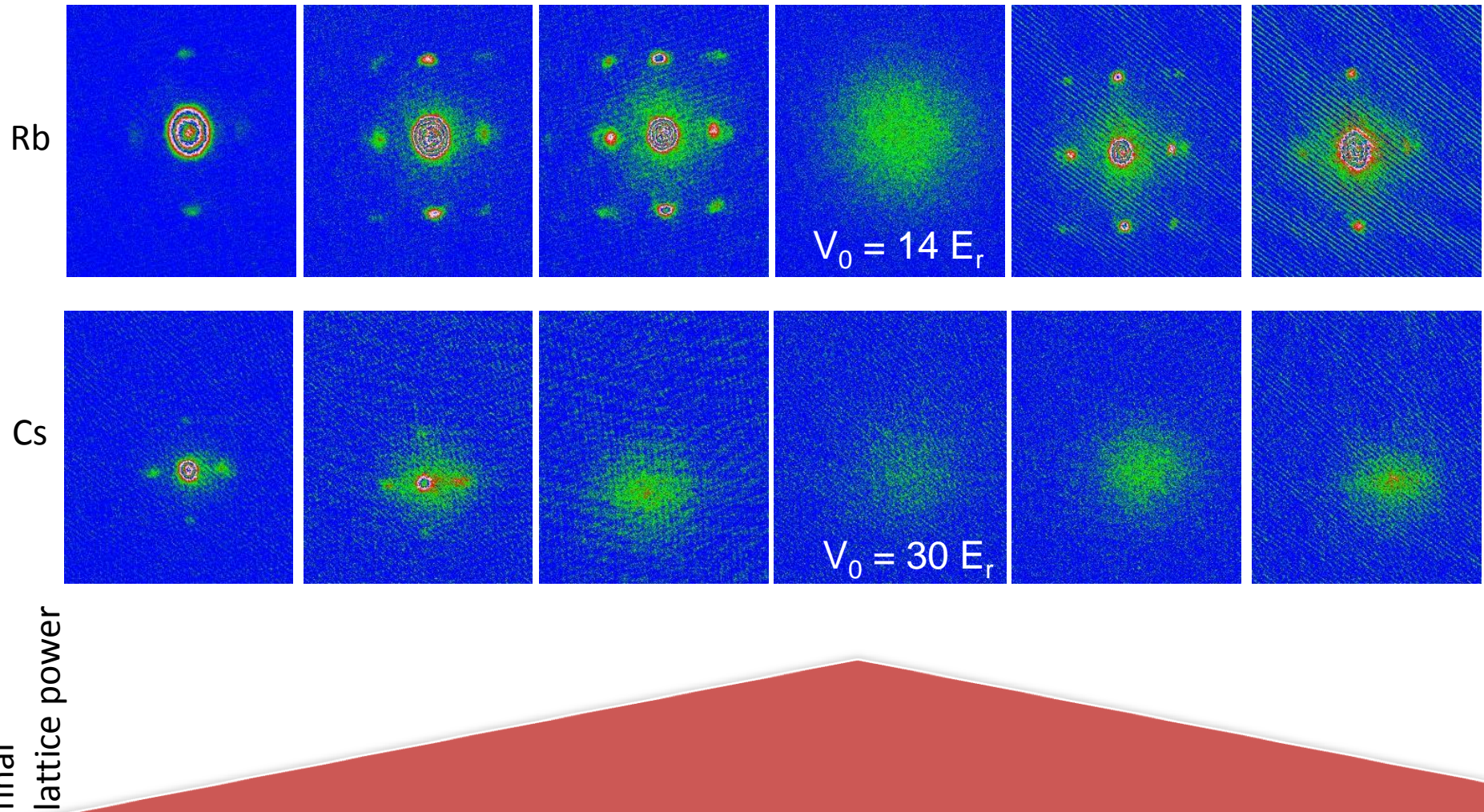
dipolar lattice physics !
(or melt for RbCs BEC)

make Feshbach molecules

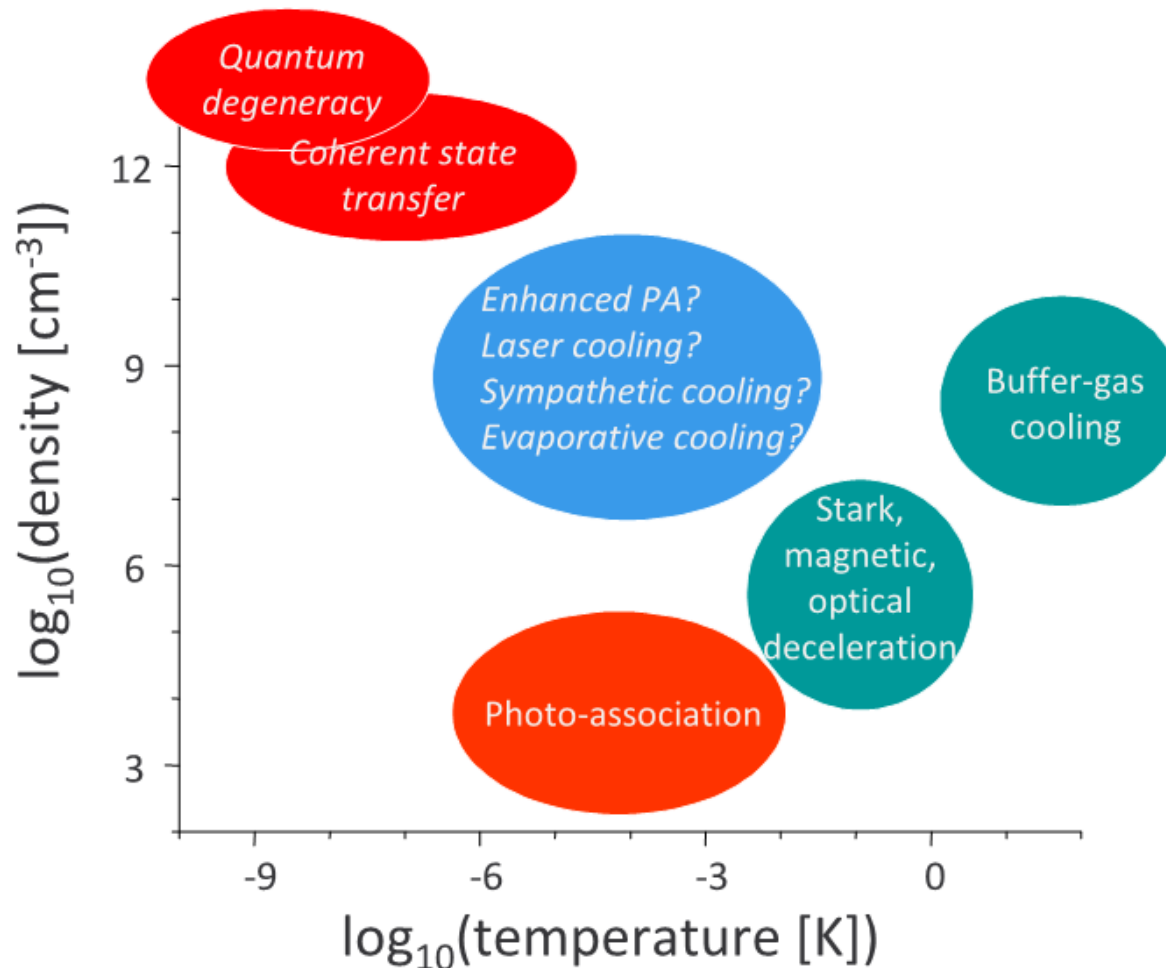


Now make Feshbach molecules, and do STIRAP (Much simpler version works for Rb₂, Cs₂)
Lower lattice adiabatically for molecular superfluid

Dual species SF-MI: same lattice but not yet overlapped



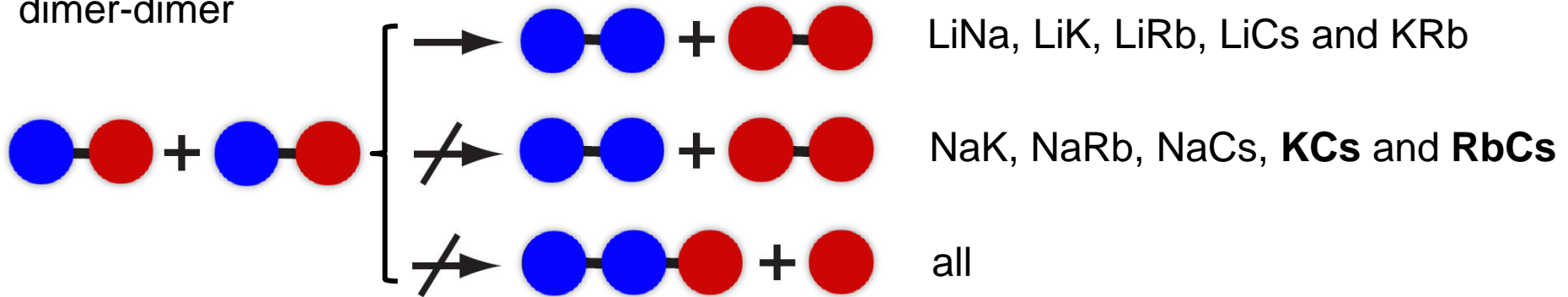
What if this is not enough for degeneracy?
(KRb currently at $1.5T_F$, PSD 0.1)



What if this is not enough for degeneracy?

3D Evaporative cooling of ground state RbCs with no electric field?

dimer-dimer



Piotr S. Zuchowski and Jeremy M. Hutson PRA **81**, 060703 (2010).

Three body loss?



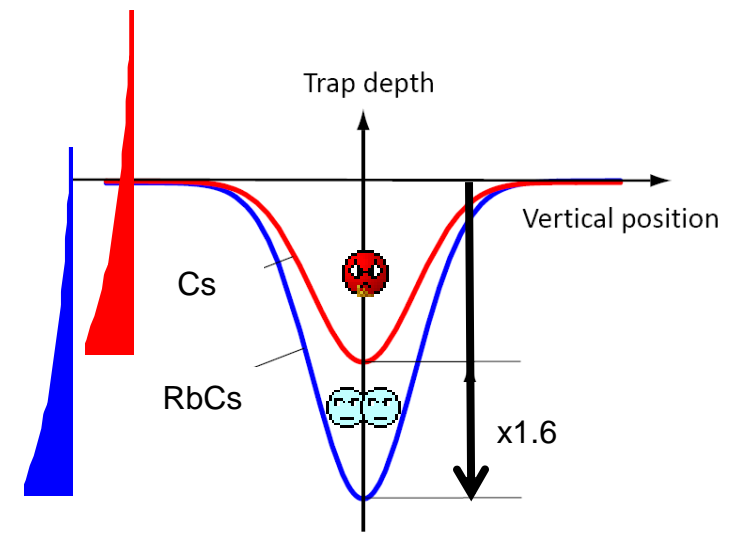
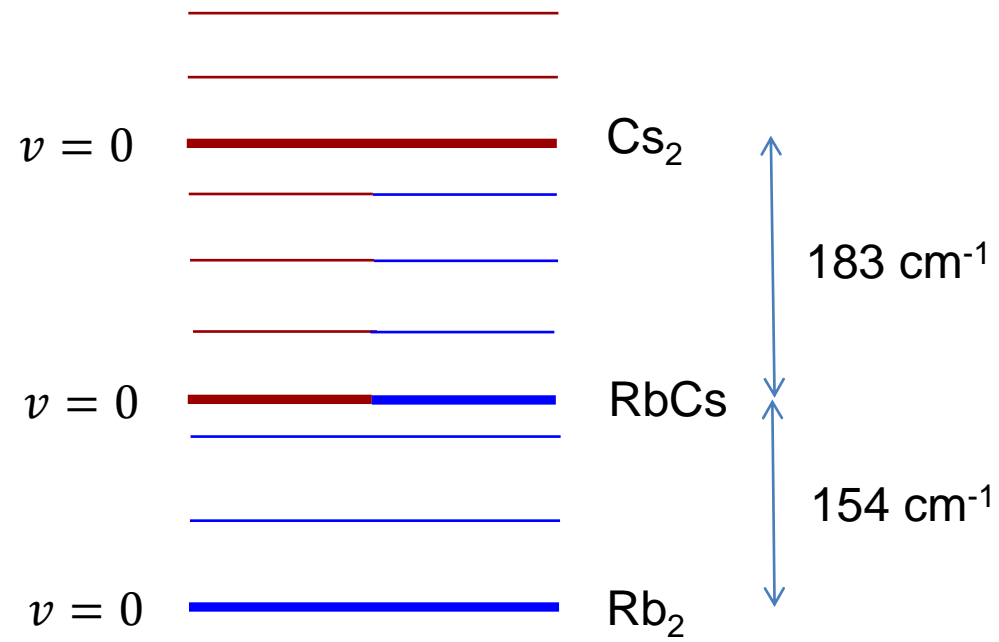
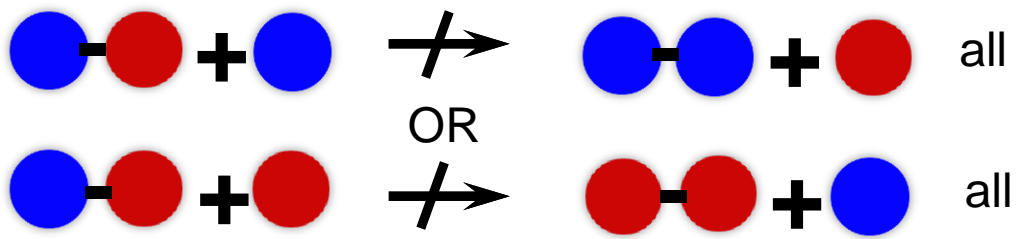
- Spin changing collisions driven by collision anisotropy
- Centrifugal barrier $\sim \mu^{-3/2} C_6^{-1/2}$

$C_6 = 140000a_0$ calculated for RbCs

Kotochigova NJP **12**, 073041 (2010).

