

# Electroweak lights from DM annihilations

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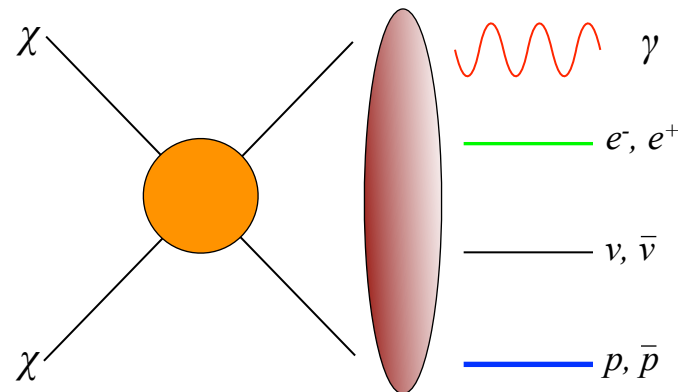
ÉCOLE POLYTECHNIQUE  
FÉDÉRALE DE LAUSANNE

Based on:

*Ciafaloni, Cirelli, Comelli, DS, Riotto, Urbano*

arXiv:1104.2996 (to appear on JCAP)

# Introduction



Radiation of EW gauge bosons is a **SM effect** and can have a big impact on final fluxes in 3 situations:

1. when looking at low-energy tails of the spectra, mostly populated by decay products of extra gauge bosons;
2. when some species are absent without EW corrections (e.g.  $\bar{p}$  from  $\chi\chi \rightarrow \ell^+\ell^-$ );
3. (*this talk*) when  $\sigma(2 \rightarrow 2)$  is suppressed, so  $\sigma(2 \rightarrow 3)$  can even dominate.

# Annihilations of Majorana fermions

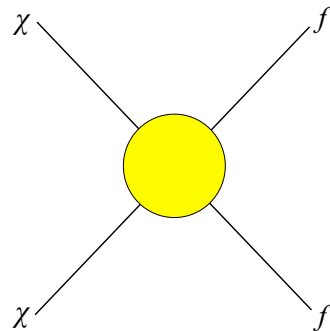
$$v\sigma_{\text{ann}} = \textcolor{red}{a} + \textcolor{green}{b}v^2 + \mathcal{O}(v^4)$$

$\uparrow$        $\uparrow$

$\textcolor{red}{s\text{-wave}}$     $\textcolor{green}{p\text{-wave}}$

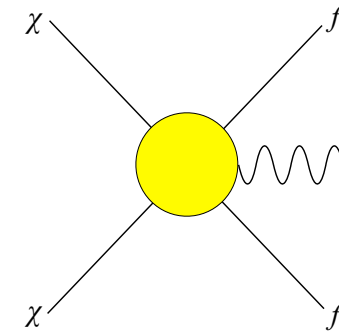
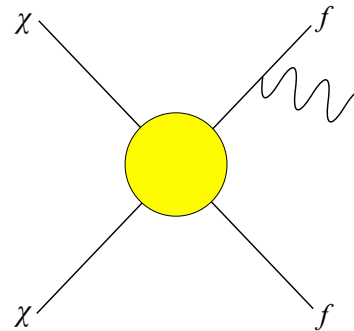
(today  $v \sim 10^{-3}$ )

For a Majorana fermion and SM singlet (e.g. Bino in SUSY)



only  $p\text{-wave}$   
 $(m_f \ll M_\chi)$

radiation  
 $\Rightarrow$



$\textcolor{red}{\text{there is an } s\text{-wave!}}$

# The Model

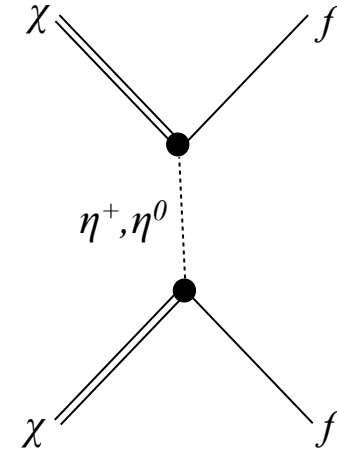
The DM couples to the SM via a heavy scalar doublet:  $S = \begin{pmatrix} \eta^+ \\ \eta^0 \end{pmatrix}$

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_\chi + \mathcal{L}_S + \mathcal{L}_{\text{int}}$$

$$\mathcal{L}_\chi = \frac{1}{2} \bar{\chi} (i \not{\partial} - M_\chi) \chi,$$

$$\mathcal{L}_S = (D_\mu S)^\dagger (D^\mu S) - M_S^2 S^\dagger S,$$

$$\begin{aligned} \mathcal{L}_{\text{int}} &= y_L \bar{\chi} (L i \sigma_2 S) + \text{h.c.} \\ &= y_L (\bar{\chi} P_L f_2 \eta^+ - \bar{\chi} P_L f_1 \eta^0) + \text{h.c.} \end{aligned}$$

