## **Estimated Instrument Performance Goals**

Telescope	1.5 m (proj. aperture) aplanatic; shared focal plane; system emissivity 1%								
	Viewing direction offset 85° from spin axis; Field of View 8°								
Instrument	LFI			HFI					
Center Freq. (GHz)	30	44	70	100	143	217	353	545	857
<b>Detector Technology</b>	HEMT LNA arrays			Bolometer arrays					
<b>Detector Temperature</b>	~20 K			0.1 K					
<b>Cooling Requirements</b>	H <sub>2</sub> sorption cooler			H <sub>2</sub> sorption + 4 K J-T stage + Dilution cooler					
Number of Unpol.	0	0	0	0	4	4	4	4	4
Detectors									
Number of Linearly	4	6	12	8	8	8	8	0	0
<b>Polarised Detectors</b>									
Angular Resolution	33	24	14	9.5	7.1	5	5	5	5
(FWHM, arcmin)									
Bandwidth (GHz)	6	8.8	14	33	47	72	116	180	283
Average $\Delta T/T_{I}^{*}$ per	2.0	2.7	4.7	2.5	2.2	4.8	14.7	147	6700
pixel <sup>#</sup>									
Average $\Delta T/T_{U,O}^*$ per	2.8	3.9	6.7	4.0	4.2	9.8	29.8		
pixel <sup>#</sup>									
* Sensitivity (1σ) to intensity (Stokes I) fluctuations observed on the sky, in thermodynamic temperature (x10 <sup>-6</sup> ) units, relative to the									

average temperature of the CMB (2.73 K), achievable after two sky surveys (14 months).

<sup>#</sup> A pixel is a square whose side is the FWHM extent of the beam.

\* Sensitivity (1σ) to polarised intensity (Stokes U and Q) fluctuations observed on the sky, in thermodynamic temperature (x10<sup>-6</sup>) units, relative to the average temperature of the CMB (2.73 K), achievable after two sky surveys (14 months).