Evaporative cooling of molecules:
 ground state molecules are precious.
 perhaps one can use atoms instead?

dimer-monomer collisions: one will be forbidden



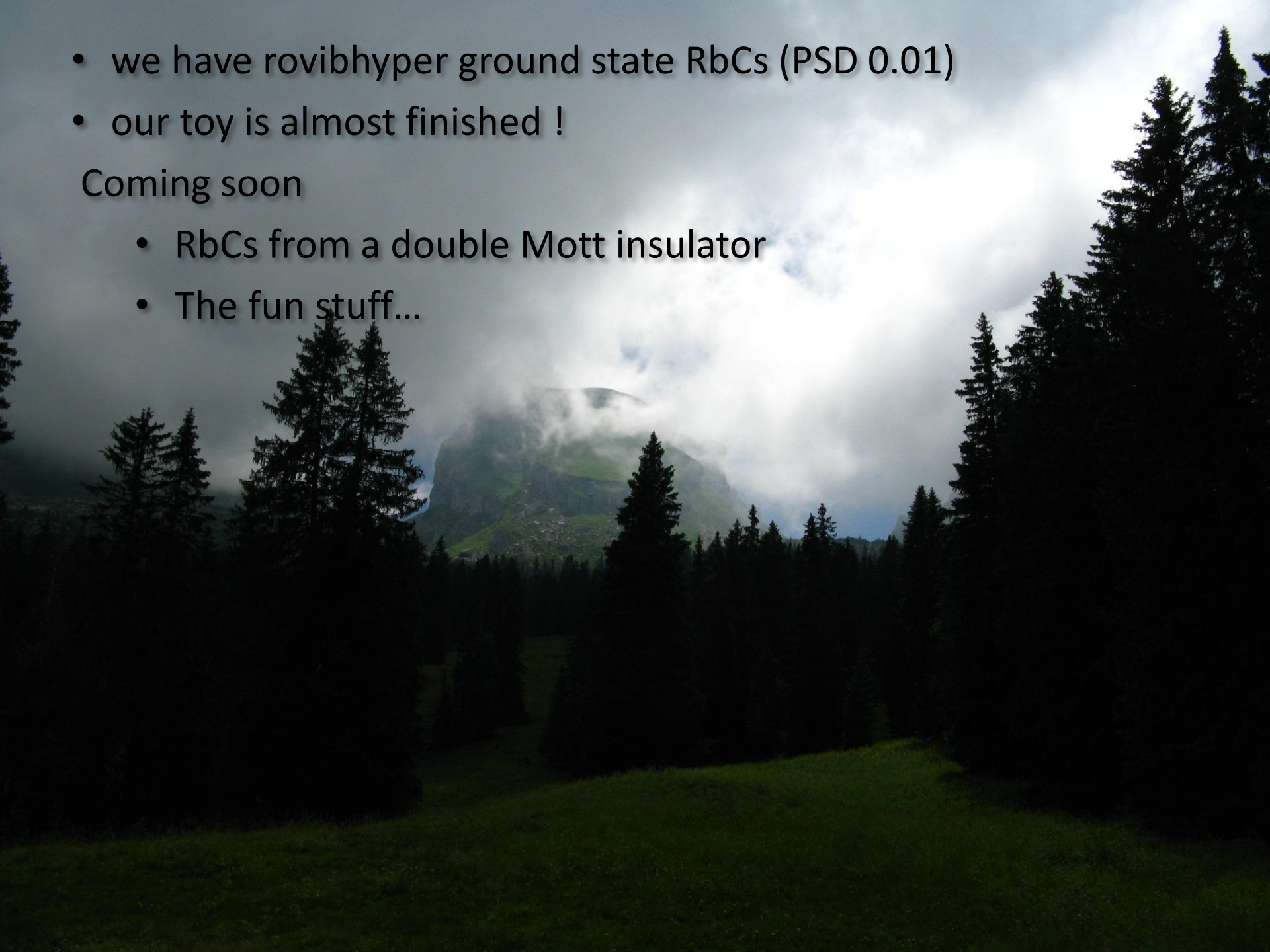
Use Cs atoms as a coolant

- requires high Cs-Cs, RbCs-Cs thermalization rates
- requires low Cs-Cs-Cs, RbCs-RbCs-RbCs, RbCs-RbCs-Cs, RbCs-Cs-Cs three body recombination rates (may need to goto 20G!)

- we have rovihyper ground state RbCs (PSD 0.01)
- our toy is almost finished !

Coming soon

- RbCs from a double Mott insulator
- The fun stuff...





€SF



Conference on Cold and Ultracold Molecules

November 18-23, 2012
University Center Obergurgl ("near" Innsbruck)

- Direct and indirect cooling techniques
- Controlled quantum chemistry
- Ultracold molecules for tests of quantum physics
- Molecular quantum gases
- Frontiers in molecular quantum control
- ...

Organizers: Guido Pupillo (chair), Francesca Ferlaino, Hanns-Christoph Nägerl



More to come...

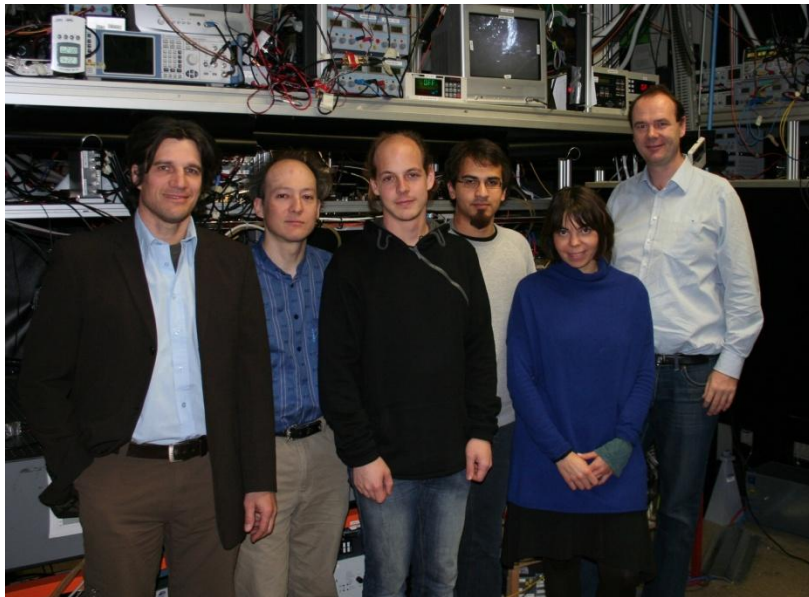
<http://www.esf.org/index.php?id=9144>



Foundations and
Applications of
Quantum Science

FWF





Eur. Phys. J. D (2011)
DOI: 10.1140/epjd/e2011-20015-6

THE EUROPEAN
PHYSICAL JOURNAL D

Regular Article

Production of a dual-species Bose-Einstein condensate of Rb and Cs atoms

A.D. Lercher¹, T. Takekoshi^{1,*}, M. Debatin¹, B. Schuster¹, R. Rameshan¹, F. Ferlaino¹, R. Grimm^{1,2}, and H.-C. Nägerl¹

PCCP

Dynamic Article Links 

Cite this: DOI: 10.1039/c1cp21769k

www.rsc.org/pccp

PAPER

Molecular spectroscopy for ground-state transfer of ultracold RbCs molecules

Markus Debatin,^a Tetsu Takekoshi,^{ab} Raffael Rameshan,^a Lukas Reichsöllner,^a Francesca Ferlaino,^{*a} Rudolf Grimm,^{ab} Romain Vexiau,^c Nadia Bouloufa,^c Olivier Dulieu^c and Hanns-Christoph Nägerl^a

Received 1st June 2011, Accepted 24th July 2011

DOI: 10.1039/c1cp21769k

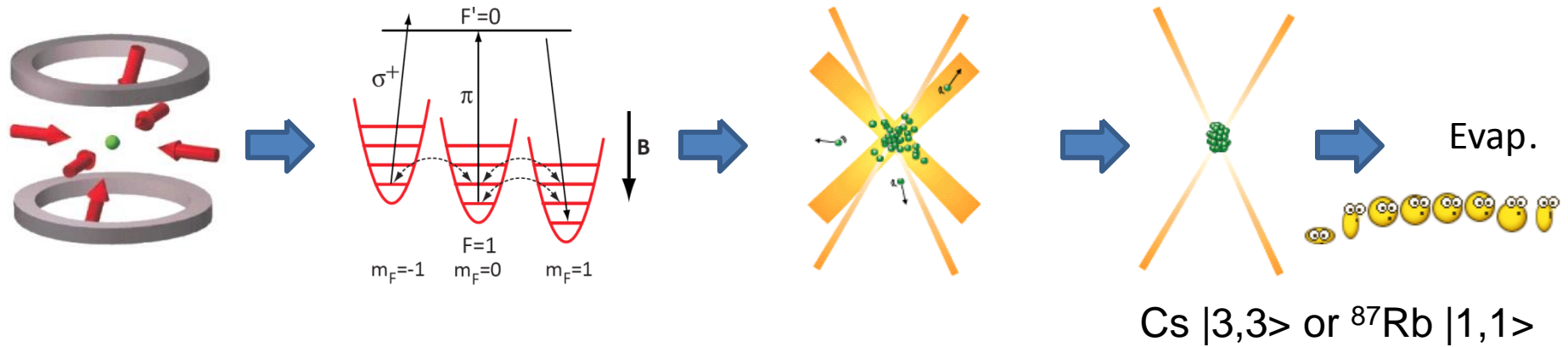
PHYSICAL REVIEW A **85**, 032506 (2012)

Towards the production of ultracold ground-state RbCs molecules: Feshbach resonances, weakly bound states, and the coupled-channel model

Tetsu Takekoshi,^{1,2} Markus Debatin,¹ Raffael Rameshan,¹ Francesca Ferlaino,¹ Rudolf Grimm,^{1,2} Hanns-Christoph Nägerl,¹ C. Ruth Le Sueur,³ Jeremy M. Hutson,³ Paul S. Julienne,⁴ Svetlana Kotochigova,⁵ and Eberhard Tiemann⁶

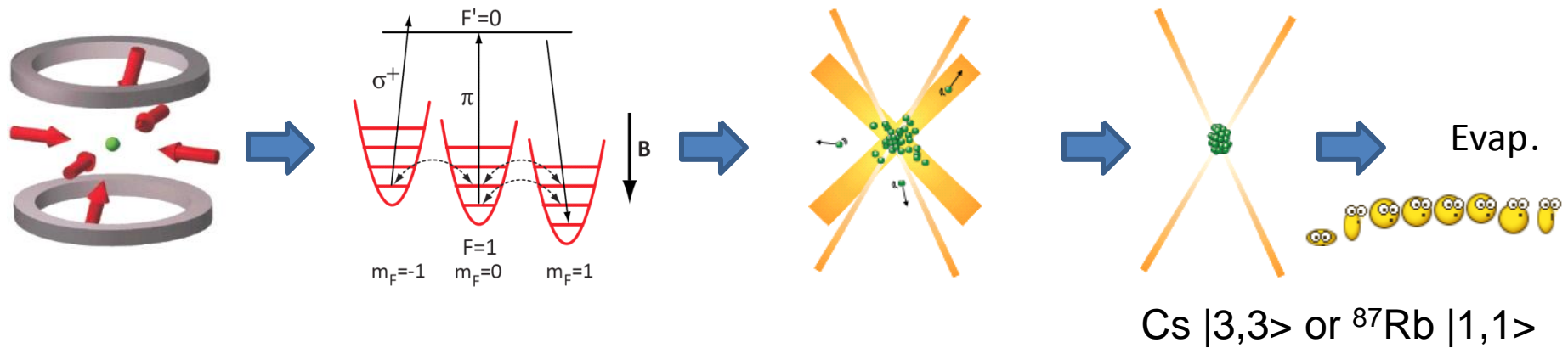
Cs BEC requires optical traps.

Optical BEC of Rb makes dual species apparatus simpler.