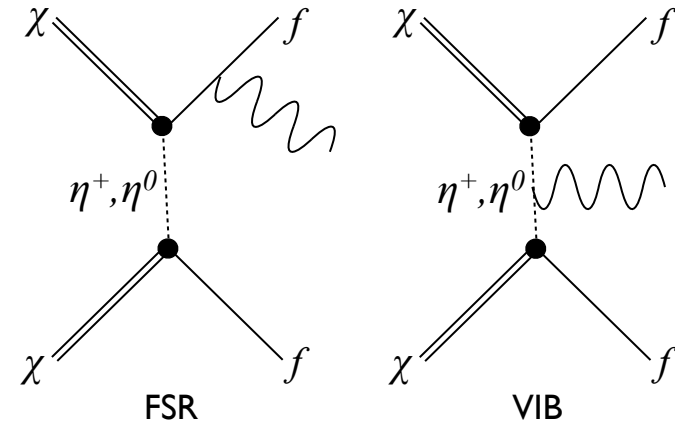


Mass parameters:  $M_\chi, M_S$

$$\leadsto M_\chi, \textcolor{red}{r} \equiv (M_S/M_\chi)^2 \geq 1$$

$$v\sigma \sim \frac{1}{M_\chi^2} \frac{v^2}{r^2}$$

Now add radiation of EW gauge bosons  $\rightsquigarrow$



Schematically, the amplitude is

$$\mathcal{M} \sim \frac{1}{M_\chi} \mathcal{O}(v) \left[ \mathcal{O}\left(\frac{1}{r}\right) \Big|_{\text{FSR}} + \mathcal{O}\left(\frac{1}{r^2}\right) \Big|_{\text{FSR}} \right] + \frac{1}{M_\chi} \left[ \mathcal{O}\left(\frac{1}{r^2}\right) \Big|_{\text{VIB}} + \mathcal{O}\left(\frac{1}{r^2}\right) \Big|_{\text{FSR}} \right]$$

and the cross section

$$v\sigma(\chi\chi \rightarrow f\bar{f}Z) \sim \frac{\alpha_W}{M_\chi^2} \left[ \mathcal{O}\left(\frac{v^2}{r^2}\right) + \mathcal{O}\left(\frac{v^2}{r^3}\right) + \mathcal{O}\left(\frac{1}{r^4}\right) \right]$$

Important lesson:

- limiting the expansion to  $\mathcal{O}(1/r)$  in the amplitude keeps the annihilation in  $p$ -wave.
- at  $\mathcal{O}(1/r^2)$ , with VIB diagrams, the  $s$ -wave is opened.

*When does the 3-body process dominate over the 2-body one?*

Estimate:

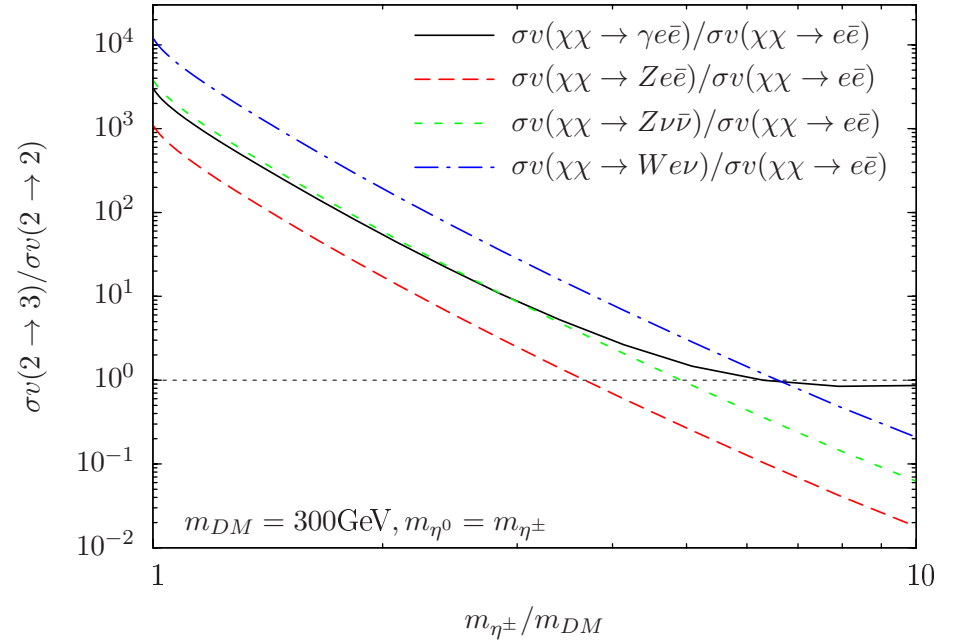
$$v\sigma(2 \rightarrow 2) \sim \frac{1}{M_\chi^2} \frac{v^2}{r^2}$$