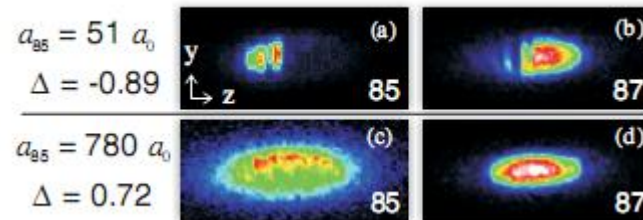
Cs BEC requires optical traps.

Optical BEC of Rb makes dual species apparatus simpler.



Our dream:

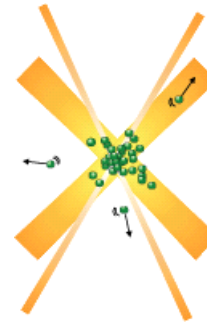
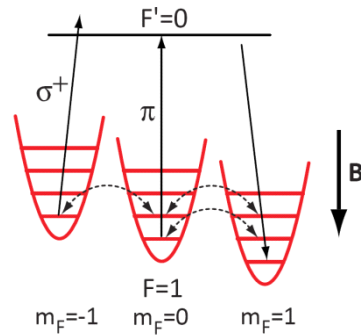
like ^{87}Rb ^{85}Rb mixture



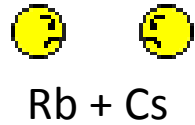
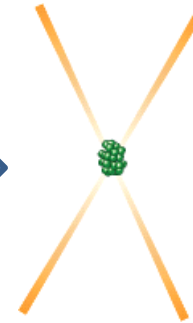
S.B. Papp, J.M. Pino, C.E. Wieman PRL **101** 040402 (2008).

Making Feshbach molecules requires a high phase space density mixture.

mixture problems bad, but not insurmountable



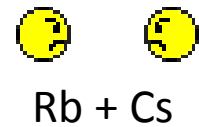
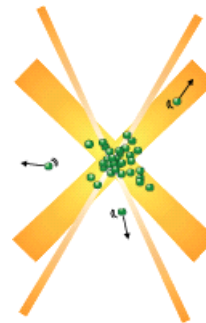
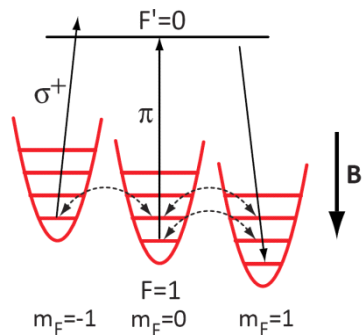
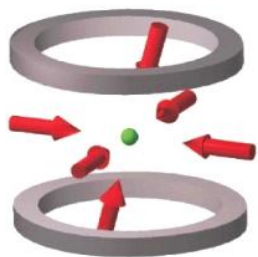
really bad mixture problems



Starting point: all-optical Cs BEC

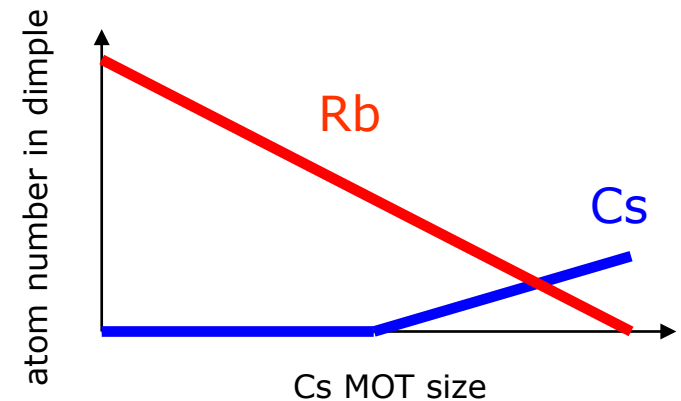
Making Feshbach molecules requires a high phase space density mixture

mixture problems bad, but not insurmountable



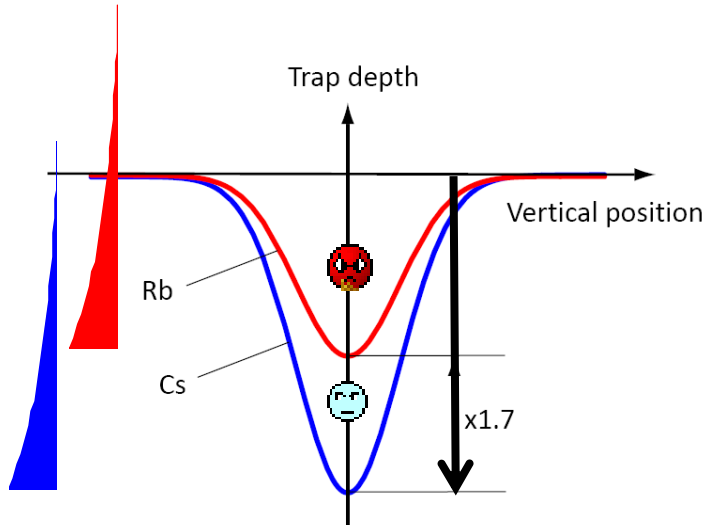
Evap.

really bad mixture problems



Important for Rb-Cs mixtures:

Traps are deeper for Rb than for Cs.
(evaporative heat load mostly on Rb)



high Rb/Cs thermalization

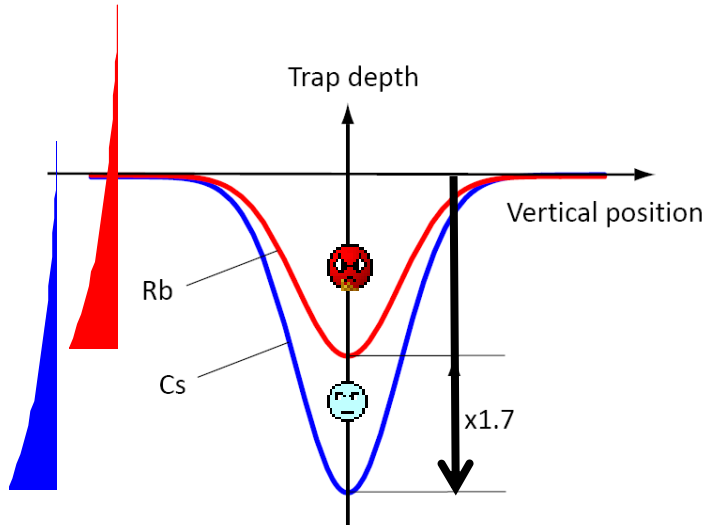
Bad for simultaneous evaporation
Good for Cs cooling! (efficient)

Bad luck #1



Important for Rb-Cs mixtures:

Traps are deeper for Rb than for Cs.
(evaporative heat load mostly on Rb)



high Rb/Cs thermalization

Bad for simultaneous evaporation
Good for Cs cooling! (efficient)

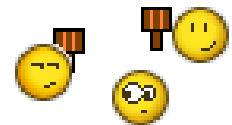
Large interspecies background scattering length
($a_{\text{RbCs}} \sim 649a_0$ from coupled channel model)

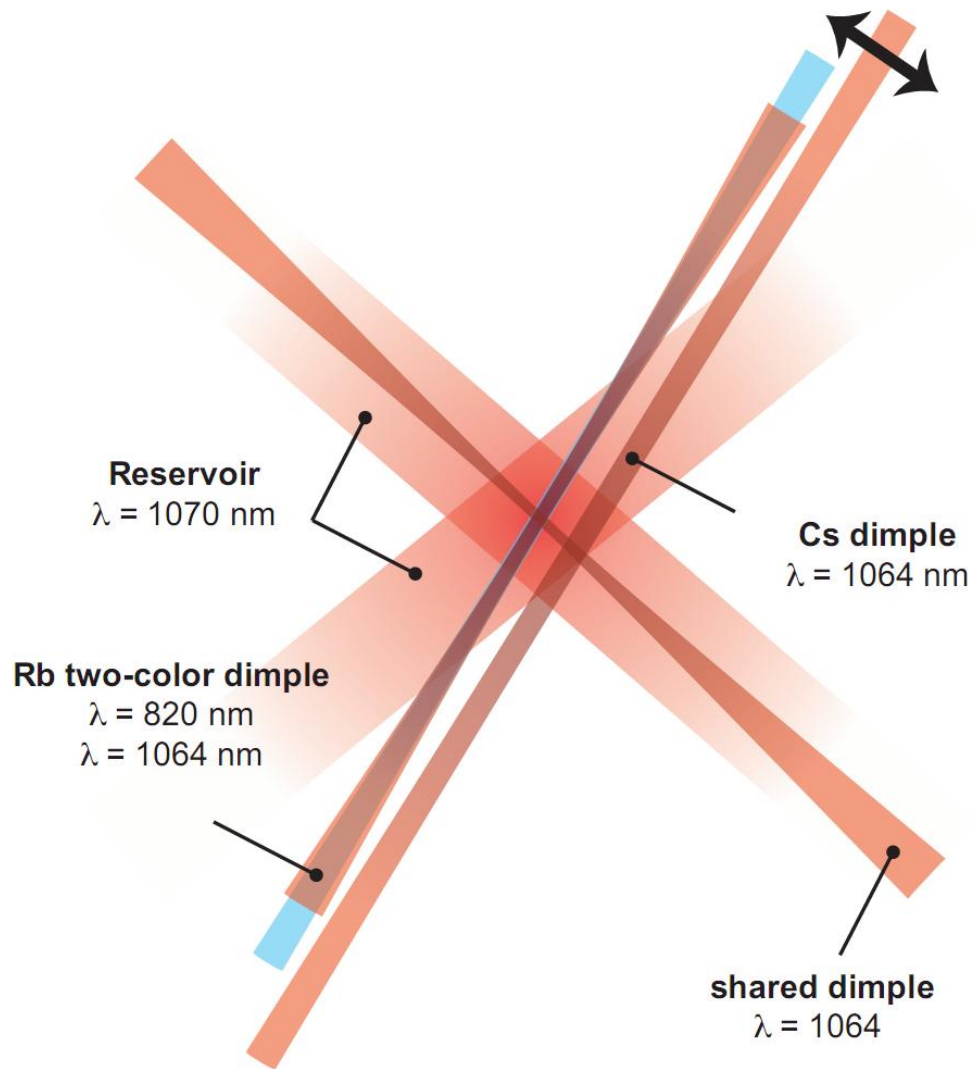


Large three-body recombination rates
For example $K_{\text{RbRbCs}} \sim |a_{\text{RbCs}}|^4 \approx 10^{-24} \text{ cm}^6 \text{ s}^{-1}$
(measured)

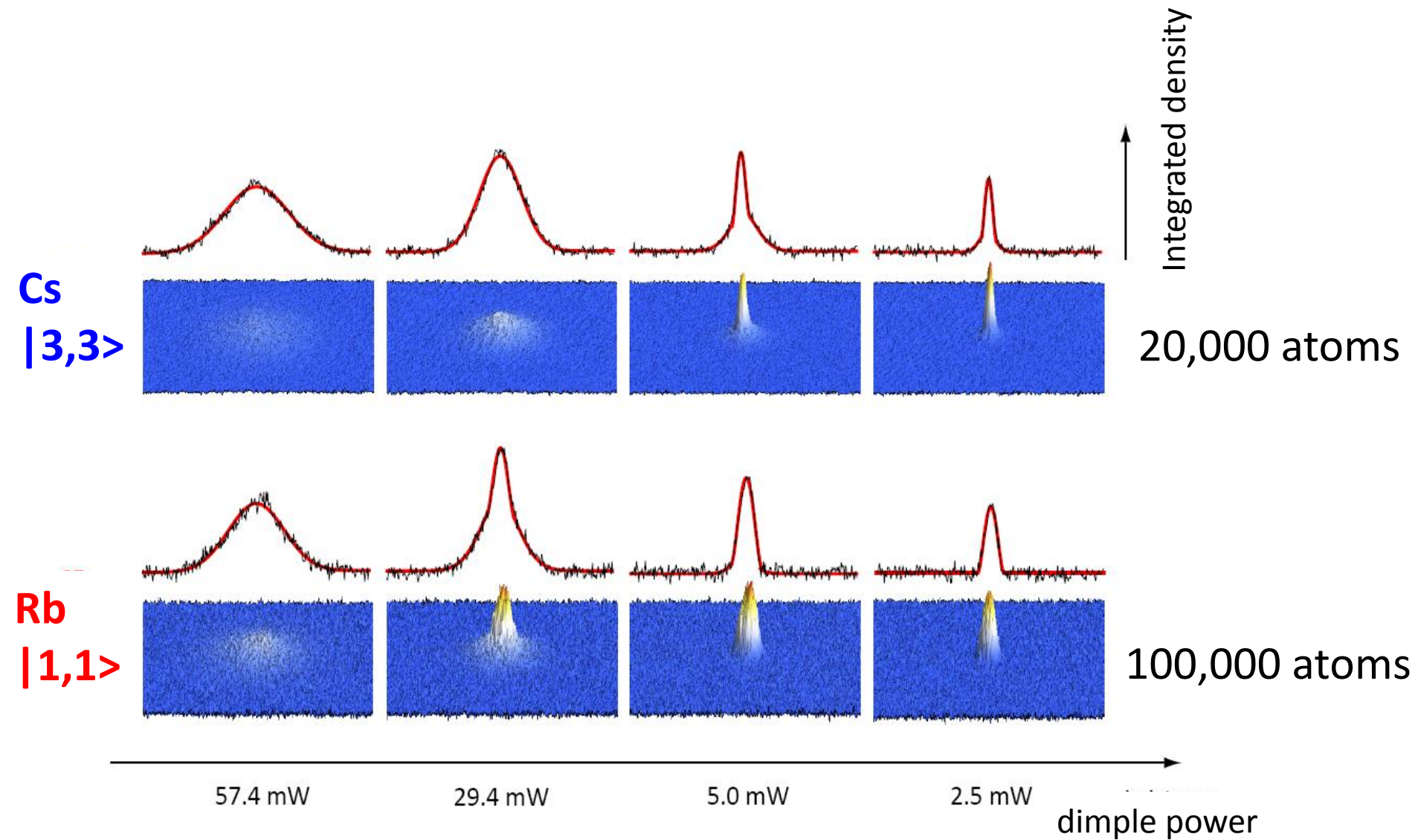


Ratio of „good“ to „bad“ collisions
 $= 4\pi a_{\text{RbCs}}^2 / (K_{\text{RbRbCs}} n_{\text{Rb}})$

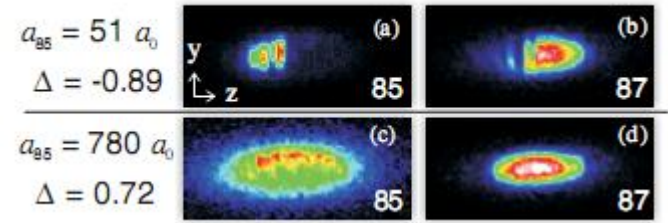
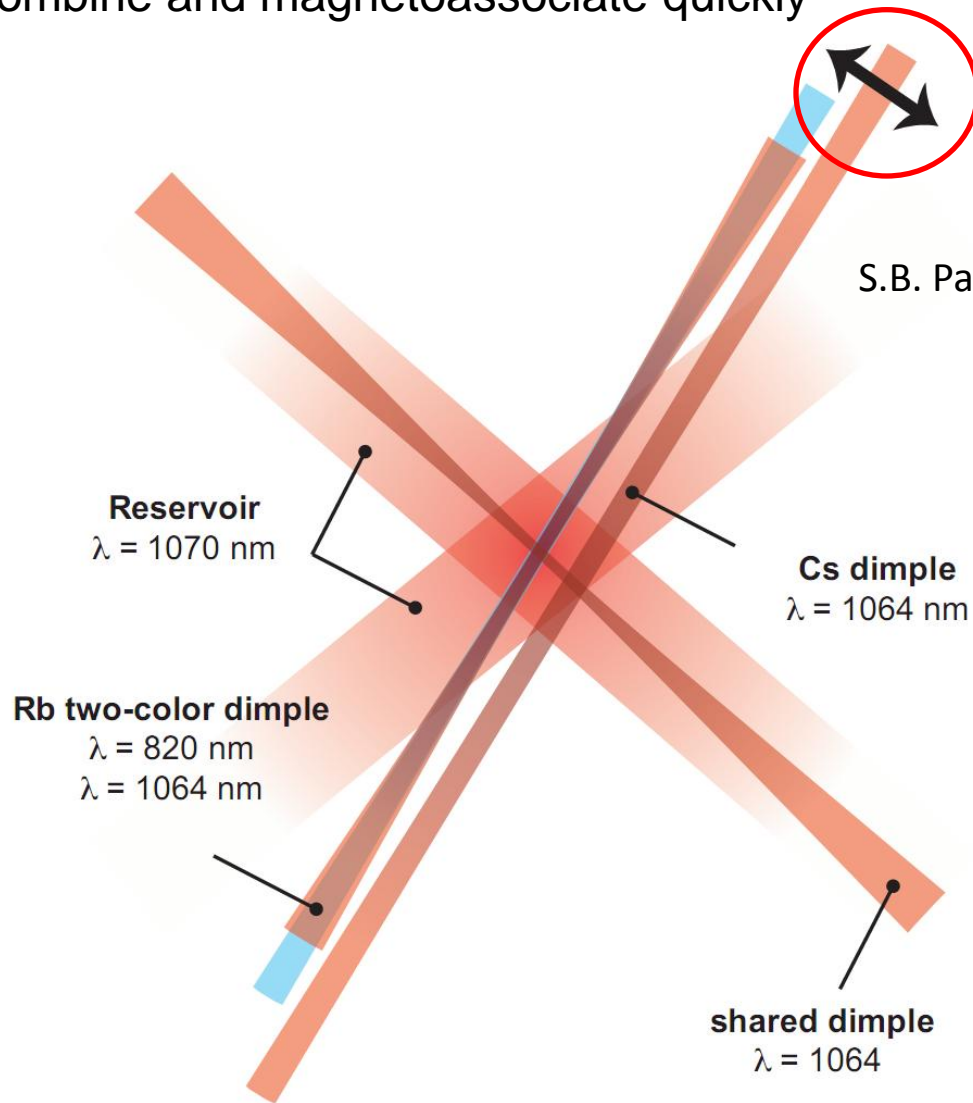




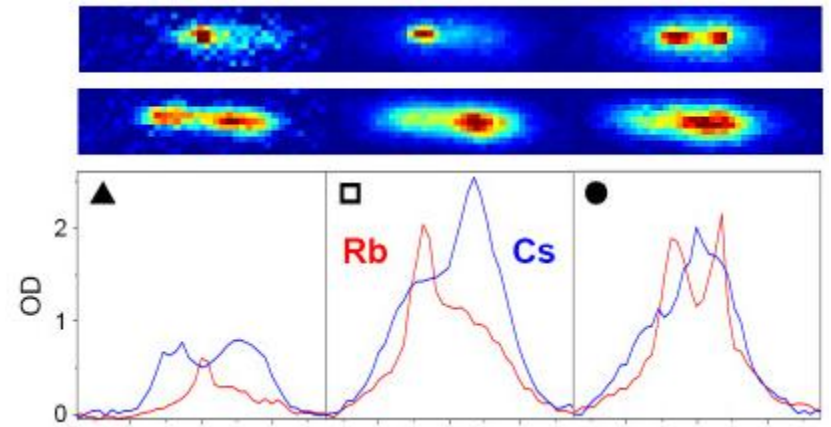
- Raman cooling
- reservoir (spin filter)
- separate dimples



Combine and magnetoassociate quickly



S.B. Papp, J.M. Pino, C.E. Wieman PRL **101** 040402 (2008).



Durham group PRA **84** 011603 (2011).

$$\frac{a_{RbRb} - a_{CsCs}}{a_{RbCs}^2} - 1 > 0 \quad \text{miscible}$$



Rb superfluid-Mott insulator transition

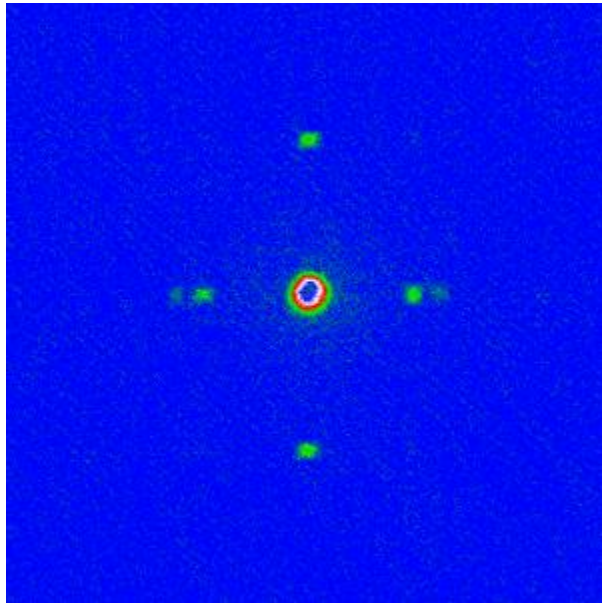
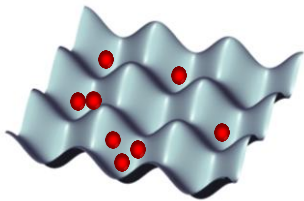


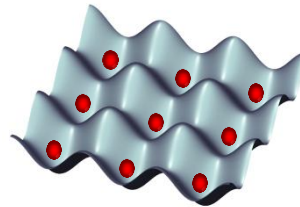
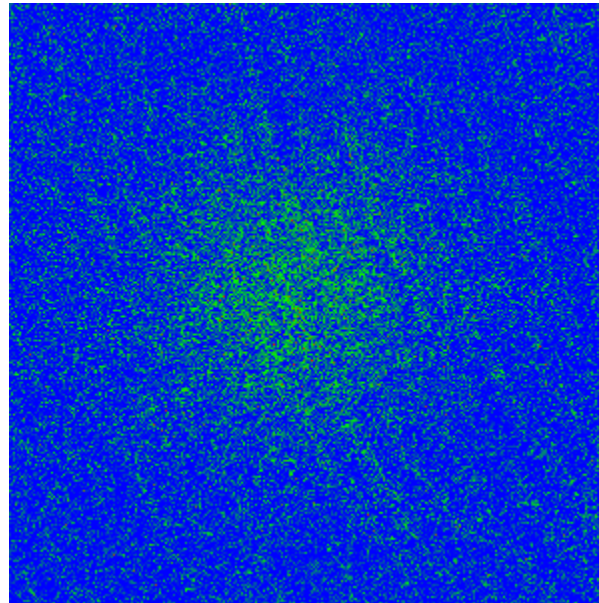
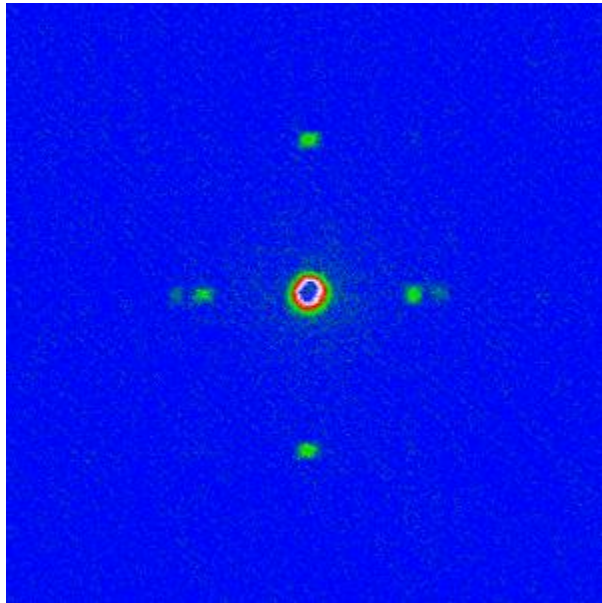
Image after expansion – matter wave interference



Superfluid state $J \gg U$

- delocalised
- poissonian distribution
- phase coherence
- interference pattern

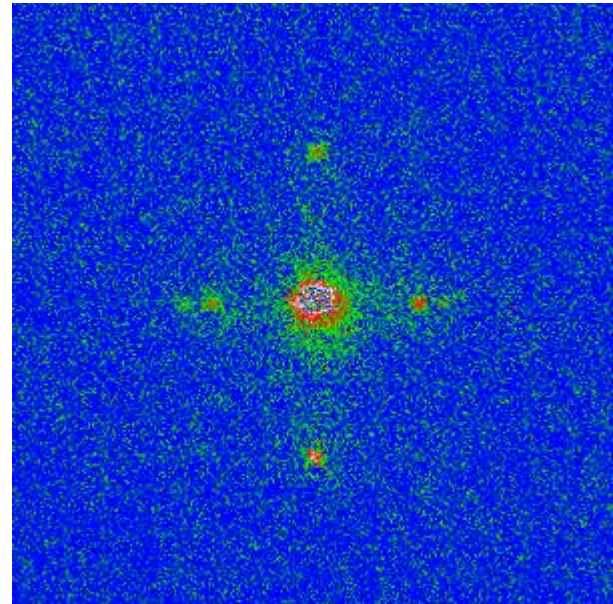
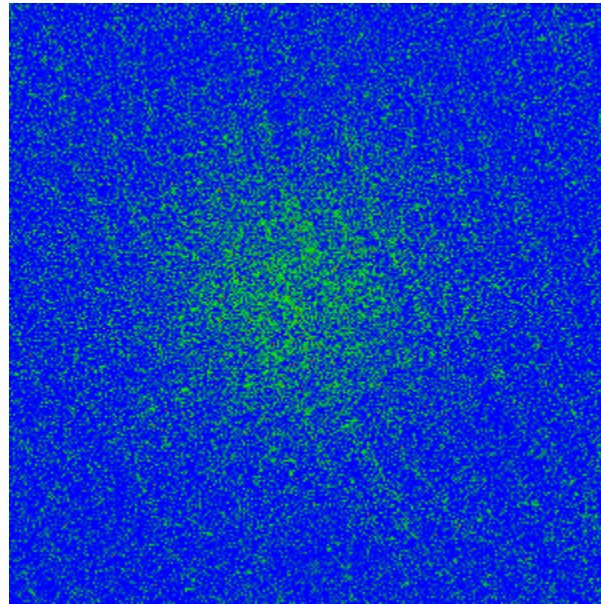
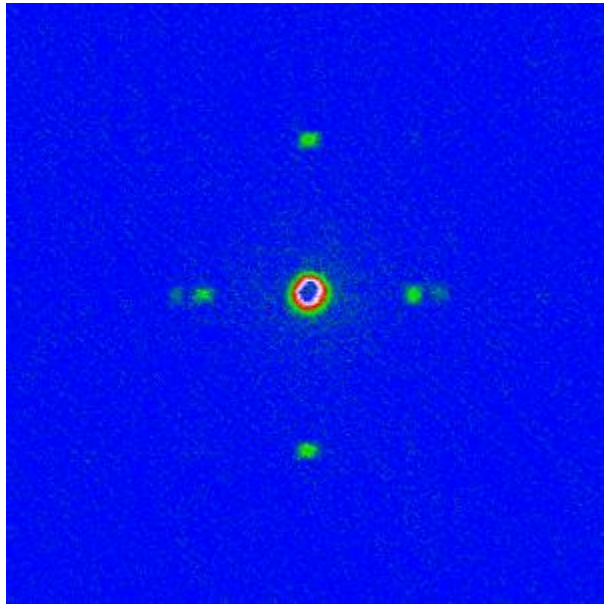
Rb superfluid-Mott insulator transition