$$v\sigma(2 \rightarrow 3) \sim \frac{1}{M_\chi^2} \frac{\alpha_W}{4\pi} \frac{1}{r^4}$$

$\sigma(2 \rightarrow 3) \gtrsim \sigma(2 \rightarrow 2)$  when

$$r \lesssim \sqrt{\frac{\alpha_W}{4\pi}} \frac{1}{v} \sim \mathcal{O}(10)$$

$$(r \equiv M_S^2/M_\chi^2)$$



[Figure from: Garny, Ibarra, Vogl – 1105.5367]

# Effective Field Theory

Integrate out the heavy scalar  $S$ :

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\chi} + \frac{1}{r} \frac{\mathcal{O}_6}{M_{\chi}^2} + \frac{1}{r^2} \frac{\mathcal{O}_8}{M_{\chi}^4} + \dots$$

The lowest-dimensional operator gives a  $p$ -wave annihilation:

$$\mathcal{O}_6 = \frac{1}{2} |y_L|^2 [\bar{\chi} \gamma_{\mu} \gamma_5 \chi] [\bar{L} \gamma^{\mu} P_L L] \quad \Longrightarrow \quad v \sigma(\chi\chi \rightarrow f \bar{f} Z)|_{\mathcal{O}_6} \propto \frac{|y_L|^4}{M_{\chi}^2} \frac{v^2}{r^2}$$

- The  $s$ -wave appears due to  $\mathcal{O}_8$ .  
 $\mathcal{O}_8$  can be more important than  $\mathcal{O}_6$  despite larger dimensionality.
- **Warning:** in this case, naive dimensional analysis fails to assess the relative importance of operators in the expansion.

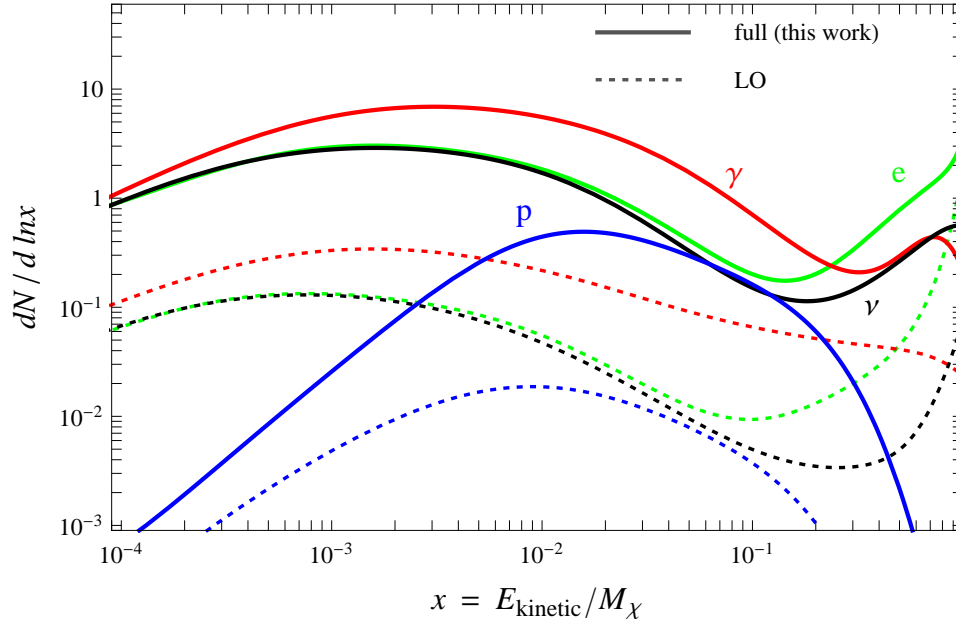
# More Quantitative Analysis

$$\chi\chi \rightarrow e^+e^-, \nu\bar{\nu}, e^+e^-\gamma, e^+e^-Z, \nu\bar{\nu}Z, e^\pm\nu W^\mp$$

- our MC: generates primary annihilation events ( $2 \rightarrow 3$ ) according to the  $|\mathcal{M}|^2$  distribution
- PYTHIA 8.1: for showering + hadronization + decay to final stable SM particles.  
(Technical remark: PYTHIA 6 does not include  $\gamma \rightarrow f\bar{f}$  branchings in the showering).
- extract energy spectra at interaction point for each species
- propagation in the galactic halo



# Energy spectra at the interaction point



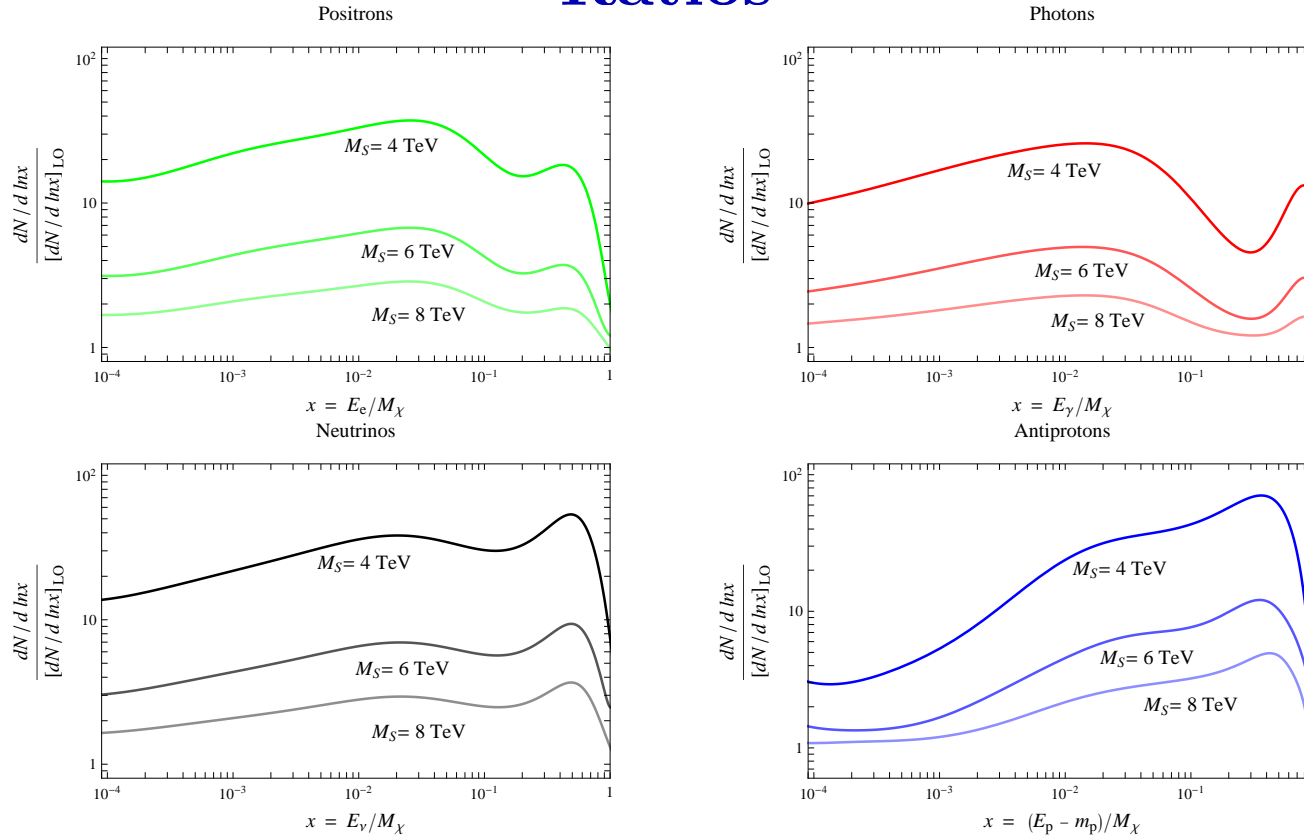
$$M_S = 4 \text{ TeV}$$

$$M_\chi = 1 \text{ TeV}$$

$$\frac{dN_f}{d\ln x} = \frac{1}{\sigma(2 \rightarrow 2)} \frac{d\sigma(\chi\chi \rightarrow f + X)}{d\ln x}$$