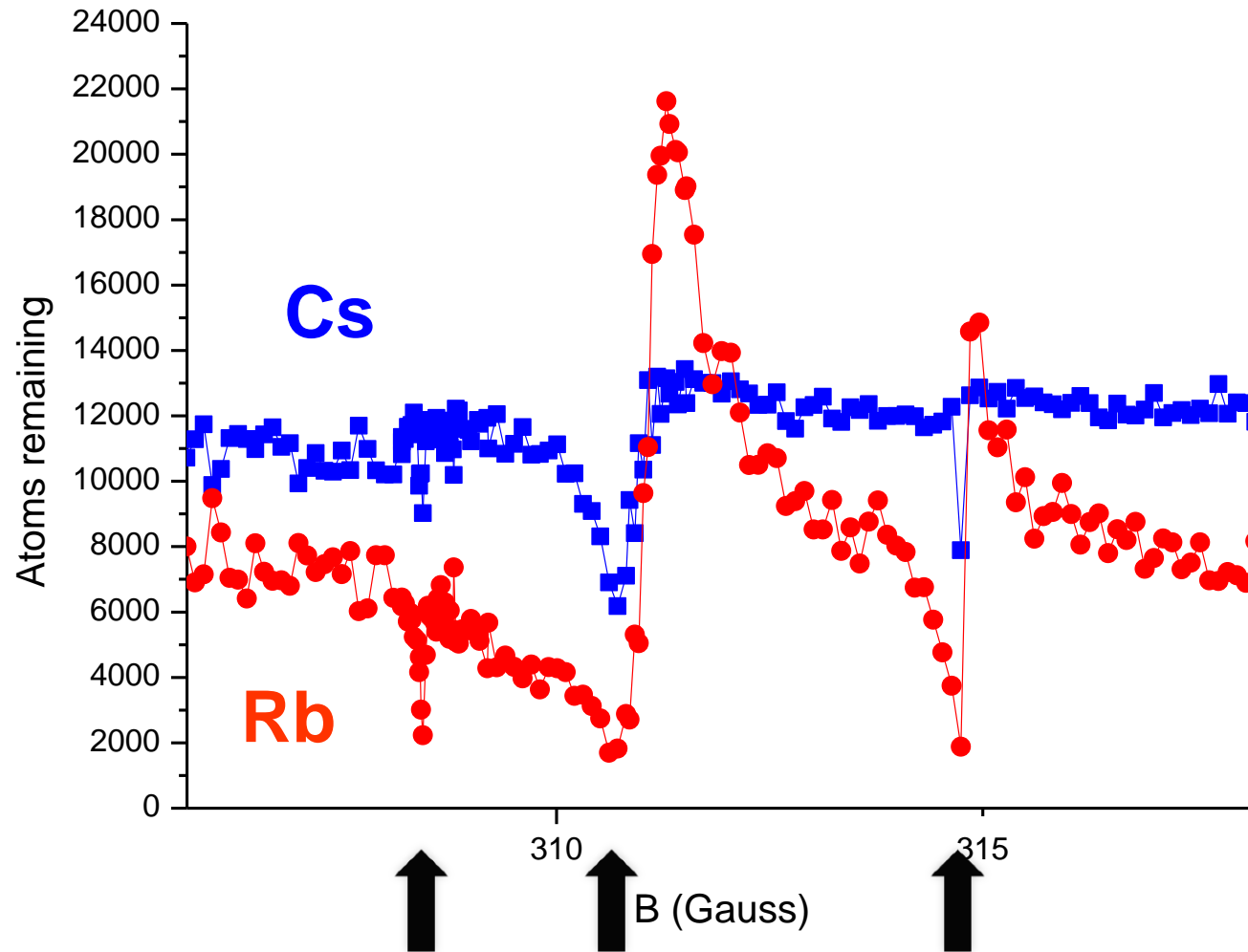
Mott insulator state $J \ll U$

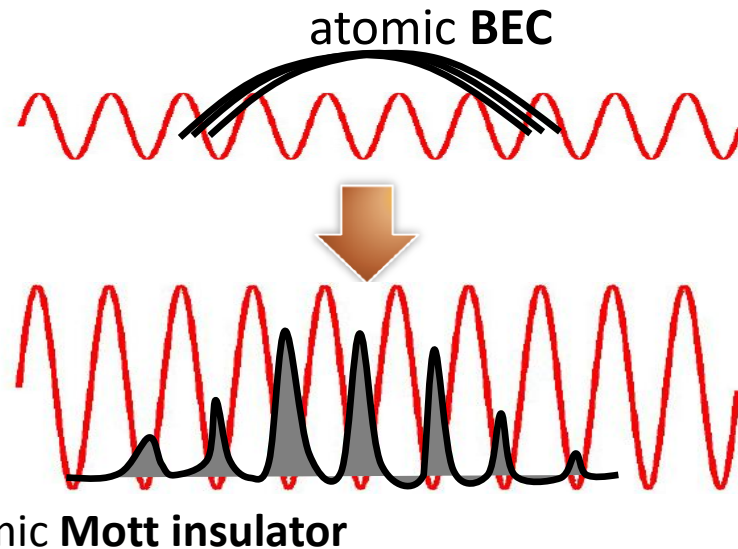
- localized atoms
- no phase coherence
- no interference pattern
- fixed atom number per site

Rb superfluid-Mott insulator transition



Data for model – Feshbach resonances

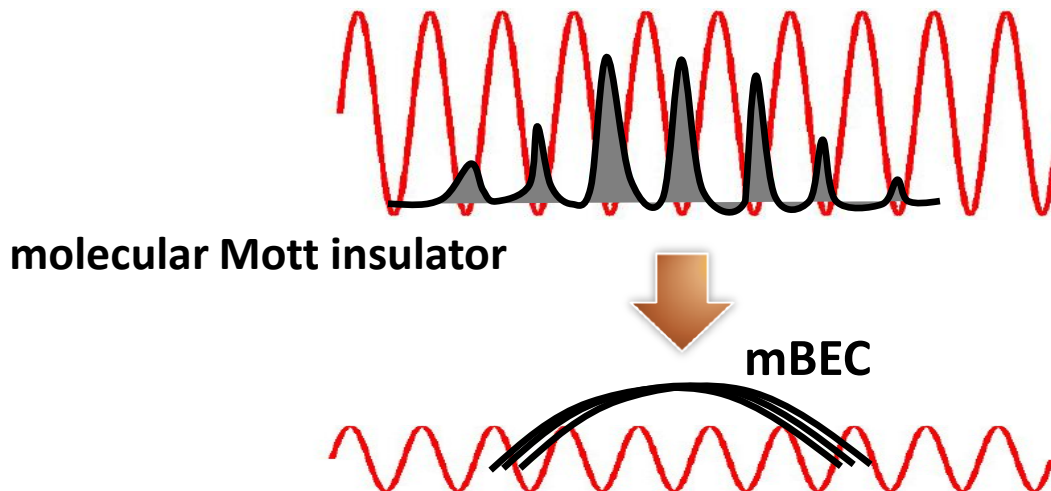




Increasing the lattice depth
Phase transition
from **the superfluid BEC**
to a localized
Mott-insulator state

Proposal: P. Zoller et al., 1998

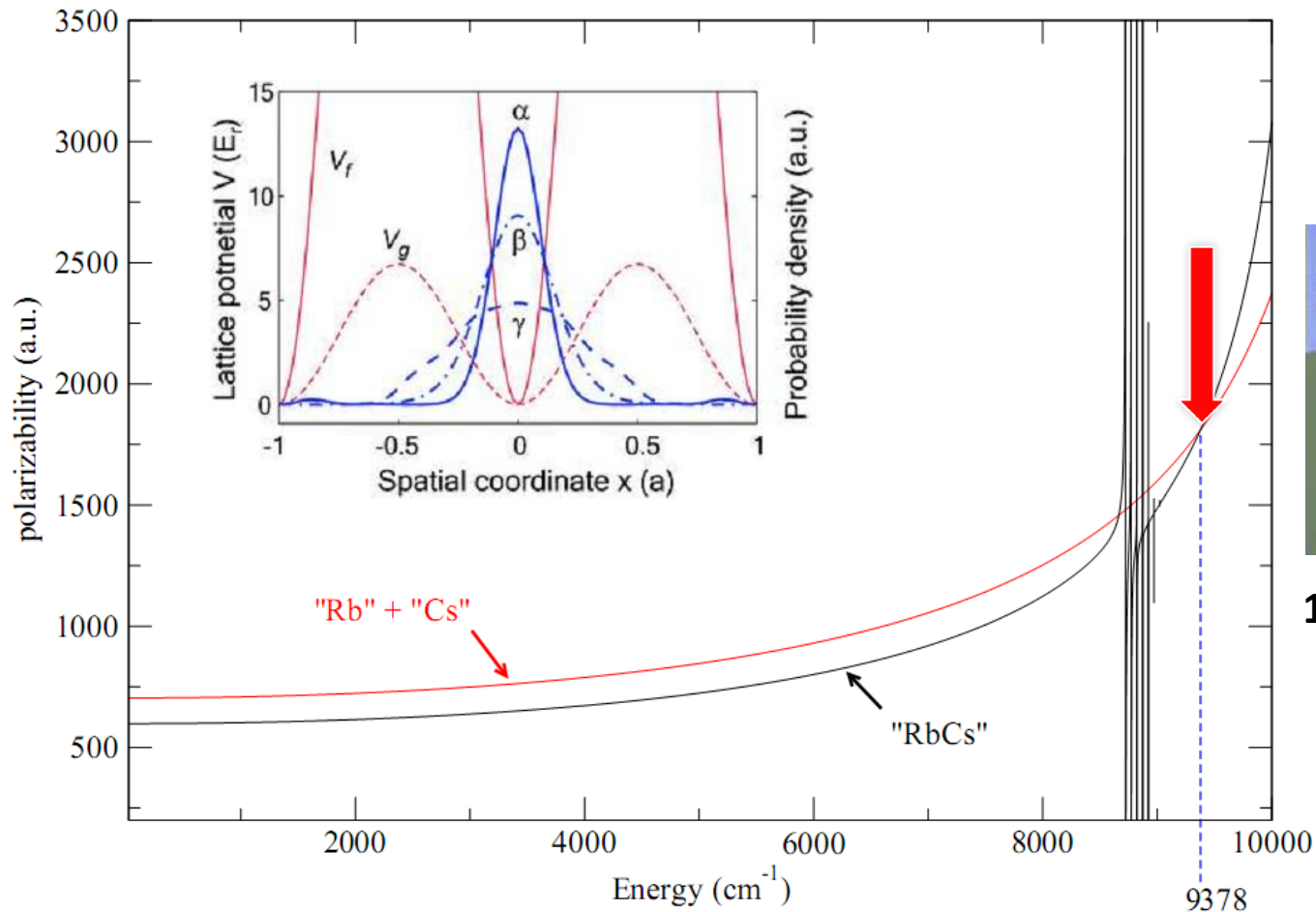
This should also work the other way round!!!



Reduction of lattice depth:
Phase transition
from a localized
Mott-insulator state
of molecules
to a **molecular**
BEC („mBEC“)

Proposal: P. Zoller et al., 2002

Romain Vexieu, Nadia Bouloufa, Oliver Dulieu



Luck!!



$$1/1064\text{nm} = 9398 \text{ cm}^{-1}$$

magic wavelength

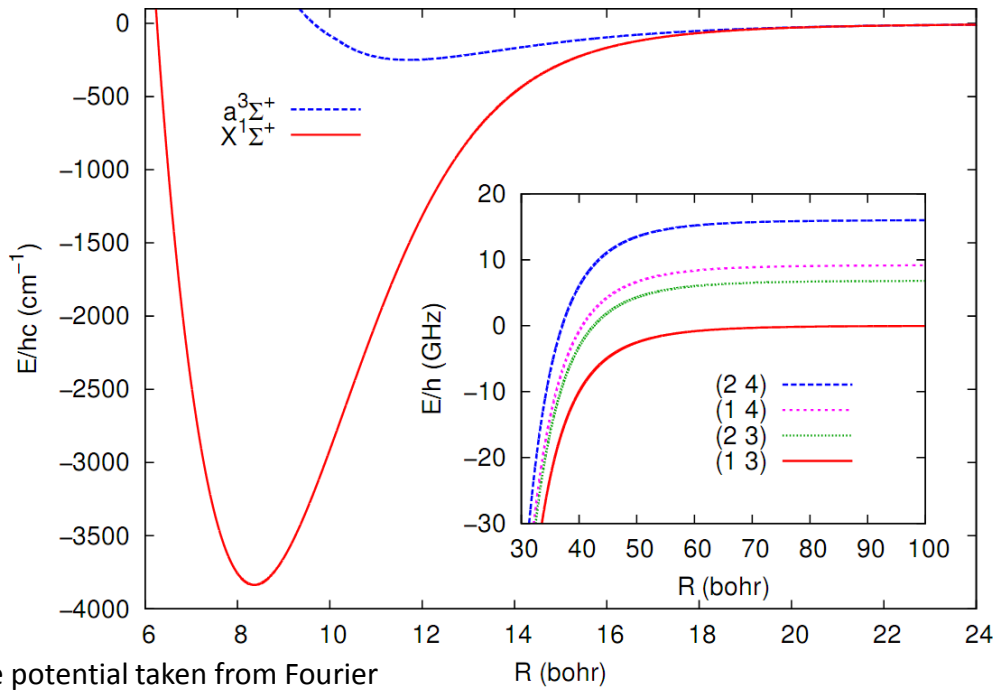
Coupled channel model

$$V_S^{SR}(R) = A_S + B_S[R_S^{SR}/a_0]^N$$

$$N = -7.31737405186660705$$

$$V_S(R) = \sum_{i=0}^n a_i \xi^i(R) \quad \xi(R) = \frac{R - R_m}{R + bR_m}$$

$$V_S^{LR}(R) = -C_6/R^6 - C_8/R^8 - C_{10}/R^{10} \pm V_{\text{exch}}(R)$$



C. Ruth LeSueur, Jeremy M. Hutson (Durham)
 Paul S. Julienne (JQI, NIST, UMD)
 Svetlana Kotochigova (Temple)
 Eberhard Tiemann (Hannover)

Initial medium range potential taken from Fourier transform spectroscopy study (Riga)
 O. Docenko *et al.*, PRA **83**, 052519 (2011).
 level energies accurate to ~ 1GHz x h

spin-spin and 2nd order spin-orbit (avoided crossing strengths)

TABLE III: Parameters of the fitted potential. [12 Aug 2011]

	fitted value	95% confidence limit	sensitivity limit
Short range			
$B_0^{SR} (E_h)$	147.675	0.49177	0.00007
$B_1^{SR} (E_h)$	428.9211	0.86538	0.00003
A_{2SO}^{long}	0.0001350	0.0000006	0.0000136
Long range			
$C_6 (E_h a_0^6)$	5694.8300	0.3108	0.0005
$C_8 (E_h a_0^8)$	80198.08	8750.1	0.3
derived parameters	value	uncertainty	
a_S (bohr)	993	6.9	
a_T (bohr)	512.4	1.4	

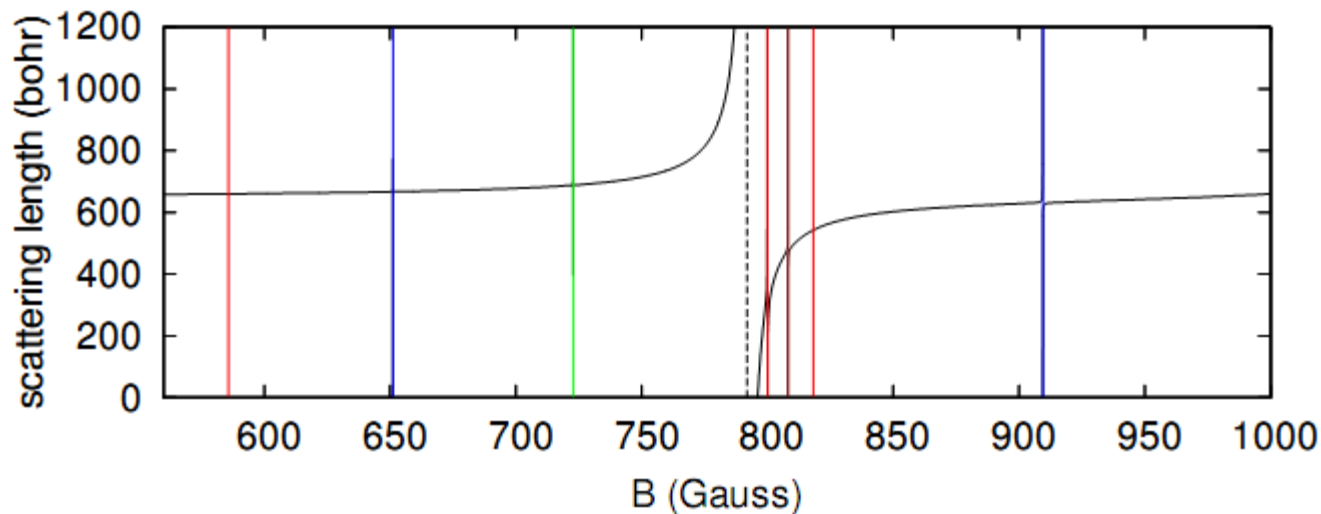
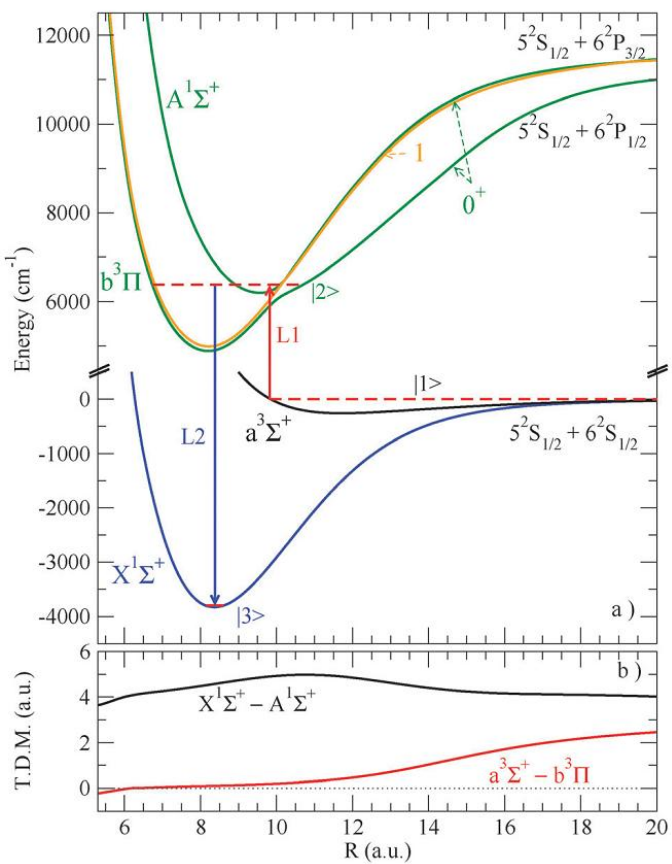


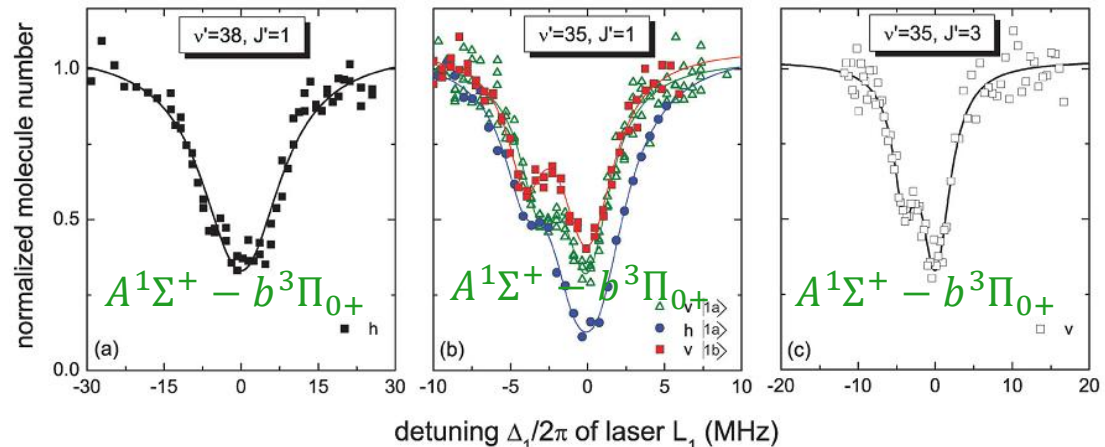
FIG. 10: [Color online.] RbCs scattering length at the $|1, 1\rangle + |3, 3\rangle$ threshold at fields above 560 G, calculated using the final fitted potential at $E = 160$ nK. Resonance positions are marked by vertical lines, with the value of M_F of the corresponding bound state indicated using the same color scheme as in Fig. 7.

Ground state transfer

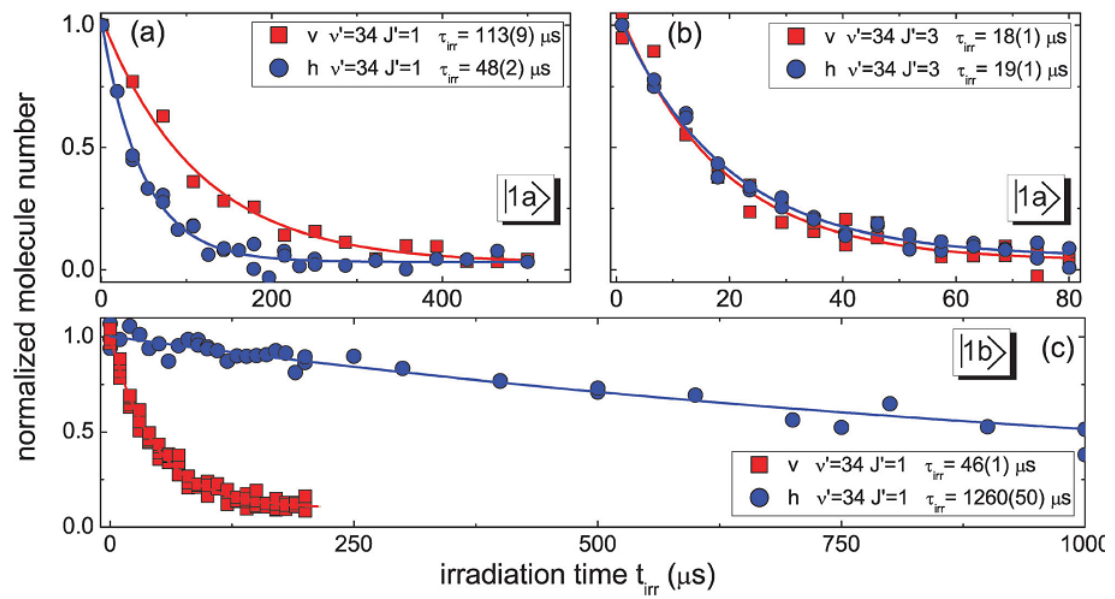


$$\Omega_1 = (\Gamma/\tau_{\text{irr}})^{1/2}$$

$$= \vec{\mu} \cdot \vec{E} / \hbar$$

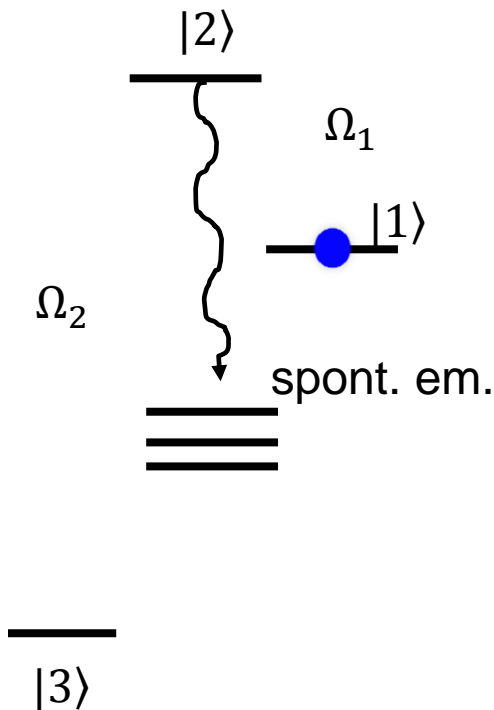


broader than expected from calculations (Vexieau, Bouloufa, Dulieu)