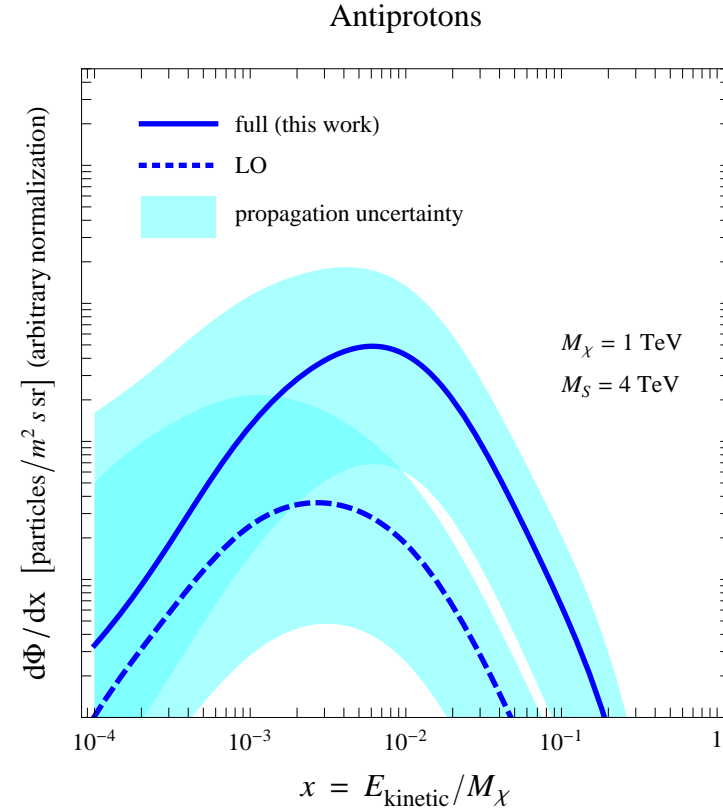
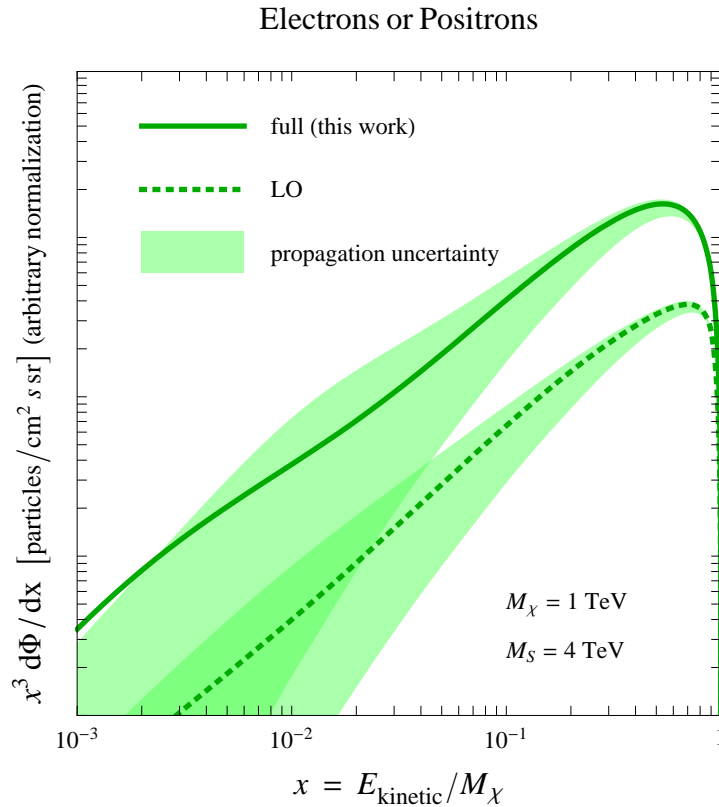
- \* “LO” means adding EW radiation at the lowest order, keeping only the  $\mathcal{O}(1/r)$  in the amplitude ( $p$ -wave).
- \* Bump of primary hard photons due to  $s$ -wave annihilation  $\chi\chi \rightarrow e^+e^-\gamma$ .
- \* Large low-energy tails due to showering and hadronization of  $W, Z$ .

# Ratios



- \*  $dN/dE/[dN/dE]_{\text{LO}} \sim \mathcal{O}(10 - 100)$
- \* Of course, much larger enhancement wrt not including EW corrections.

# Propagated fluxes



- Neutral particles ( $\gamma, \nu$ ) just go straight.
- Propagation does not spoil the effect.

# Conclusions

- Majorana DM annihilates through  $s$ -wave once EW radiation is included.
- Care when using EFT: the naive dim analysis can be misleading. This effect is missed by  $\mathcal{O}_6$ .
- The resulting spectra get substantially enhanced by factors  $\mathcal{O}(10 - 100)$  (with respect to  $p$ -wave only).  
Even more drastic effect with respect to the case without EW corrections.
- EW corrections are a SM effect (no exotics!).  
Reliable calculations of fluxes for DM ID should take them into account.