Ground state spectroscopy:

$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 0 & \Omega_2 \\ 0 & \Omega_2 & 0 \end{bmatrix}$$



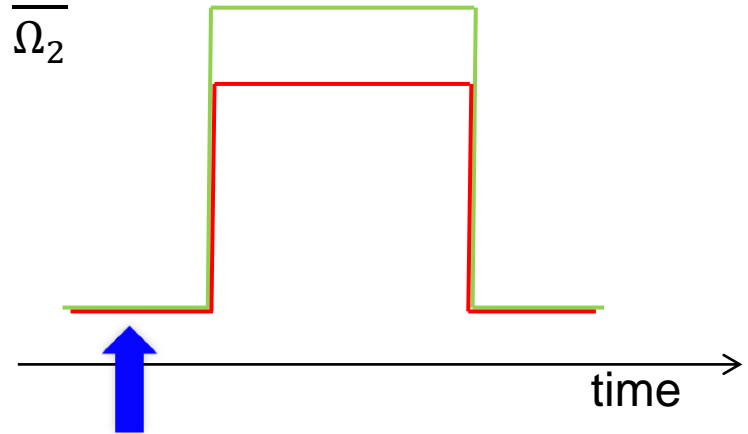
Eigenstates with light on:

$$|a^+\rangle = \sin \Theta \sin \Phi |1\rangle + \cos \Phi |2\rangle + \cos \Theta \sin \Phi |3\rangle$$

$$|a^0\rangle = \cos \Theta |1\rangle - \sin \Theta |3\rangle$$

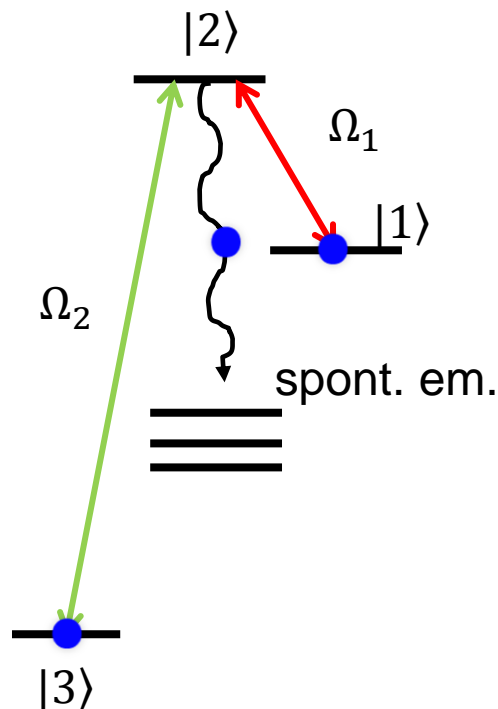
$$|a^-\rangle = \sin \Theta \cos \Phi |1\rangle - \sin \Phi |2\rangle + \cos \Theta \cos \Phi |3\rangle$$

$$\tan \Theta = \frac{\Omega_1}{\Omega_2}$$



Ground state spectroscopy:

$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 0 & \Omega_2 \\ 0 & \Omega_2 & 0 \end{bmatrix}$$



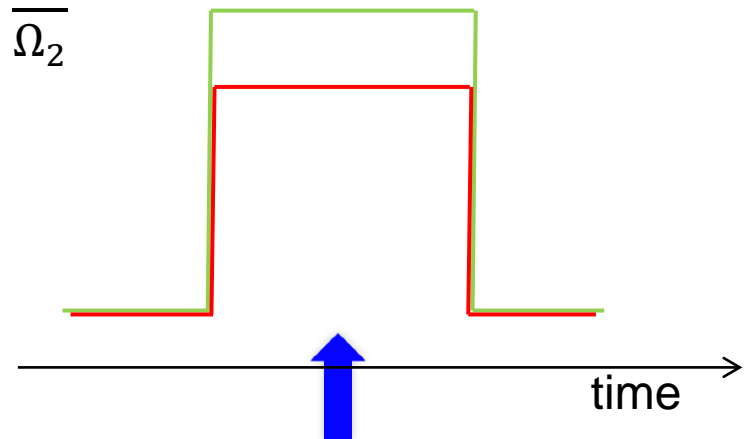
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$$|a^0\rangle = \cos \Theta |1\rangle - \sin \Theta |3\rangle$$

$$|a^-\rangle = \sin \Theta \cos \Phi |1\rangle - \sin \Phi |2\rangle + \cos \Theta \cos \Phi |3\rangle$$

$$\tan \Theta = \frac{\Omega_1}{\Omega_2}$$

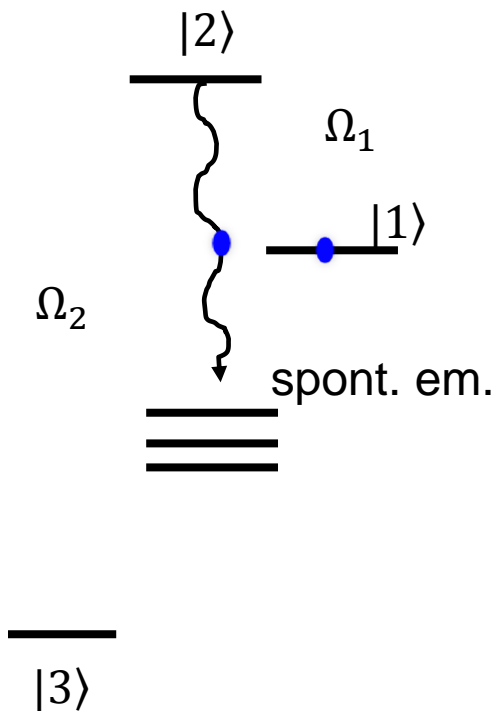


projection $\cos \Theta$ onto dark state

K. Bergmann, H. Theuer, and B.W. Shore: Rev. Mod. Phys. 70, 1003 (1998)

Ground state spectroscopy:

$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 0 & \Omega_2 \\ 0 & \Omega_2 & 0 \end{bmatrix}$$



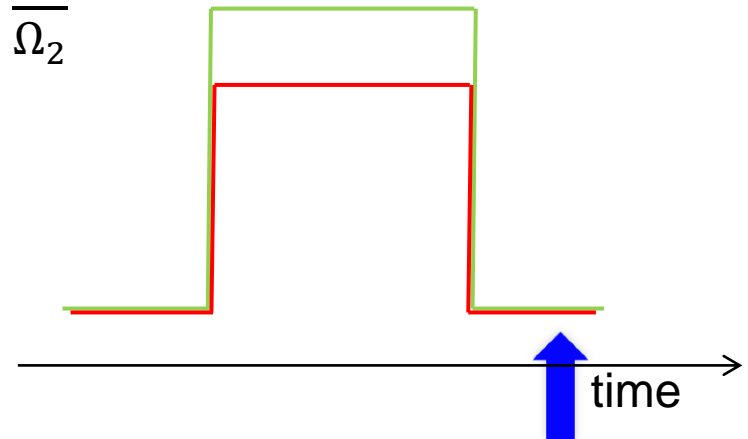
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$$|a^+\rangle = \sin \Theta \sin \Phi |1\rangle + \cos \Phi |2\rangle + \cos \Theta \sin \Phi |3\rangle$$

$$|a^0\rangle = \cos \Theta |1\rangle - \sin \Theta |3\rangle$$

$$|a^-\rangle = \sin \Theta \cos \Phi |1\rangle - \sin \Phi |2\rangle + \cos \Theta \cos \Phi |3\rangle$$

$$\tan \Theta = \frac{\Omega_1}{\Omega_2}$$



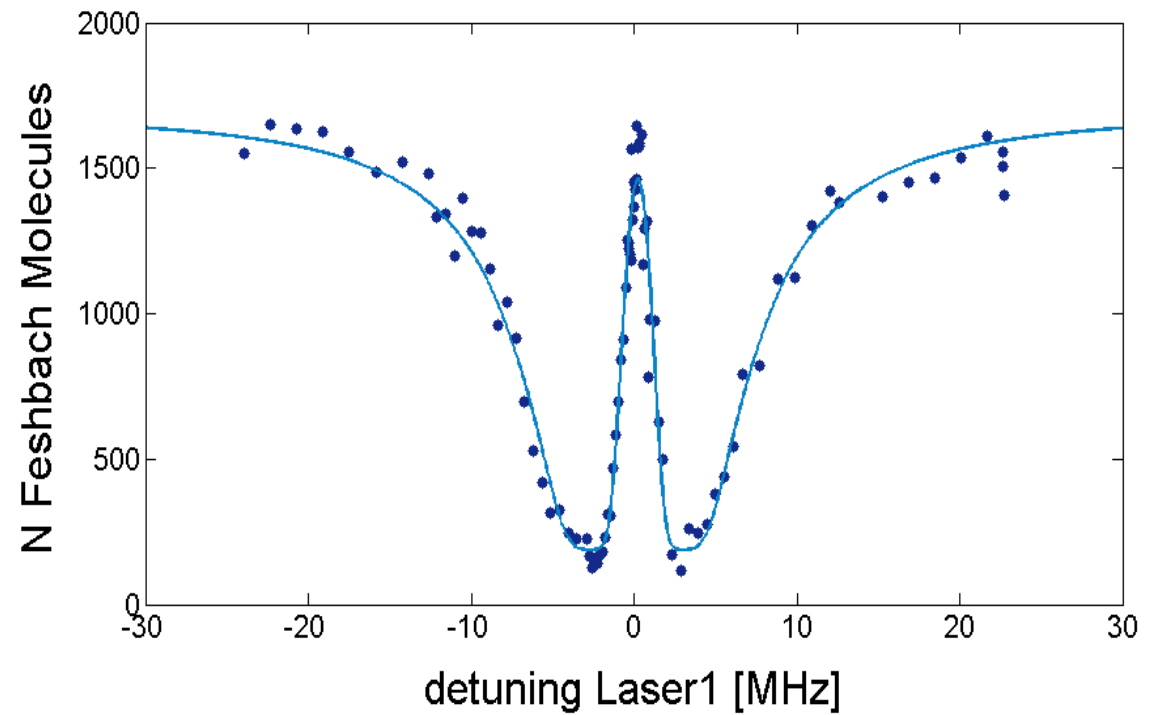
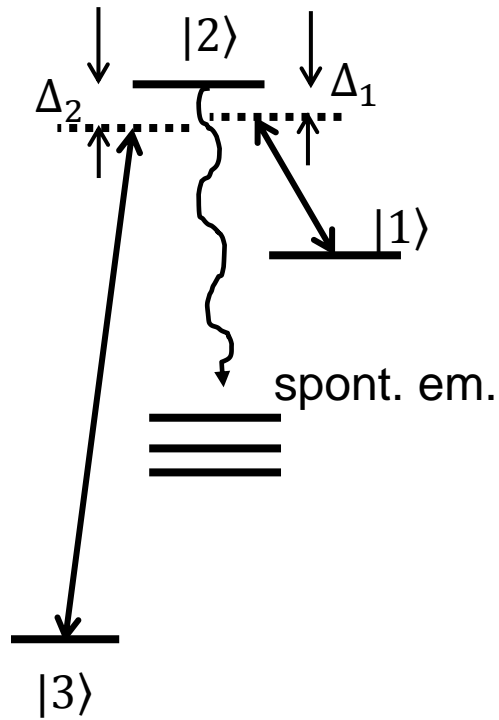
projection $\cos \Theta$ onto $|1\rangle$

$$P_{1 \rightarrow a^0 \rightarrow 1} = P_{1 \rightarrow a^0} P_{a^0 \rightarrow 1} = \cos^4 \Theta = \frac{\Omega_1^4}{(\Omega_1^2 + \Omega_2^2)^2}$$

K. Bergmann, H. Theuer, and B.W. Shore: Rev. Mod. Phys. 70, 1003 (1998)

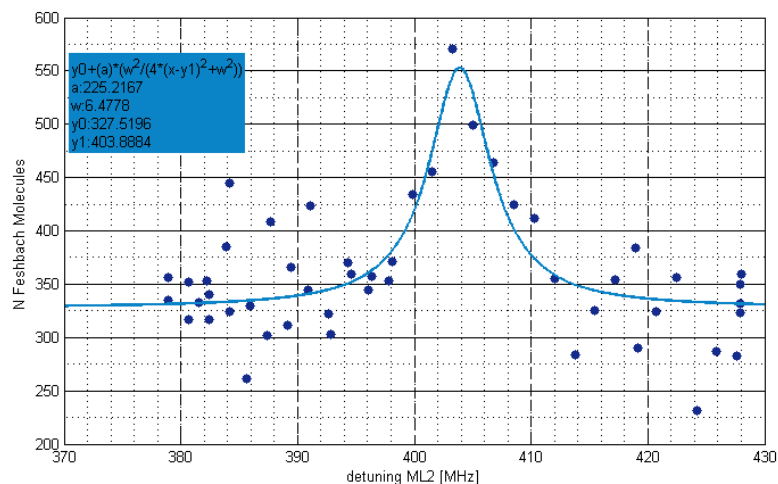
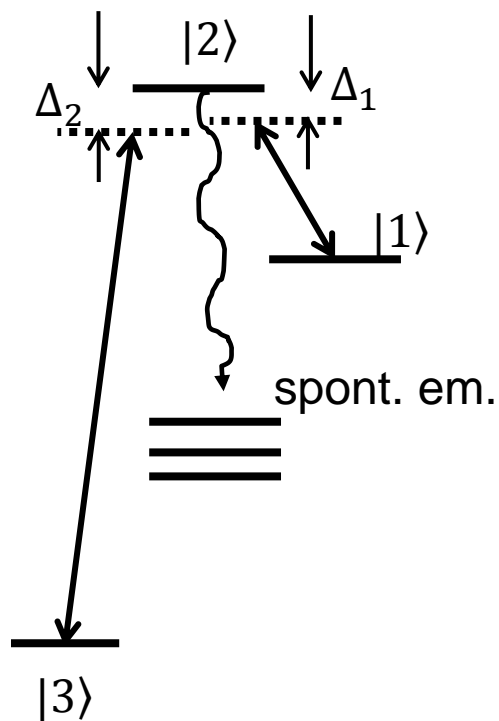
Ground state spectroscopy:

$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 2\Delta_1 & \Omega_2 \\ 0 & \Omega_2 & 2(\Delta_1 - \Delta_2) \end{bmatrix}$$



Ground state spectroscopy:

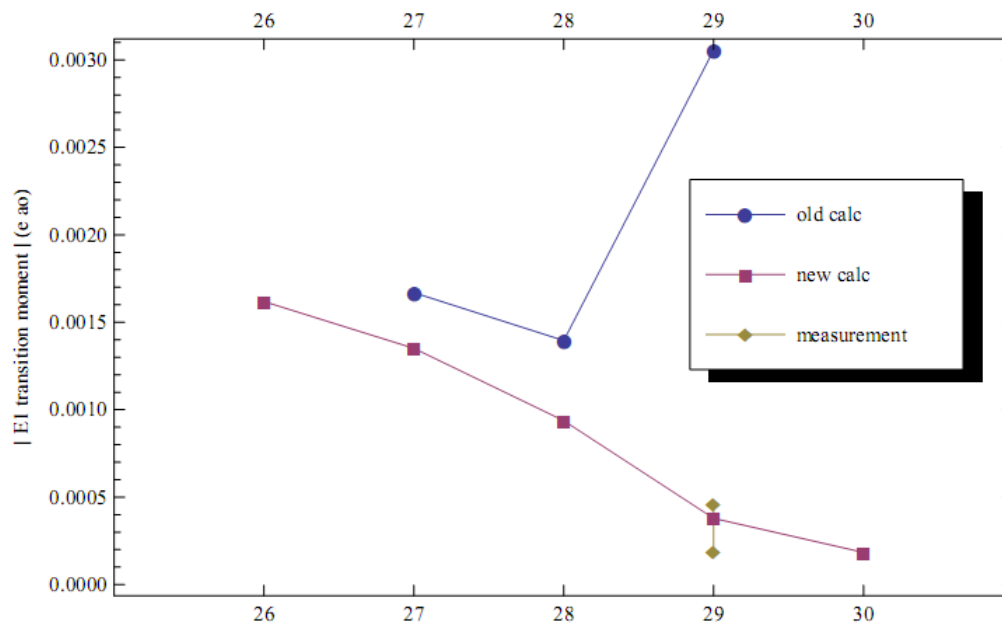
$$H(t) = \frac{\hbar}{2} \begin{bmatrix} 0 & \Omega_1 & 0 \\ \Omega_1 & 2\Delta_1 & \Omega_2 \\ 0 & \Omega_2 & 2(\Delta_1 - \Delta_2) \end{bmatrix}$$



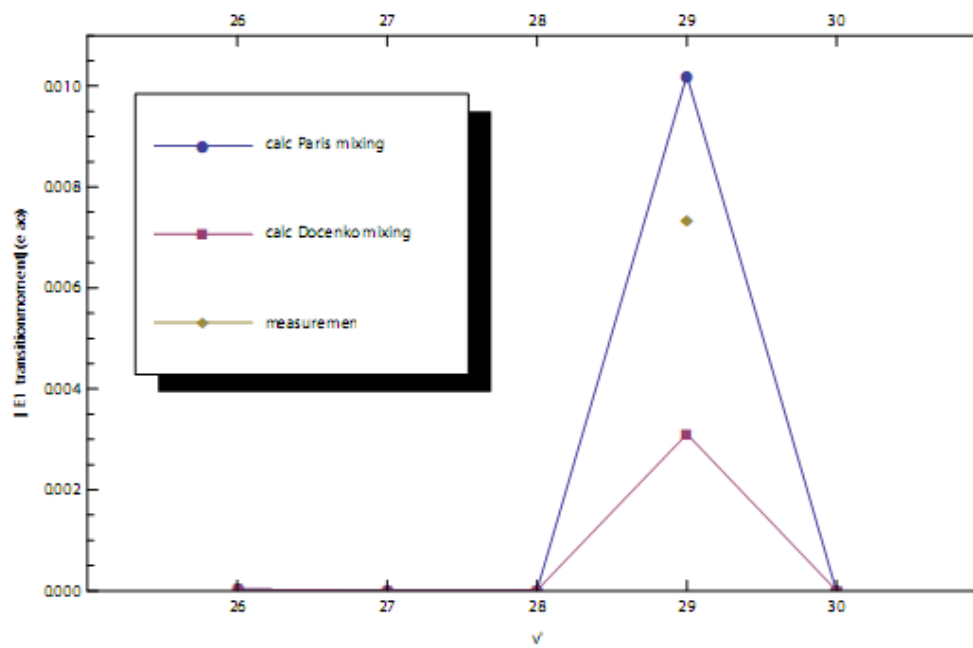
- Have mapped out ground state $v=0$ $N=0,2$.
- Rotational constants agree very well with Fourier transform spectroscopy experiments

iso	J	N	s	n	Trm(cm-1)	<R>(Ang)	Fr(A)	Fr(b0)	Fr(b1)	Fr(b2)	(cm-1)
2	1	52	2	27	10019.87	4.69195	0.23704	0.76296	0	0	6208.30
2	1	53	3	24	10023.59	4.51642	0	0	1	0	6212.01
2	1	54	2	28	10057.53	4.73666	0.29303	0.70697	0	0	6245.95
2	1	55	1	0	10060.47	5.04516	0.89831	0.10169	0	0	6248.89
2	1	56	3	25	10070.94	4.525	0	0	1	0	6259.36
2	1	57	2	29	10094.18	4.76101	0.31243	0.68757	0	0	6282.60
2	1	58	1	1	10108.55	5.0381	0.88108	0.11892	0	0	6296.98
2	1	59	3	26	10118.18	4.53366	0	0	1	0	6306.60
2	1	60	2	30	10130.05	4.78599	0.33485	0.66515	0	0	6318.47
2	1	61	1	2	10157.9	5.03051	0.85921	0.14079	0	0	6346.32
2	1	62	3	27	10165.3	4.54241	0.00001	0.00001	0.99998	0	6353.72
2	1	63	2	31	10165.42	4.80475	0.34841	0.65157	0.00001	0	6353.84
2	1	64	2	32	10200.1	4.84594	0.40076	0.59924	0	0	6388.52
2	1	65	1	3	10208.36	5.00023	0.79685	0.20315	0	0	6396.78
2	1	66	3	28	10212.3	4.55122	0	0	1	0	6400.73
2	1	67	2	33	10234.73	4.85737	0.39963	0.60037	0	0	6423.16
2	1	68	1	4	10259.03	4.99905	0.76369	0.21395	0.00036	0	6447.45
2	1	69	3	29	10259.19	4.56026	0.00029	0.00007	0.99964	0	6447.62
2	1	70	2	34	10269.07	4.87554	0.41152	0.58808	0	0	6457.50
2	1	71	2	35	10303.22	4.89033	0.41885	0.58115	0	0	6491.65
2	1	72	3	30	10305.96	4.56905	0	0	1	0	6494.39
2	1	73	1	5	10309.96	4.99505	0.76648	0.23352	0	0	6498.38
2	1	74	2	36	10337.24	4.91178	0.43766	0.56234	0	0	6525.67
2	1	75	3	31	10352.62	4.57806	0	0	1	0	6541.04
2	1	76	1	6	10360.75	4.97269	0.70972	0.29028	0	0	6549.18
2	1	77	2	37	10371.33	4.94569	0.48286	0.51714	0	0	6559.76
2	1	78	3	32	10399.15	4.58713	0	0	1	0	6587.58
2	1	79	2	38	10404.19	4.90554	0.4046	0.5954	0	0	6592.62
2	1	80	1	7	10412.55	5.02509	0.77545	0.22455	0	0	6600.97

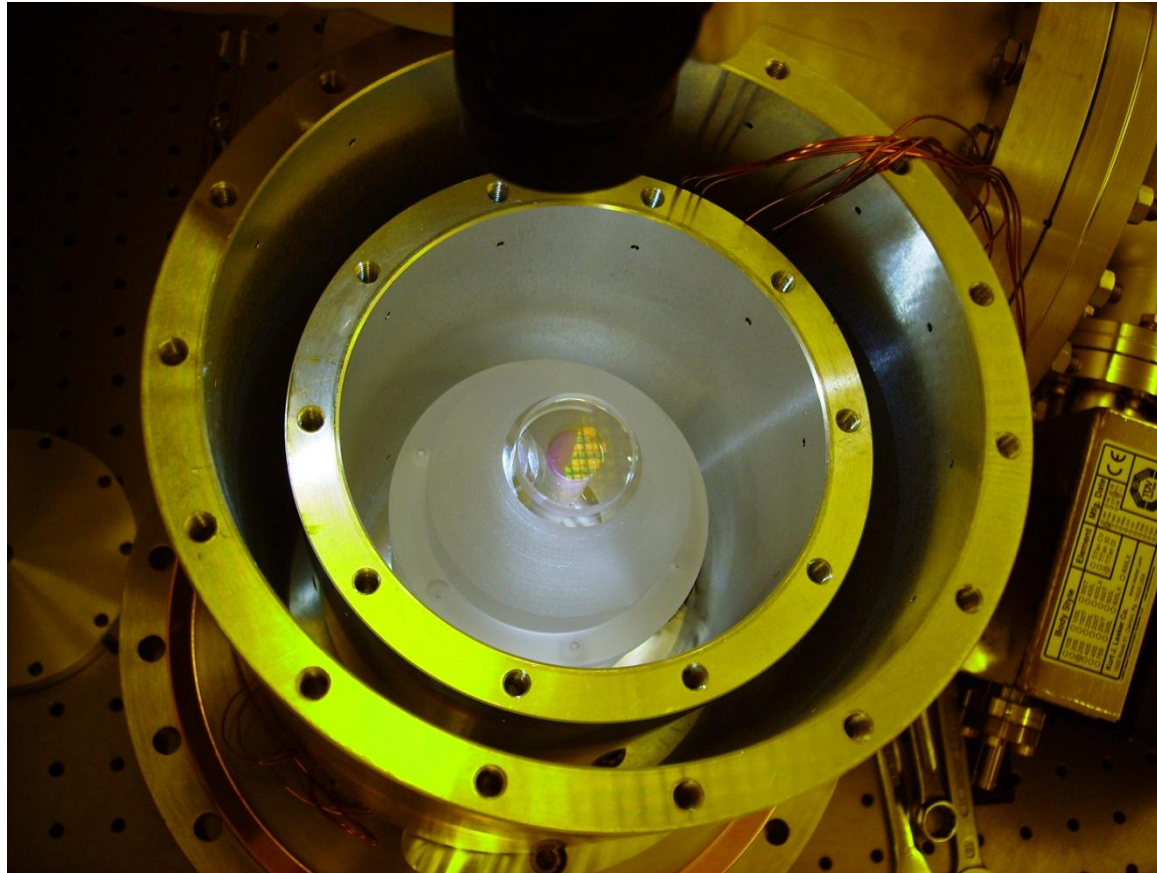
Feshbach \rightarrow b3P11 STIRAP 1 transition



b3P11 \rightarrow X1Sigma⁺ v''=0 transition



Supercavities!



- sub-Hz laser stabilities possible (cavities themselves are the reference)
- limited by acoustics
- finesse ~ 200000
- narrow linewidth diode lasers also being built